

Noise in radio communications

Matjaž Vidmar

LSO, FE, Ljubljana, 5.2.2017

List of figures: Noise in radio communications

- 1 - The dispute between famous scientists
- 2 - Noise spectral density
- 3 - Black-body thermal radiation
- 4 - Received thermal-noise power
- 5 - Thermal equilibrium
- 6 - Natural noise sources
- 7 - Natural sky noise
- 8 - Sun-noise example
- 9 - Receiver signal-to-noise ratio
- 10 - Chain noise temperature
- 11 - Amplifier noise figure
- 12 - Relationship $F \leftrightarrow T$
- 13 - Attenuator noise
- 14 - G/T figure of merit
- 15 - Noise of active components
- 16 - Noise parameters
- 17 - Example: the sensitivity of a GSM phone
- 18 - Change of S/N versus F
- 19 - Receiver-sensitivity measurement
- 20 - Hot/cold method
- 21 - Noise-figure meter
- 22 - Bit-Error Rate BER calculation
- 23 - BER \leftrightarrow S/N table for BPSK
- 24 - BER for different modulations
- 25 - Forward Error Correction (FEC)
- 26 - Oscillator phase noise
- 27 - Leeson's equation
- 28 - $1/f$ noise
- 29 - Resonator quality Q
- 30 - Phase-Locked Loop (PLL)
- 31 - Effects of phase noise
- 32 - Phase noise without approximations
- 33 - Width of Lorentzian spectral line
- 34 - Noise as test signal
- 35 - Cryptographic-key source
- 36 - Noise cryptography
- 37 - LFSR pseudo-random sequences
- 38 - Use of pseudo-random sequences

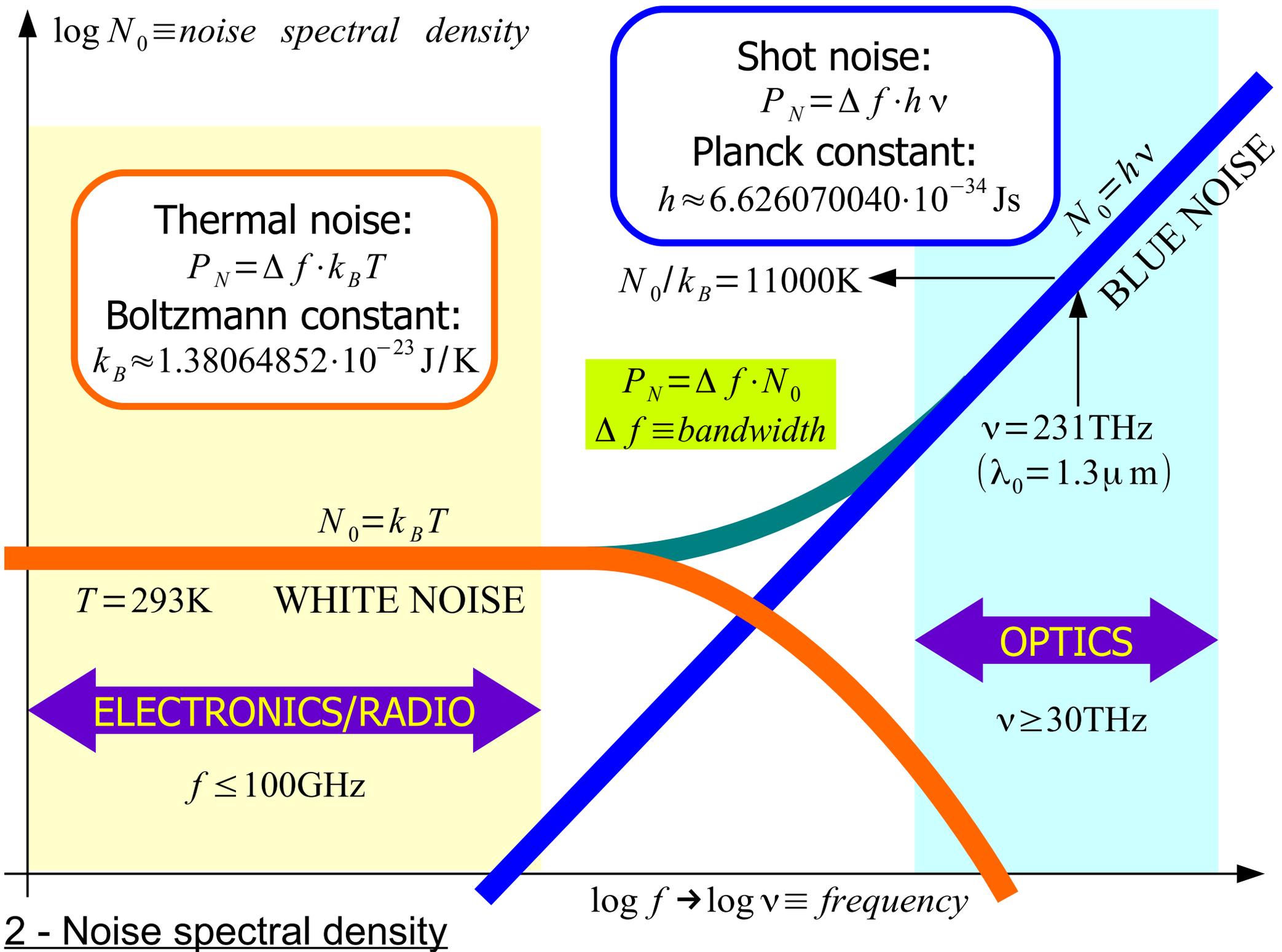
Fifth Solvay International Conference on Electrons and Photons (October 1927). The leading figures Albert Einstein and Niels Bohr disagreed:

Albert Einstein: „God does not play dice!“

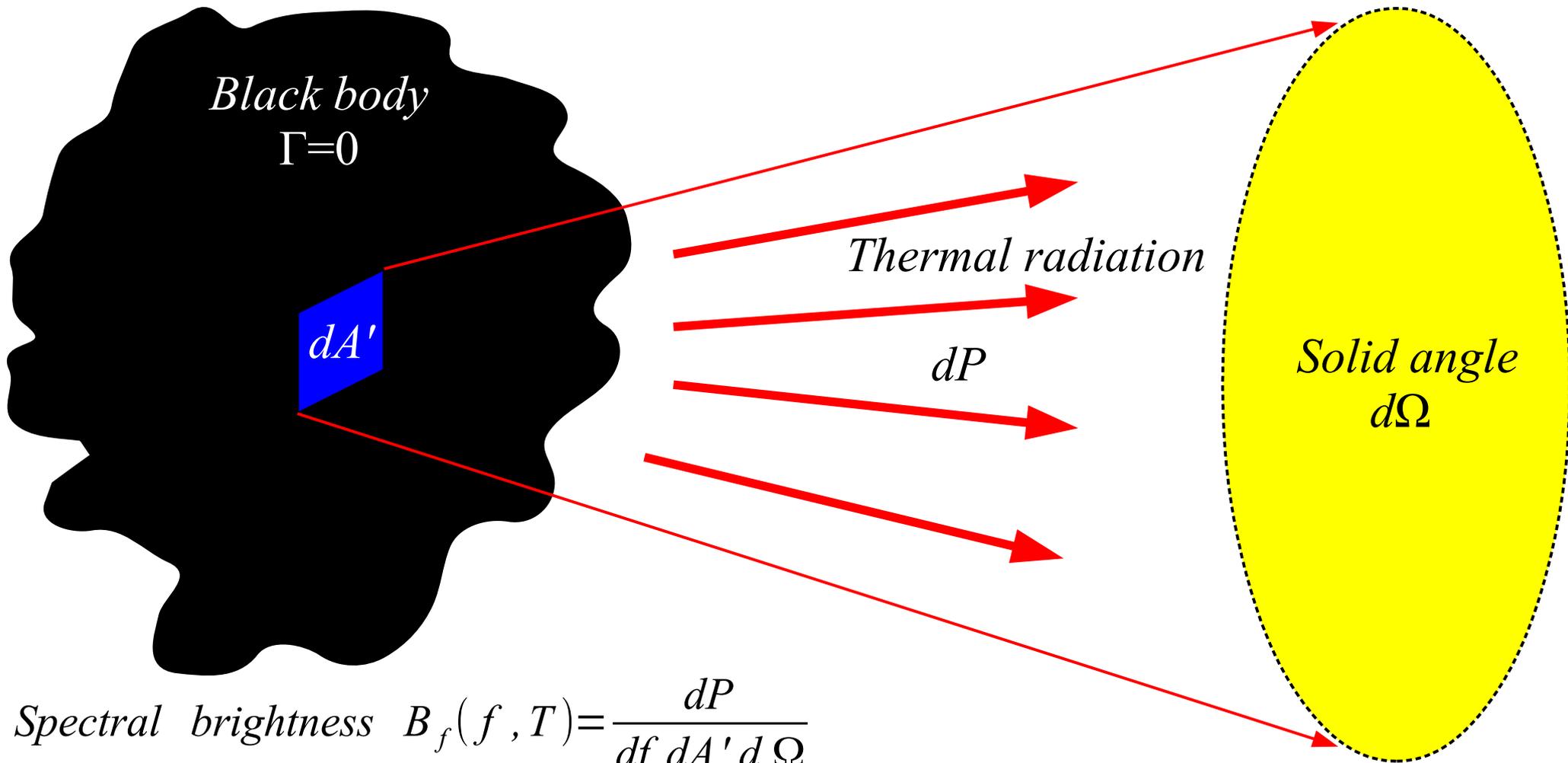
Niels Bohr: „Einstein, stop telling God what to do!“

In telecommunications random signals are called noise. Noise impairs the performance of any communication link.

Noise is a macroscopic description of quantum effects!



2 - Noise spectral density



Spectral brightness $B_f(f, T) = \frac{dP}{df dA' d\Omega}$

Planck law $B_f(f, T) = \frac{2 h f^3}{c_0^2} \cdot \frac{1}{e^{\frac{hf}{k_B T}} - 1}$

Free space ϵ_0, μ_0
 $c_0 = 299792458 \text{ m/s} \approx 3 \cdot 10^8 \text{ m/s}$

Radio $hf \ll k_B T \rightarrow$ *Rayleigh-Jeans approximation* $B_f(f, T) \approx \frac{2 k_B T f^2}{c_0^2} = \frac{2 k_B T}{\lambda^2}$

3 – Black-body thermal radiation

Free space ϵ_0, μ_0

Black body
 $\Gamma=0$

Single polarization

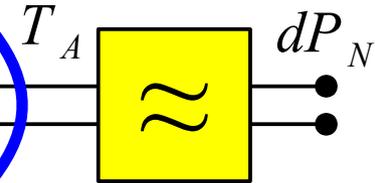
Lossless antenna
 $\eta=1$

$$B_f = \frac{2k_B T}{\lambda^2}$$

dA'

$$dP_N = \frac{1}{2} \cdot B_f \cdot \Delta f \cdot dA' \cdot \Delta \Omega$$

$A_{eff}(\Theta, \Phi)$



Bandpass filter Δf

r

$$\Delta \Omega = \frac{A_{eff}(\Theta, \Phi)}{r^2} = \frac{\lambda^2 D(\Theta, \Phi)}{4\pi r^2} = \frac{\lambda^2 |F(\Theta, \Phi)|^2}{r^2 \iint_{4\pi} |F(\Theta^*, \Phi^*)|^2 d\Omega^*}$$

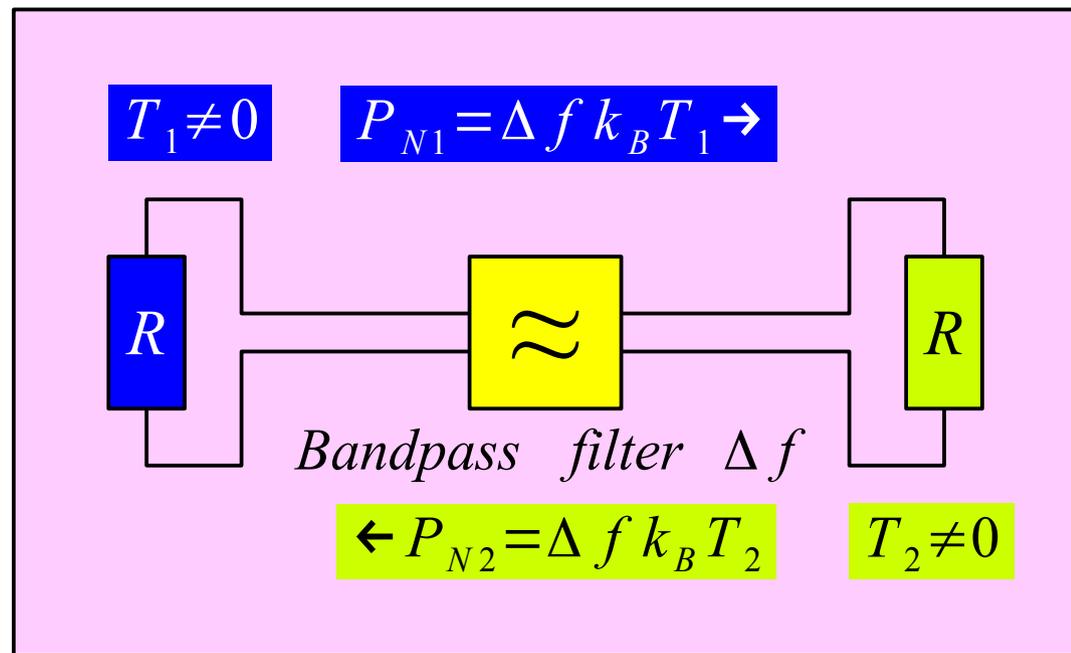
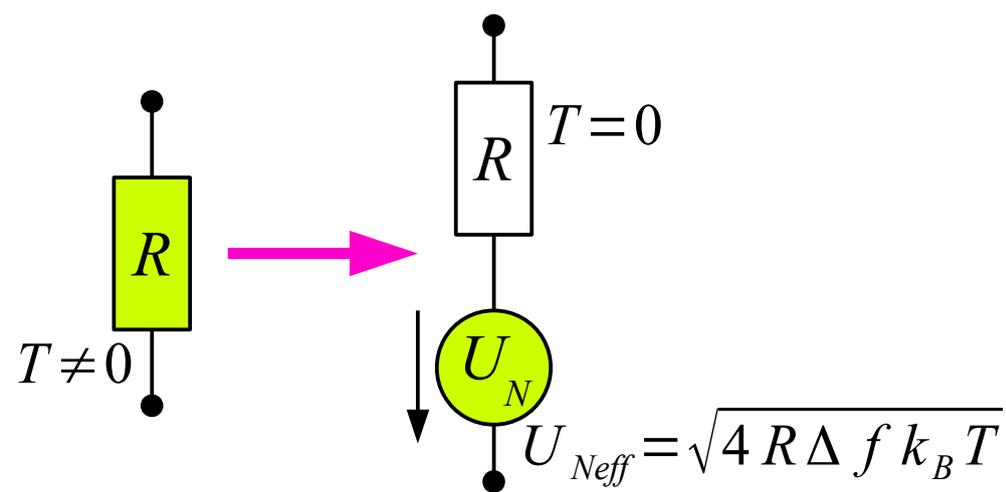
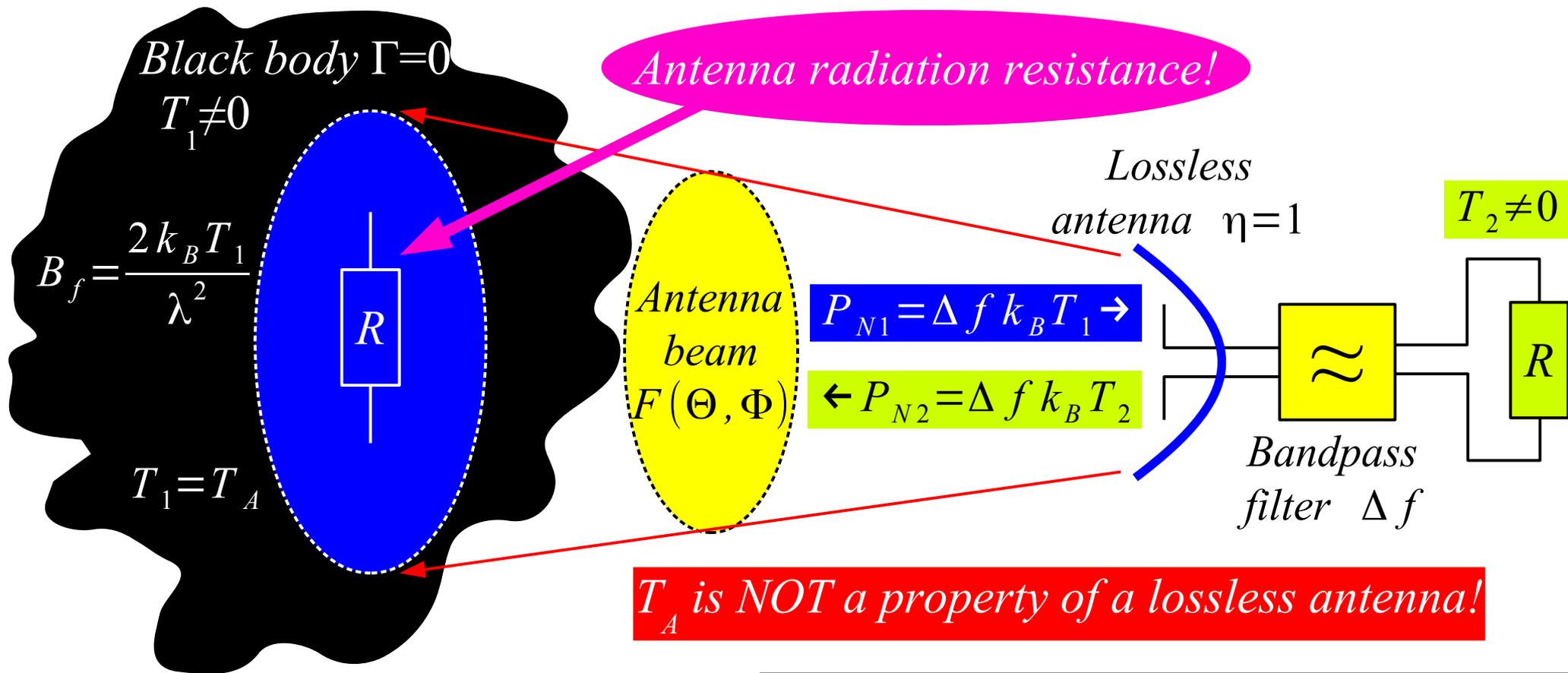
$$dA' = r^2 d\Omega$$

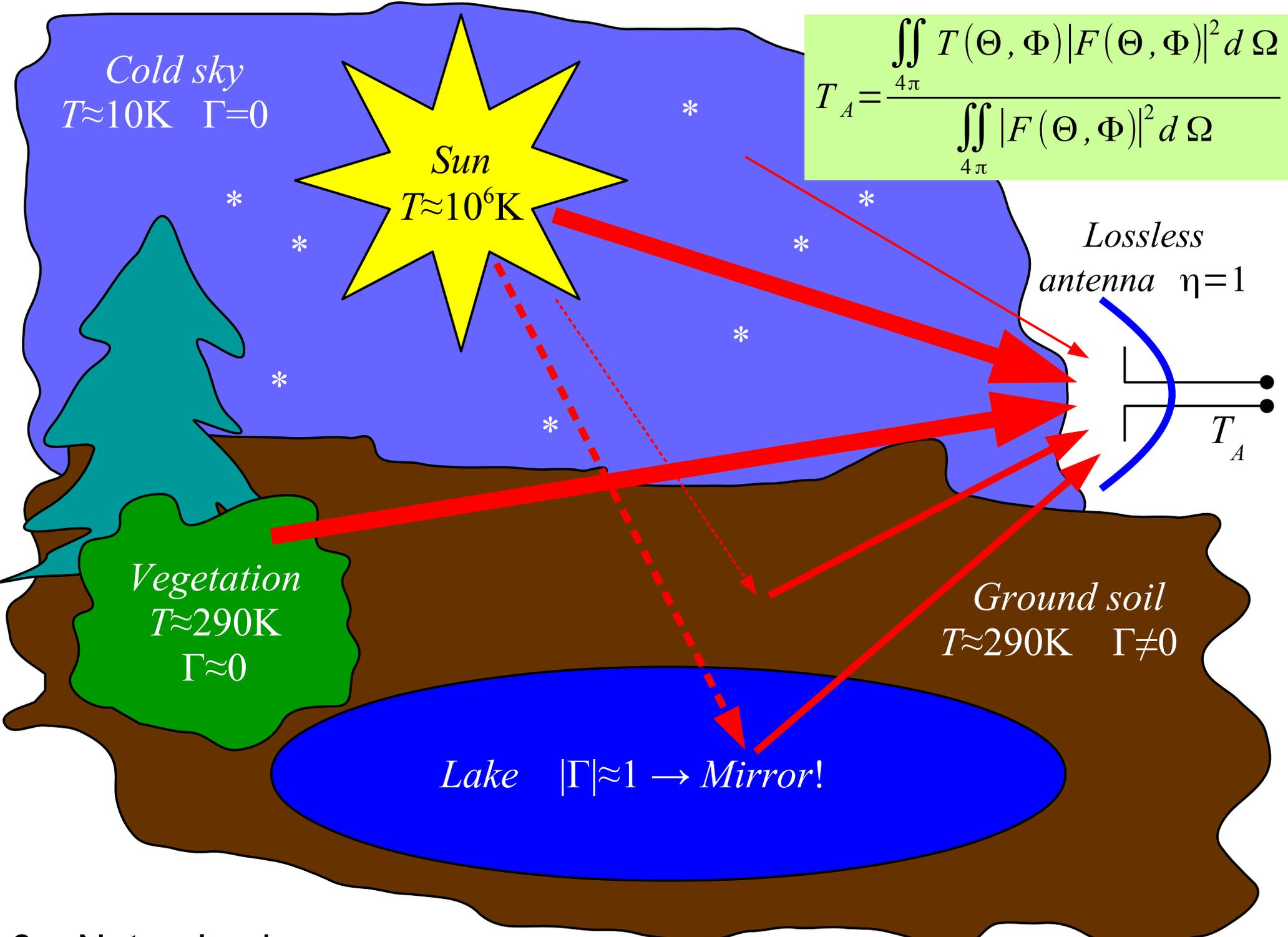
$$P_N = \iint_{A'} \frac{1}{2} \cdot B_f \cdot \Delta f \cdot dA' \cdot \Delta \Omega = \iint_{4\pi} \frac{1}{2} \cdot \frac{2k_B T(\Theta, \Phi)}{\lambda^2} \cdot \Delta f \cdot r^2 d\Omega \cdot \frac{\lambda^2 |F(\Theta, \Phi)|^2}{r^2 \iint_{4\pi} |F(\Theta^*, \Phi^*)|^2 d\Omega^*}$$

$$P_N = \Delta f k_B \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega} = \Delta f k_B T_A$$

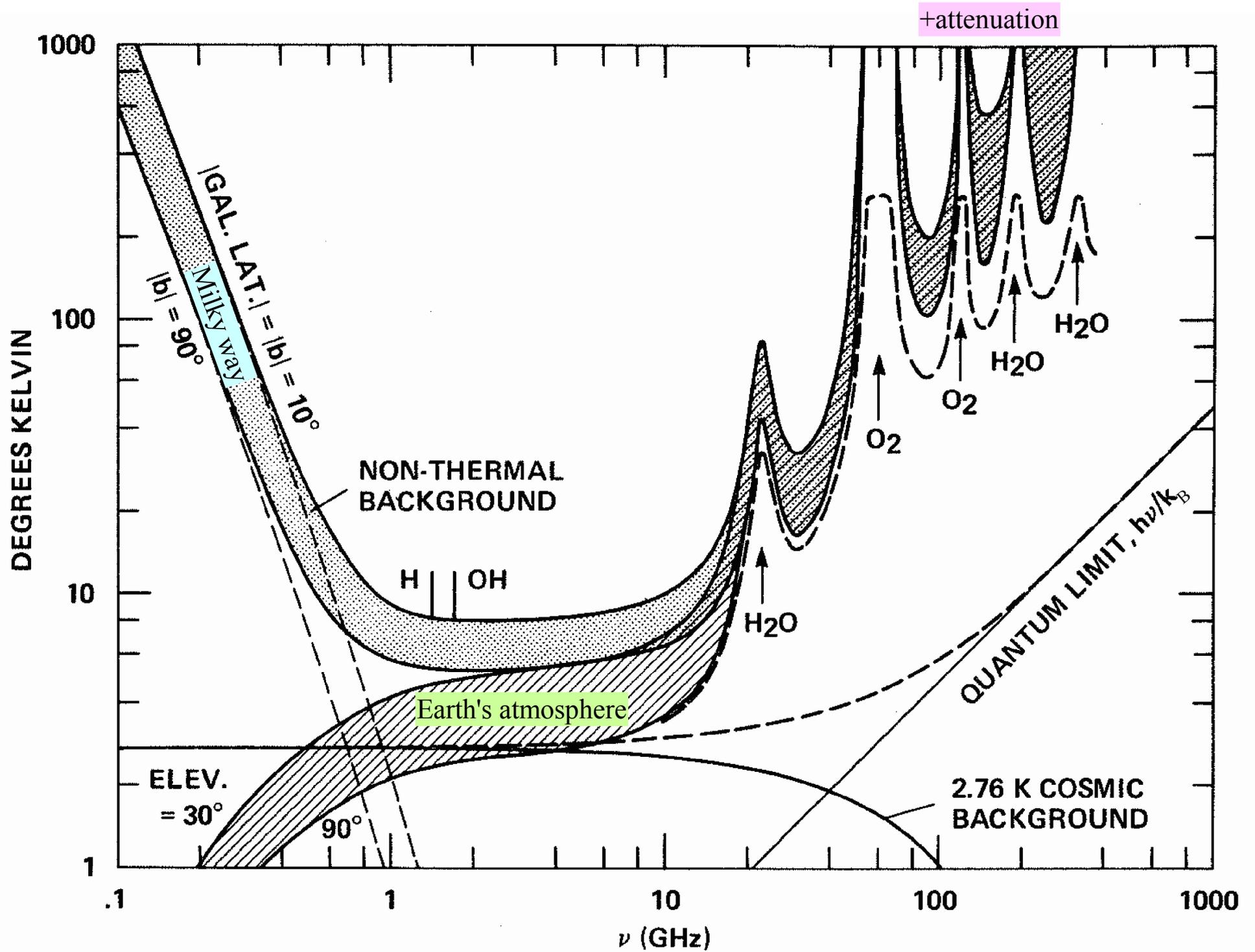
$$T_A = \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

4 – Received thermal-noise power

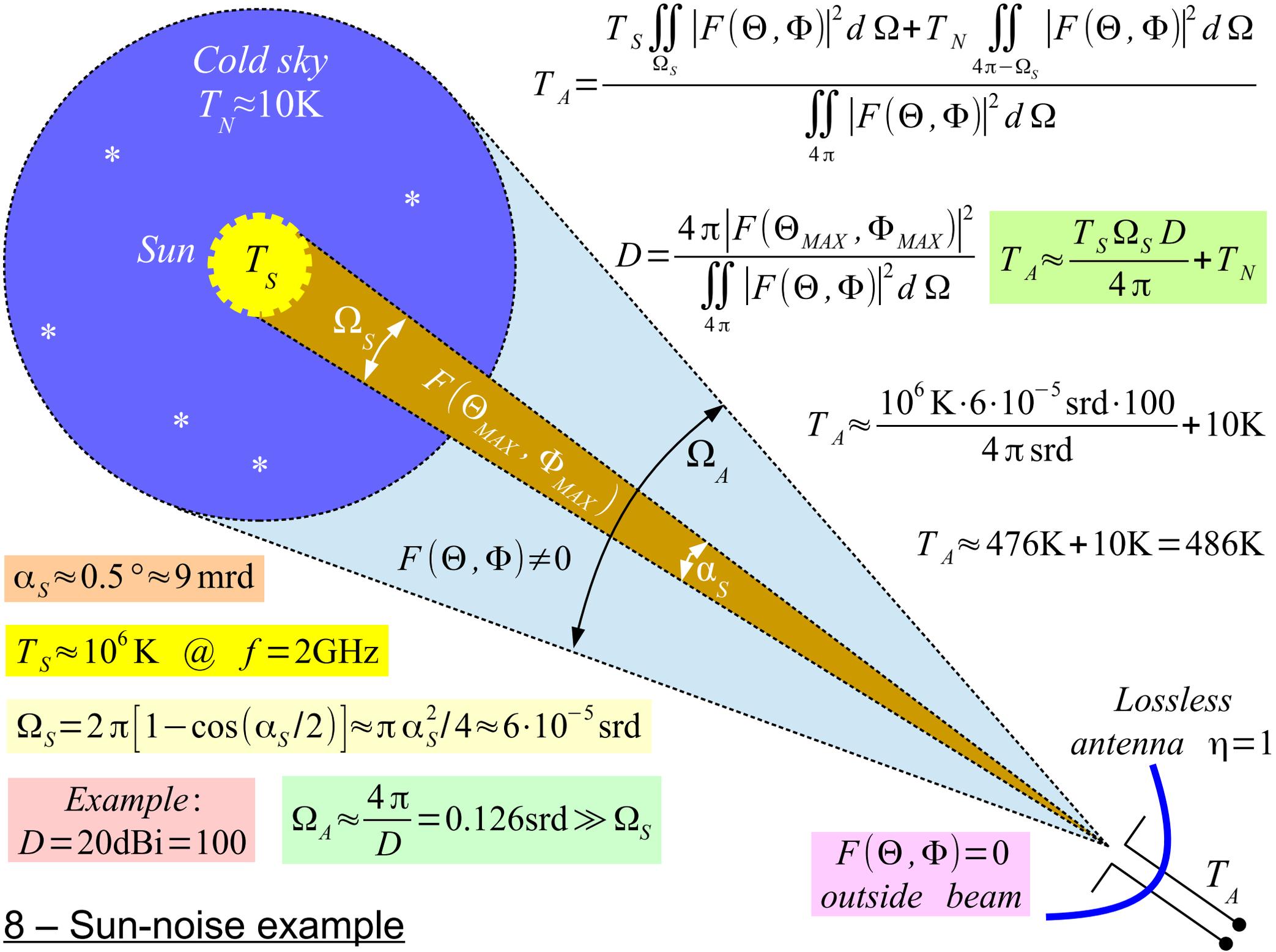




6 – Natural noise sources



7 – Natural sky noise



$$T_A = \frac{T_S \iint_{\Omega_S} |F(\Theta, \Phi)|^2 d\Omega + T_N \iint_{4\pi - \Omega_S} |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

$$D = \frac{4\pi |F(\Theta_{MAX}, \Phi_{MAX})|^2}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega} \quad T_A \approx \frac{T_S \Omega_S D}{4\pi} + T_N$$

$$T_A \approx \frac{10^6 \text{K} \cdot 6 \cdot 10^{-5} \text{srd} \cdot 100}{4\pi \text{srd}} + 10\text{K}$$

$$T_A \approx 476\text{K} + 10\text{K} = 486\text{K}$$

$$\alpha_S \approx 0.5^\circ \approx 9 \text{ mrd}$$

$$T_S \approx 10^6 \text{K} @ f = 2\text{GHz}$$

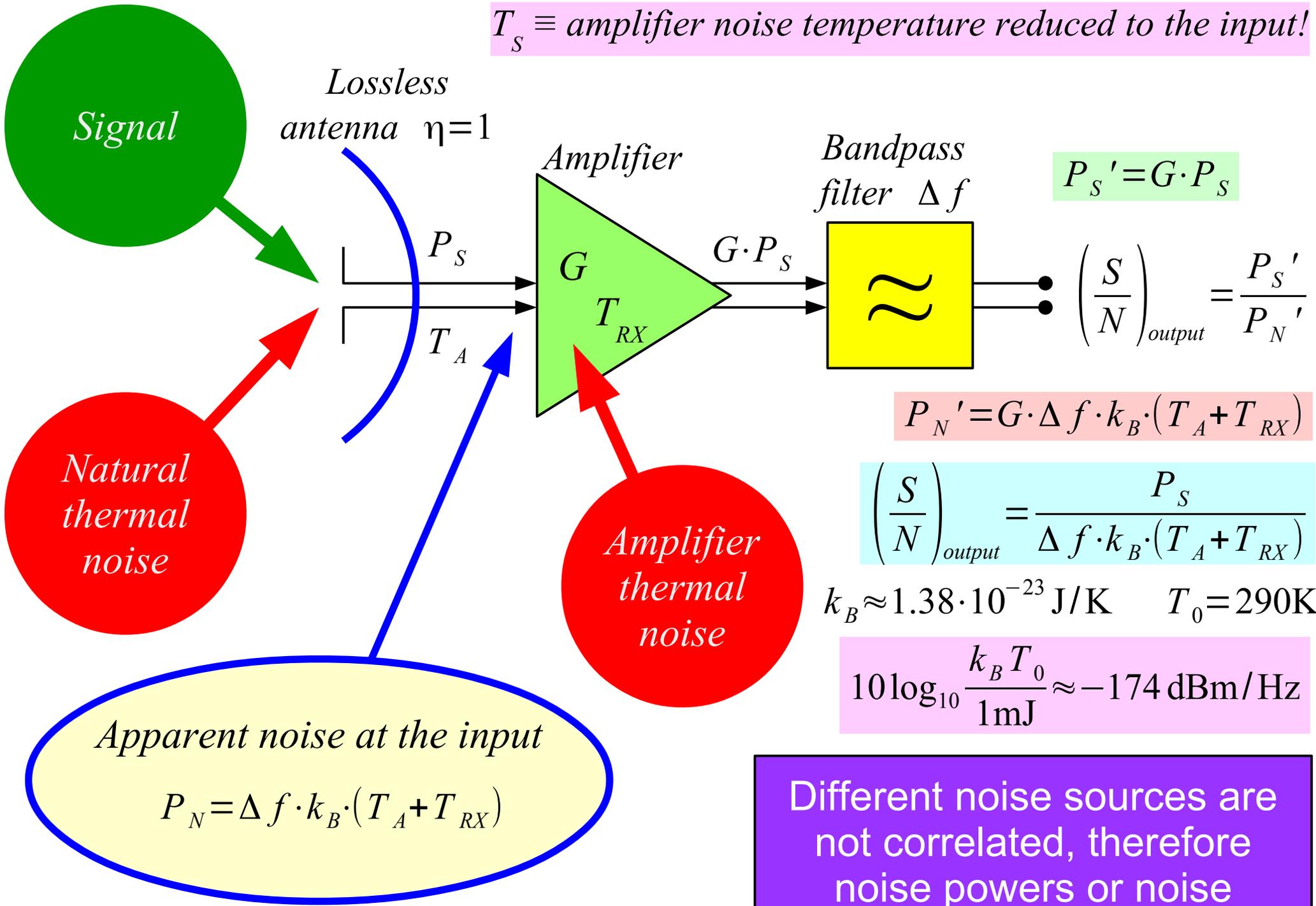
$$\Omega_S = 2\pi [1 - \cos(\alpha_S/2)] \approx \pi \alpha_S^2 / 4 \approx 6 \cdot 10^{-5} \text{srd}$$

Example:
 $D = 20\text{dBi} = 100$

$$\Omega_A \approx \frac{4\pi}{D} = 0.126 \text{srd} \gg \Omega_S$$

8 – Sun-noise example

$T_s \equiv$ amplifier noise temperature reduced to the input!

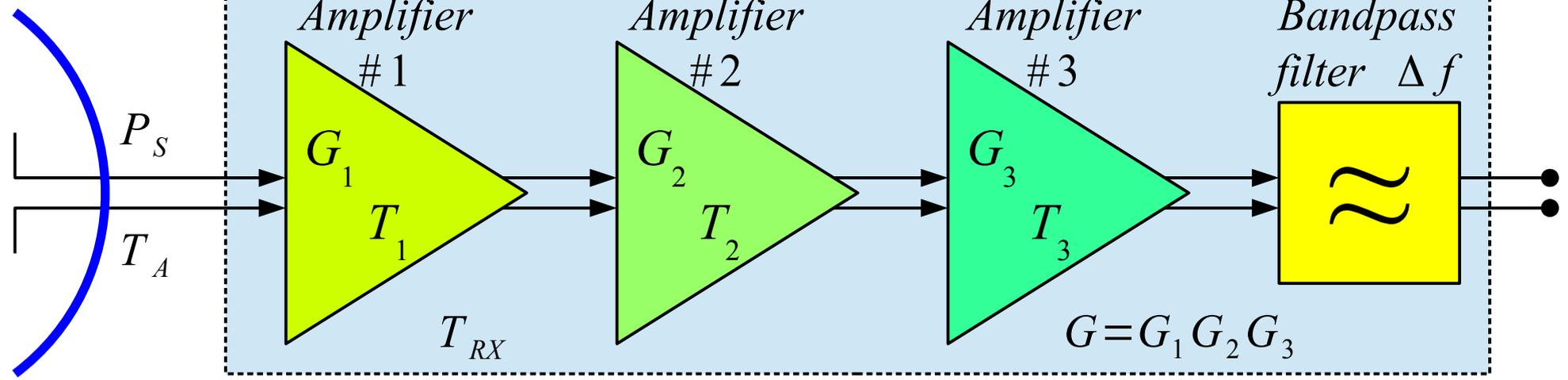


9 – Receiver signal-to-noise ratio

Different noise sources are not correlated, therefore noise powers or noise temperatures are summed!

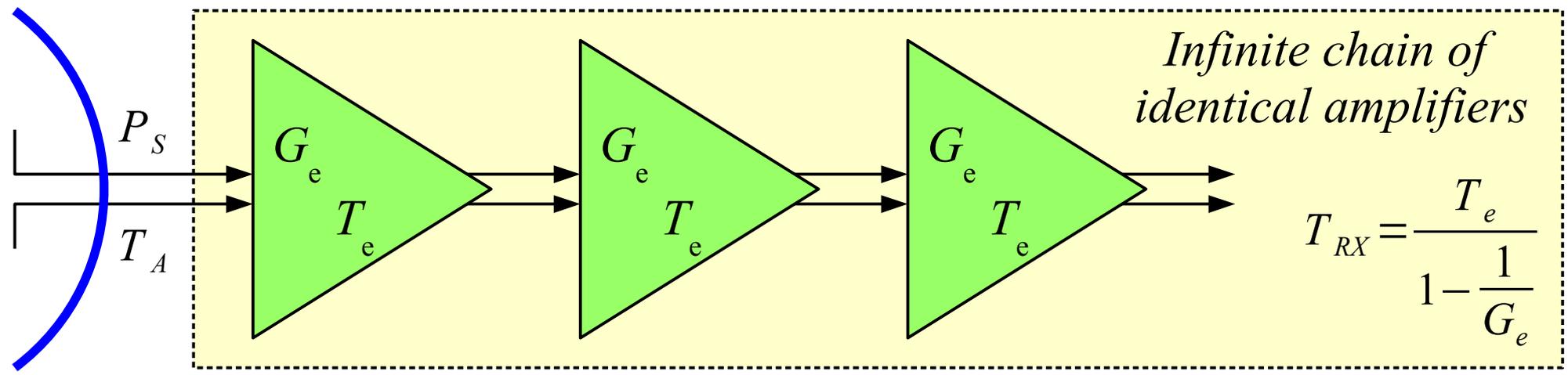
$$P_s' = G_3 G_2 G_1 P_s$$

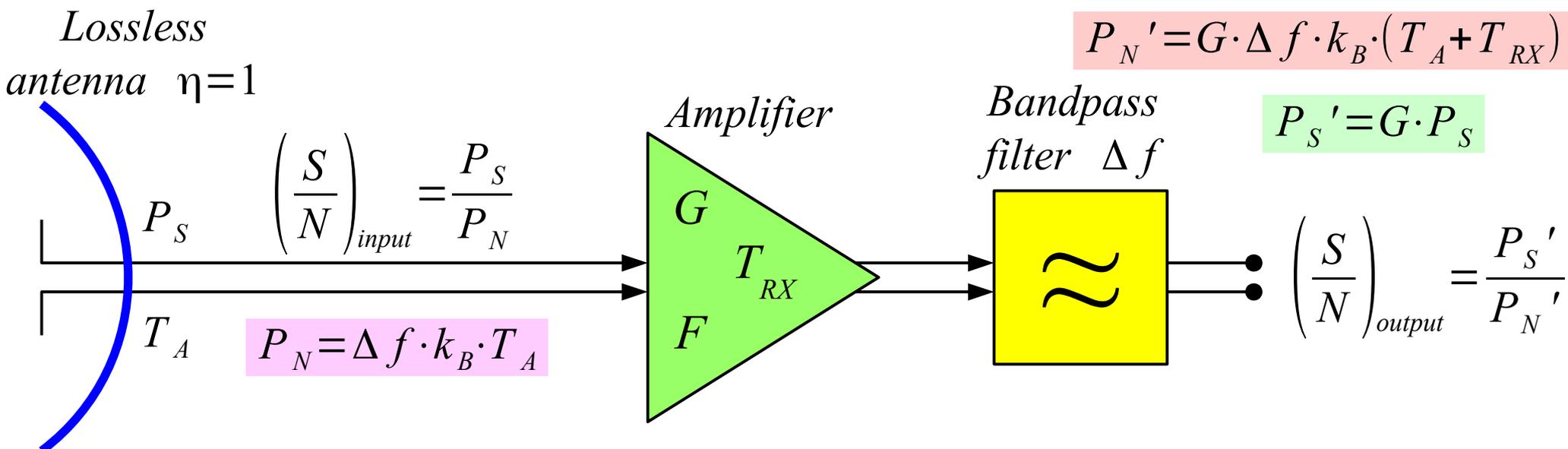
Lossless antenna $\eta = 1$



$$P_N' = \Delta f k_B [G_3 G_2 G_1 (T_A + T_1) + G_3 G_2 T_2 + G_3 T_3]$$

$$P_N' = G_3 G_2 G_1 \Delta f k_B (T_A + T_{RX}) \quad \rightarrow \quad T_{RX} = T_1 + \frac{T_2}{G_1} + \frac{T_3}{G_1 G_2} + \dots$$





Nonsense definition of the noise figure:

$$F = \frac{\left(\frac{S}{N}\right)_{input}}{\left(\frac{S}{N}\right)_{output}} = \frac{\frac{P_S}{\Delta f k_B T_A}}{\frac{G P_S}{G \Delta f k_B (T_A + T_{RX})}} = \frac{T_A + T_{RX}}{T_A} = 1 + \frac{T_{RX}}{T_A}$$

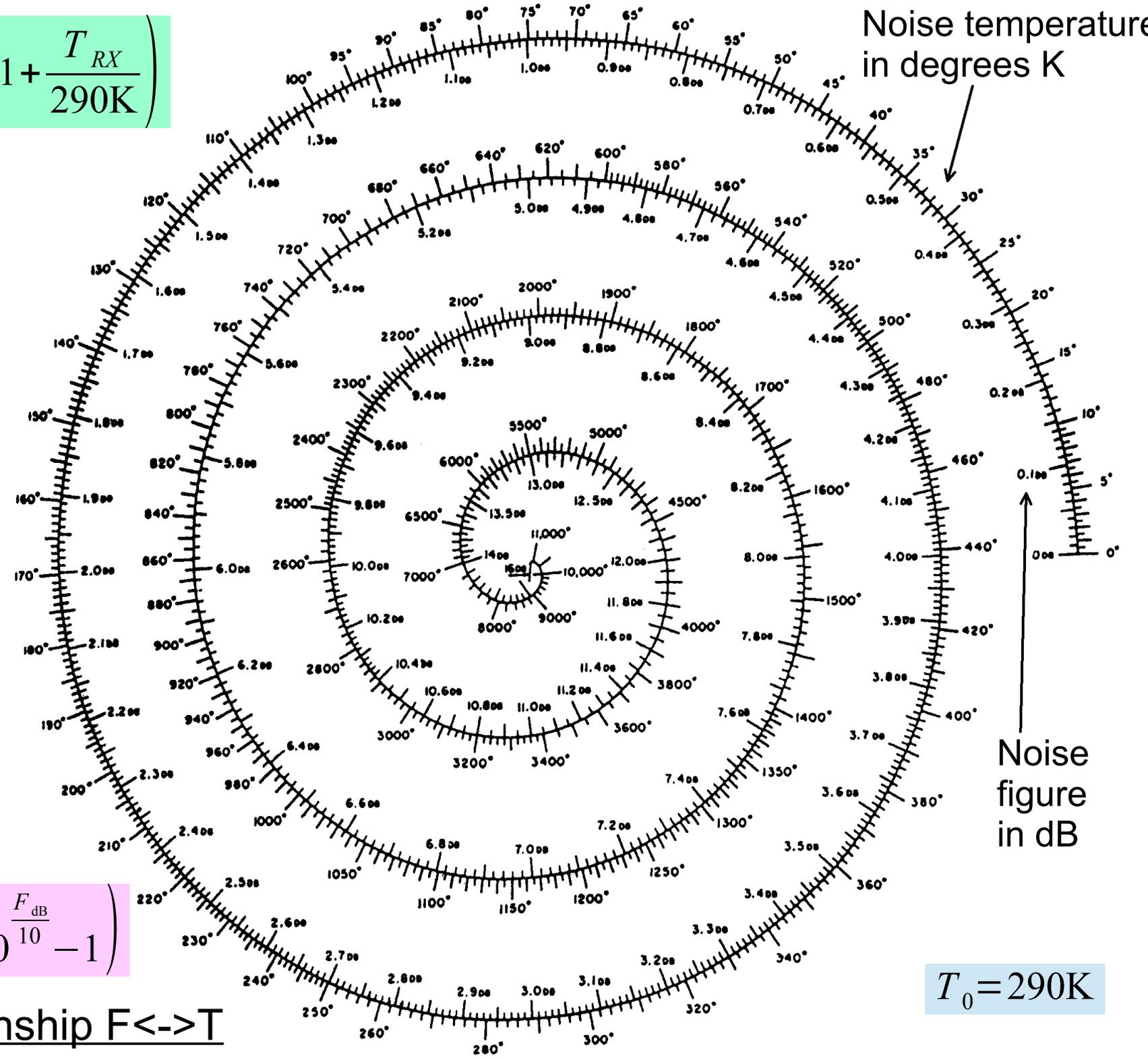
A property of an amplifier can not be a function of T_A !

Sensible definition $F = 1 + \frac{T_{RX}}{T_0}$ @ $T_0 = 290K$ $\leftrightarrow T_{RX} = T_0(F - 1)$

Logarithmic units $F_{dB} = 10 \log_{10} F = 10 \log_{10} \left(1 + \frac{T_{RX}}{T_0}\right)$ $\leftrightarrow T_{RX} = T_0 \left(10^{\frac{F_{dB}}{10}} - 1\right)$

11 – Amplifier noise figure

$$F_{dB} = 10 \log_{10} \left(1 + \frac{T_{RX}}{290K} \right)$$



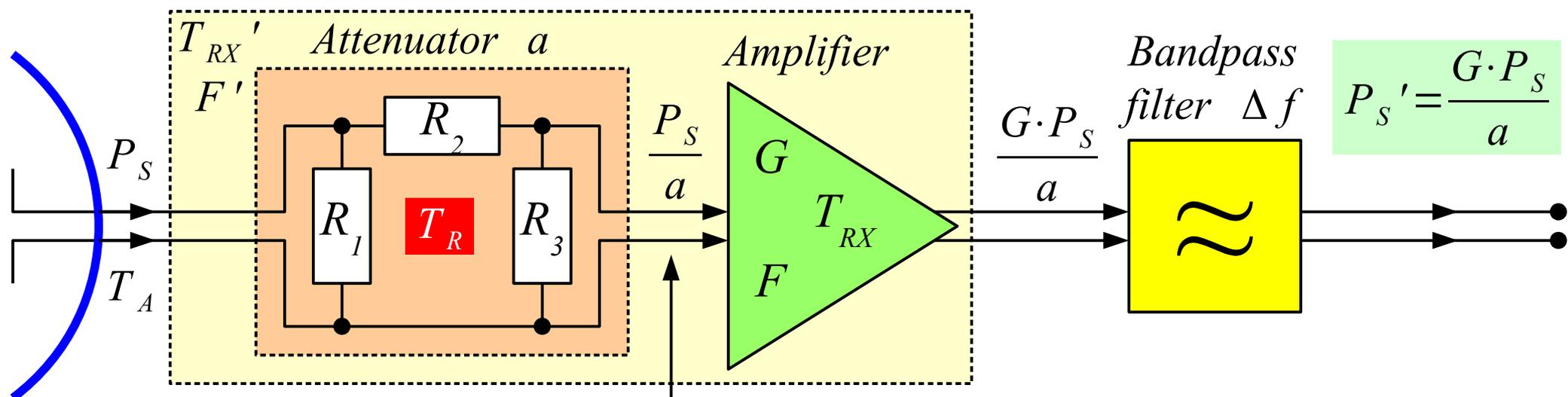
Noise temperature in degrees K

Noise figure in dB

$$T_{RX} = 290K \left(10^{\frac{F_{dB}}{10}} - 1 \right)$$

$$T_0 = 290K$$

12 – Relationship F <-> T



Lossless antenna $\eta=1$

$$\frac{T_A}{a} + T_R \left(1 - \frac{1}{a}\right)$$

$$P_N' = G \cdot \Delta f \cdot k_B \cdot \left[\frac{T_A}{a} + T_R \left(1 - \frac{1}{a}\right) + T_{RX} \right]$$

$$T_{RX}' = T_R(a-1) + a T_{RX} \quad \left(\frac{S}{N} \right)_{output} = \frac{P_S'}{P_N'} = \frac{P_S}{\Delta f \cdot k_B \cdot [T_A + T_R(a-1) + a T_{RX}]}$$

$$F' = 1 + \frac{T_{RX}'}{T_0} = 1 + \frac{T_R}{T_0}(a-1) + a \frac{T_{RX}}{T_0}$$

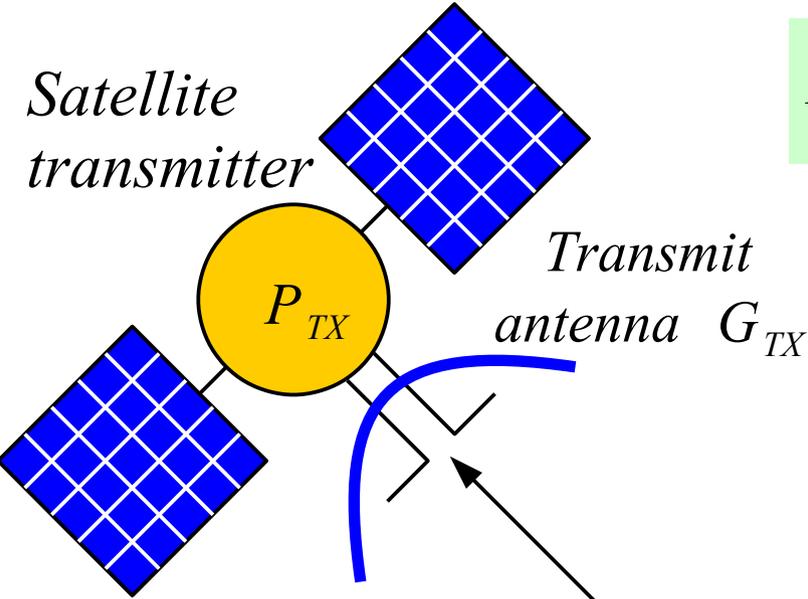
Frequent case $T_R \approx T_0 = 290K$

$$F' \approx a + a \frac{T_{RX}}{T_0} = a \left(1 + \frac{T_{RX}}{T_0} \right) = a \cdot F$$

$$F_{dB}' \approx a_{dB} + F_{dB}$$

- Attenuator examples $T_R \approx T_0 = 290K$
- $$F' \approx a \cdot F \quad F_{dB}' \approx a_{dB} + F_{dB}$$
- (1) lossy antenna $a_{dB} = -10 \log_{10} \eta$
 - (2) lossy transmission line a_{dB}
 - (3) lossy bandpass filter a_{dB}
 - (4) passive-mixer loss a_{dB}

Satellite transmitter



Free-space radio link $P_S = P_{TX} \cdot G_{TX} \cdot G_{RX} \cdot \left(\frac{\lambda}{4\pi r}\right)^2$

Transmitter

Receiver

$$\left(\frac{S}{N}\right)_{output} = P_{TX} \cdot G_{TX} \cdot \frac{1}{\Delta f \cdot k_B} \cdot \left(\frac{\lambda}{4\pi r}\right)^2 \cdot \frac{G_{RX}}{(T_A + T_{RX})}$$

System

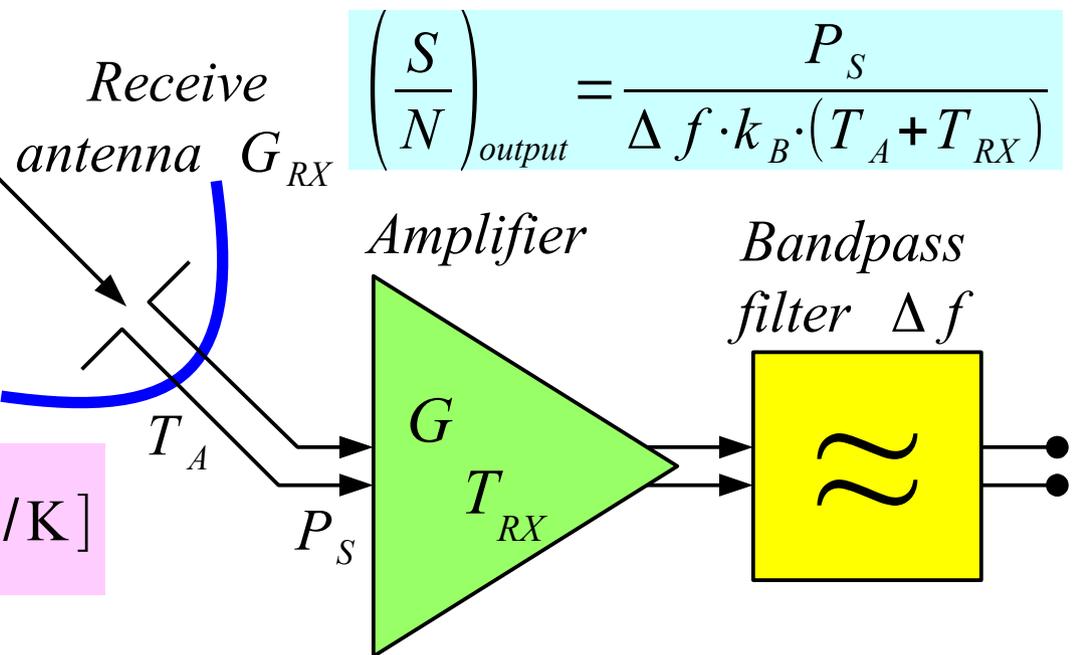
Receiving ground station

$$(G/T) = \frac{G_{RX}}{(T_A + T_{RX})} \text{ [K}^{-1}\text{]}$$

$$(G/T)_{dB/K} = 10 \log_{10} \frac{G_{RX} \cdot 1K}{(T_A + T_{RX})} \text{ [dB/K]}$$

$$(G/T)_{dB/K} = G_{RX \text{ dB}} - 10 \log_{10} \frac{T_A + T_{RX}}{1K} \text{ [dB/K]}$$

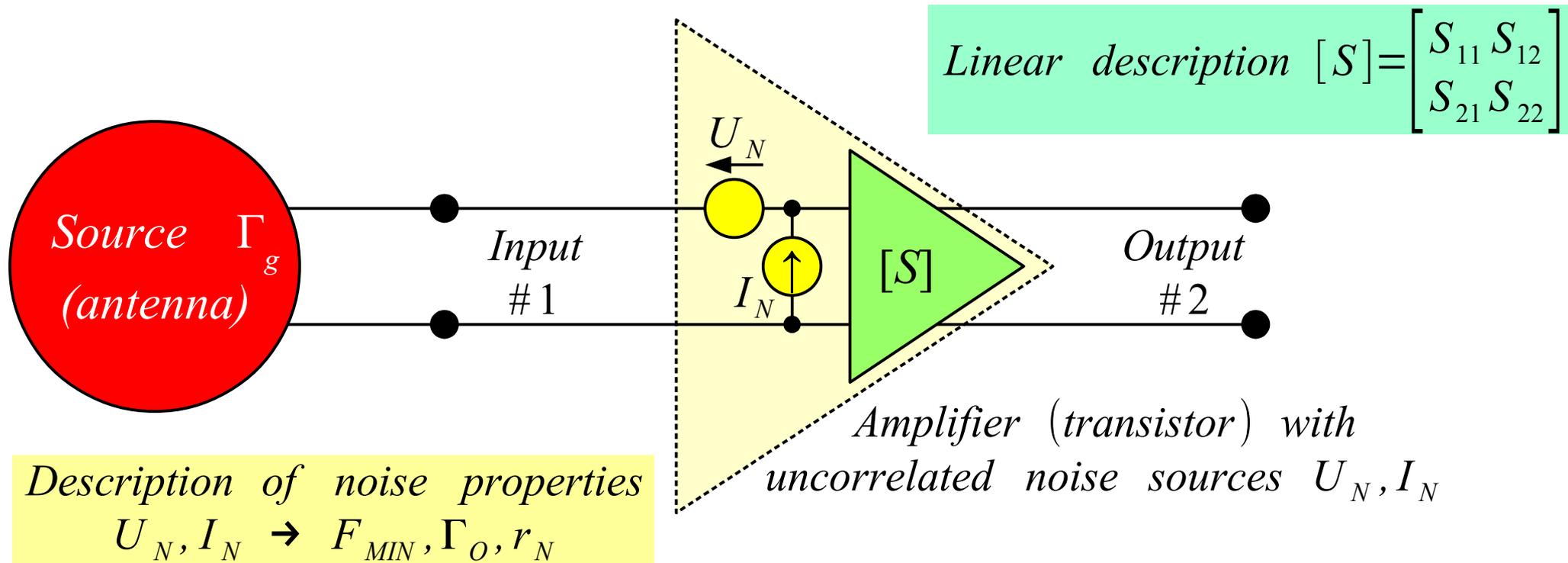
r



$$\left(\frac{S}{N}\right)_{output} = \frac{P_S}{\Delta f \cdot k_B \cdot (T_A + T_{RX})}$$

Receiving ground station

Amplifier device	Gain G [dB]	Noise temperature T_{RX} [K]	Noise figure F_{dB} [dB]
Vacuum tube with control grid (triode, pentode)	10↔20	1600↔9000	8↔15
Vacuum tube with speed modulation (klystron, TWT)	20↔50	3000↔30000	10↔20
Parametric amplifier (room temperature)	10↔15	75↔300	1↔3
Si BJT, JFET or MOSFET (room temperature)	10↔20	75↔300	1↔3
GaAs FET or HEMT (room temperature)	10↔15	20↔120	0.3↔1.5
GaAs FET ali HEMT (liquid-nitrogen 77K)	10↔15	7↔35	0.1↔0.5
Si or GaAs MMIC amplifier	10↔25	170↔1600	2↔8
Operational amplifier	40↔100	10^4 ↔ 10^9	16↔66



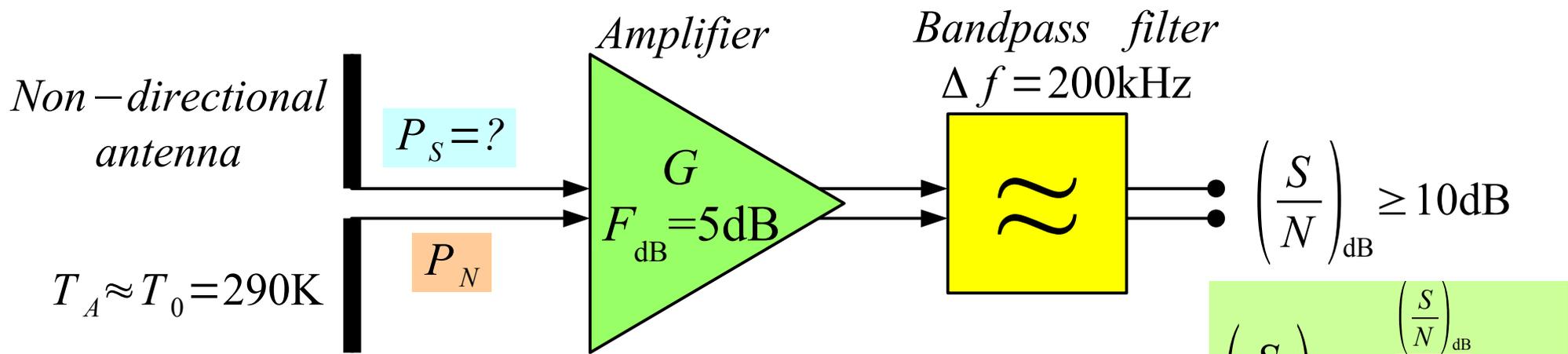
$$F = F_{MIN} + 4 \frac{R_N}{Z_K} \cdot \frac{|\Gamma_g - \Gamma_O|^2}{(1 - |\Gamma_g|^2) \cdot |1 + \Gamma_O|^2} = F_{MIN} + 4 r_N \cdot \frac{|\Gamma_g - \Gamma_O|^2}{(1 - |\Gamma_g|^2) \cdot |1 + \Gamma_O|^2}$$

$F_{MIN} \equiv$ lowest noise figure at $\Gamma_g = \Gamma_O$ in linear units (not in dB!)

$\Gamma_O \equiv$ optimum source reflectivity for F_{MIN} (unrelated to $[S]$!)

$r_N = \frac{R_N}{Z_K} \equiv$ normalized noise resistance (usually $Z_K = 50 \Omega$)

Low-noise devices are usually NOT unconditionally stable amplifiers!



$$T_{RX} = T_0 \cdot \left(10^{\frac{F_{\text{dB}}}{10}} - 1\right) = 290\text{K} \cdot (3.162 - 1) = 627\text{K} \quad k_B \approx 1.38 \cdot 10^{-23} \text{ J/K}$$

$$P_N = \Delta f \cdot k_B \cdot (T_A + T_{RX}) = 200\text{kHz} \cdot 1.38 \cdot 10^{-23} \text{ J/K} \cdot (290\text{K} + 627\text{K}) = 2.531 \cdot 10^{-15} \text{ W}$$

$$P_S = P_N \cdot \left(\frac{S}{N}\right) = 2.531 \cdot 10^{-15} \text{ W} \cdot 10 = 2.531 \cdot 10^{-14} \text{ W} \quad P_{S\text{dBm}} = 10 \log_{10} \frac{P_S}{1\text{mW}} = -106\text{dBm}$$

Simplified calculation exclusively in case $T_A \approx T_0 = 290\text{K}$

$$P_{S\text{dBm}} \approx (S/N)_{\text{dB}} + (\Delta f)_{\text{dB}\cdot\text{Hz}} + (k_B T_0)_{\text{dBm/Hz}} + F_{\text{dB}}$$

$$(k_B T_0)_{\text{dBm/Hz}} = 10 \log_{10} \frac{k_B T_0}{1\text{mJ}} \approx -174 \text{ dBm/Hz} \quad (\Delta f)_{\text{dB}\cdot\text{Hz}} = 10 \log_{10} \left(\frac{\Delta f}{1\text{Hz}}\right) = 53 \text{ dB}\cdot\text{Hz}$$

$$P_{S\text{dBm}} \approx 10\text{dB} + 53\text{dB}\cdot\text{Hz} - 174\text{dBm/Hz} + 5\text{dB} = -106\text{dBm}$$

Two different receivers #1 and #2:

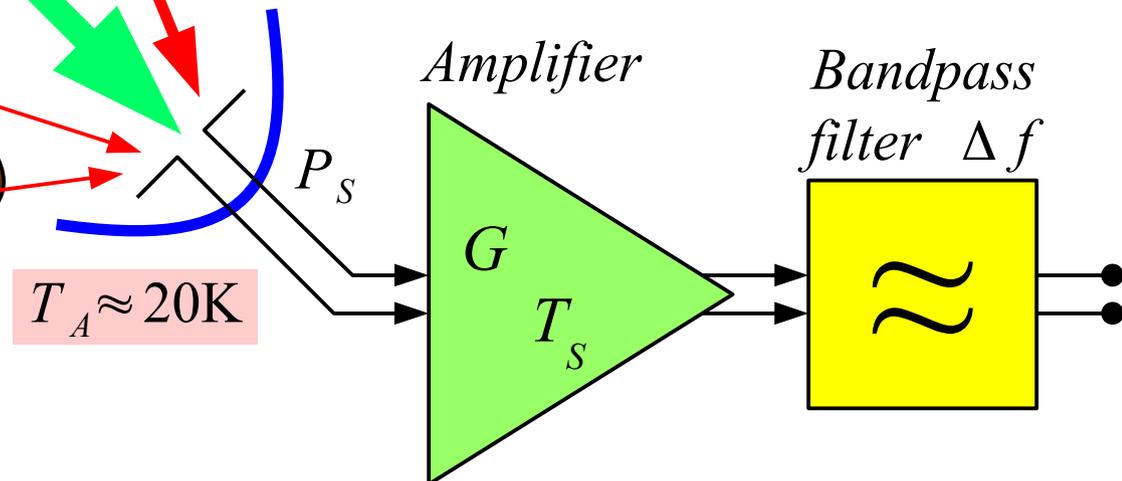
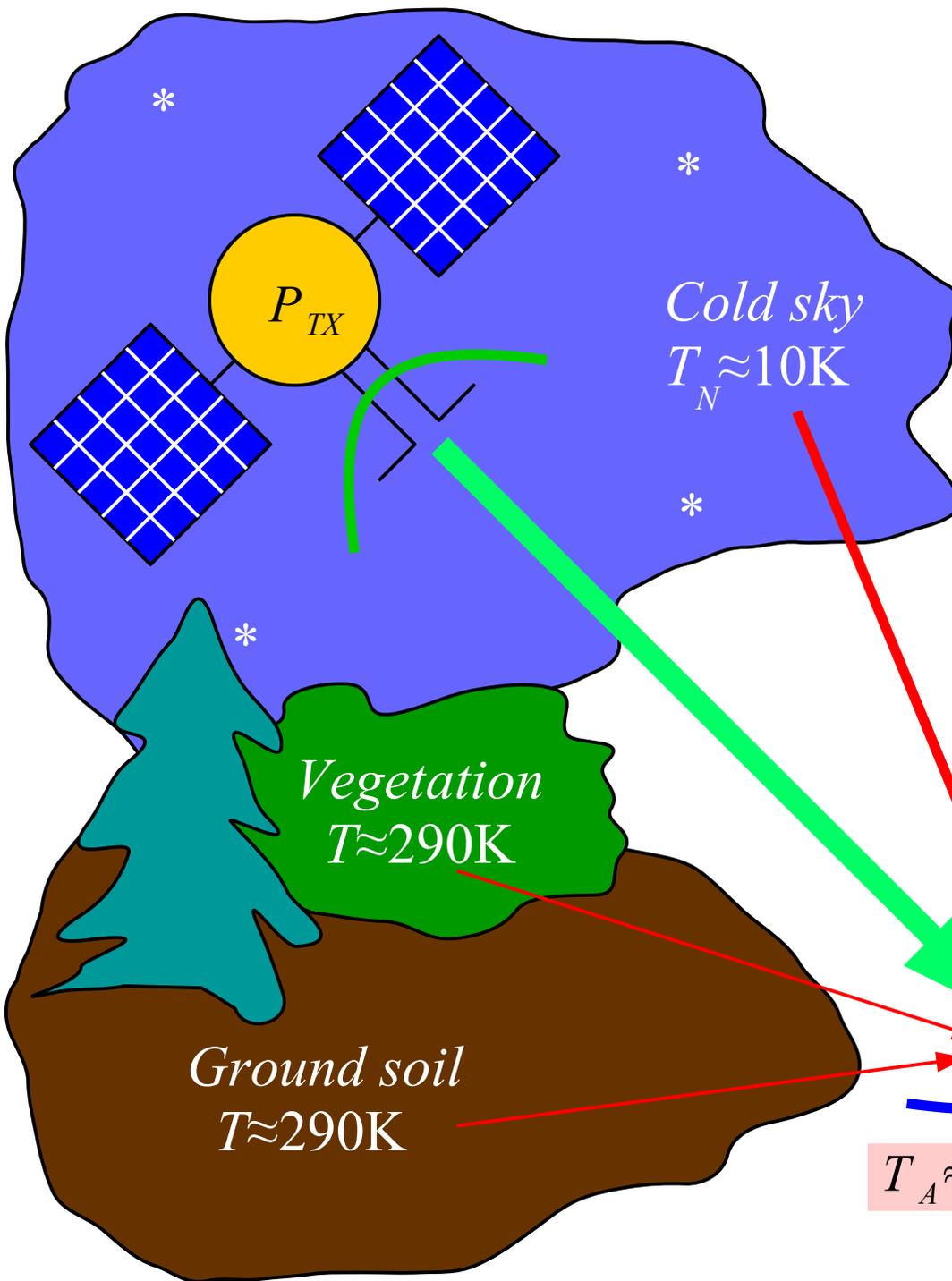
$$F_1 = 1\text{dB} \rightarrow T_{RX1} = 75\text{K}$$

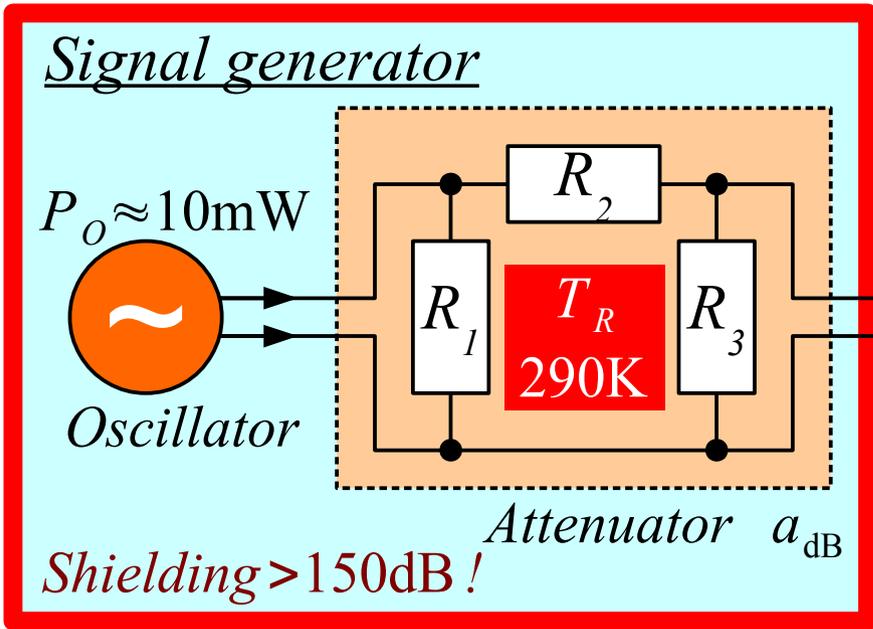
$$F_2 = 0.5\text{dB} \rightarrow T_{RX2} = 35\text{K}$$

$$\Delta F_{\text{dB}} = F_1 - F_2 = 0.5\text{dB}$$

$$\Delta \left(\frac{S}{N} \right)_{\text{dB}} = 10 \log_{10} \left[\frac{T_A + T_{RX2}}{T_A + T_{RX1}} \right]$$

$$\Delta \left(\frac{S}{N} \right)_{\text{dB}} = 10 \log_{10} \left[\frac{20\text{K} + 75\text{K}}{20\text{K} + 35\text{K}} \right] = 2.37\text{dB}$$





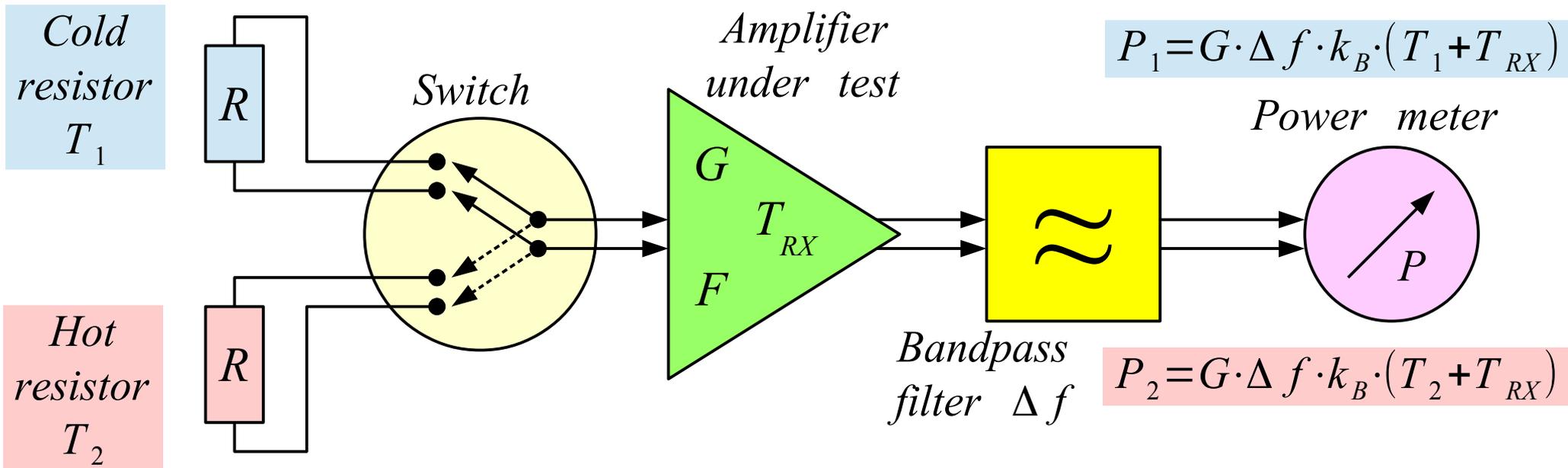
$50\text{dB} < a_{\text{dB}} < 150\text{dB}$

Coupling via radiation?

Additional requirements for the test source (signal generator) for sensitivity measurements of radio receivers:

- (1) Shielding > 150dB
- (2) $T_R = T_A = T_0 = 290\text{K}$

- (1) Required S/N before demodulation?
- (2) Required S/N after demodulation?
- (3) Required BER?



The unknowns $G \cdot \Delta f \cdot k_B$ cancel in the Y ratio!

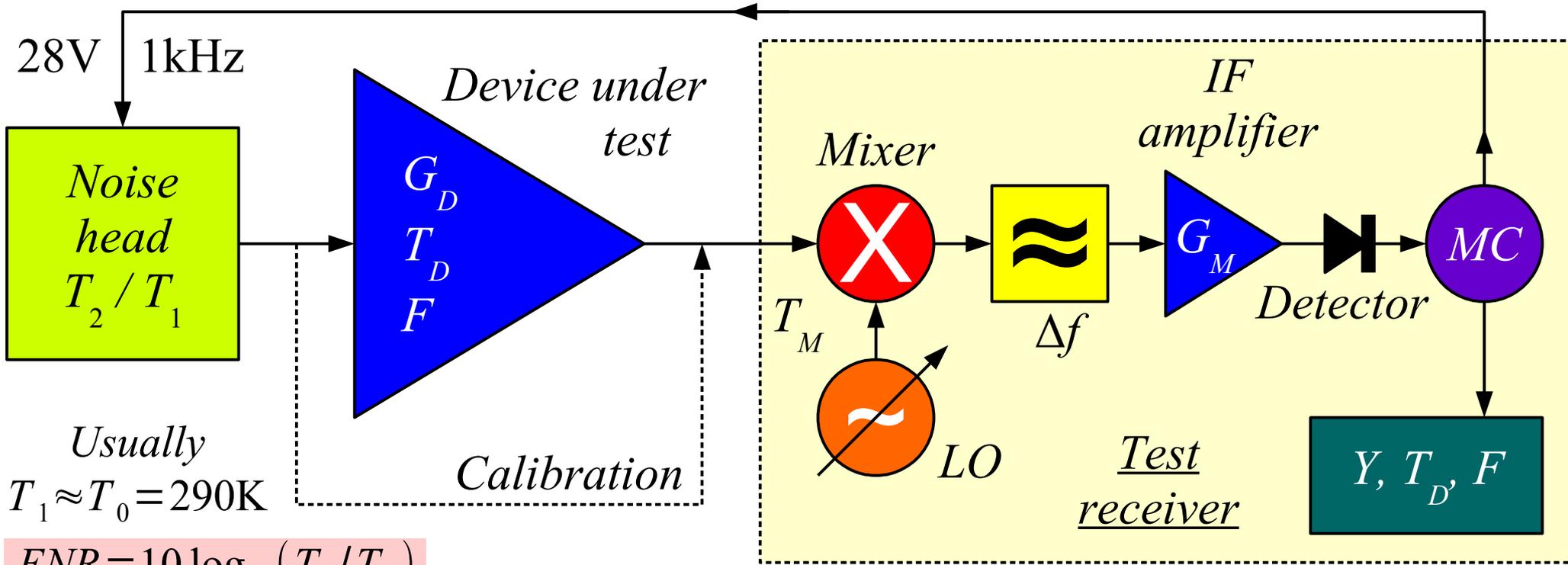
$$Y = \frac{P_2}{P_1} = \frac{T_2 + T_{RX}}{T_1 + T_{RX}}$$

$$T_{RX} = \frac{T_2 - Y \cdot T_1}{Y - 1}$$

$$T_0 = 290\text{K}$$

$$F_{dB} = 10 \log_{10} \left[1 + \frac{T_2 - Y \cdot T_1}{(Y - 1) \cdot T_0} \right]$$

Resistor type	Temperature
Antenna into cold sky	$\sim 20\text{K}$
Liquid N_2 cooled R	$\sim 77\text{K}$
Antenna into absorber	$\sim 290\text{K}$
R at room temperature	$\sim 290\text{K}$
Light-bulb filament as R	$\sim 2000\text{K}$
Ionized gas as R	$\sim 10^4\text{K}$
Avalanche breakdown	$\sim 10^6\text{K}$



Usually
 $T_1 \approx T_0 = 290\text{K}$

$$ENR = 10 \log_{10}(T_2/T_1)$$

$G_M \Delta f \equiv \text{uncertain!}$

Two measurements without calibration:

$$Y = \frac{P_2}{P_1} = \frac{T_2 + T_D + T_M/G_D}{T_1 + T_D + T_M/G_D}$$

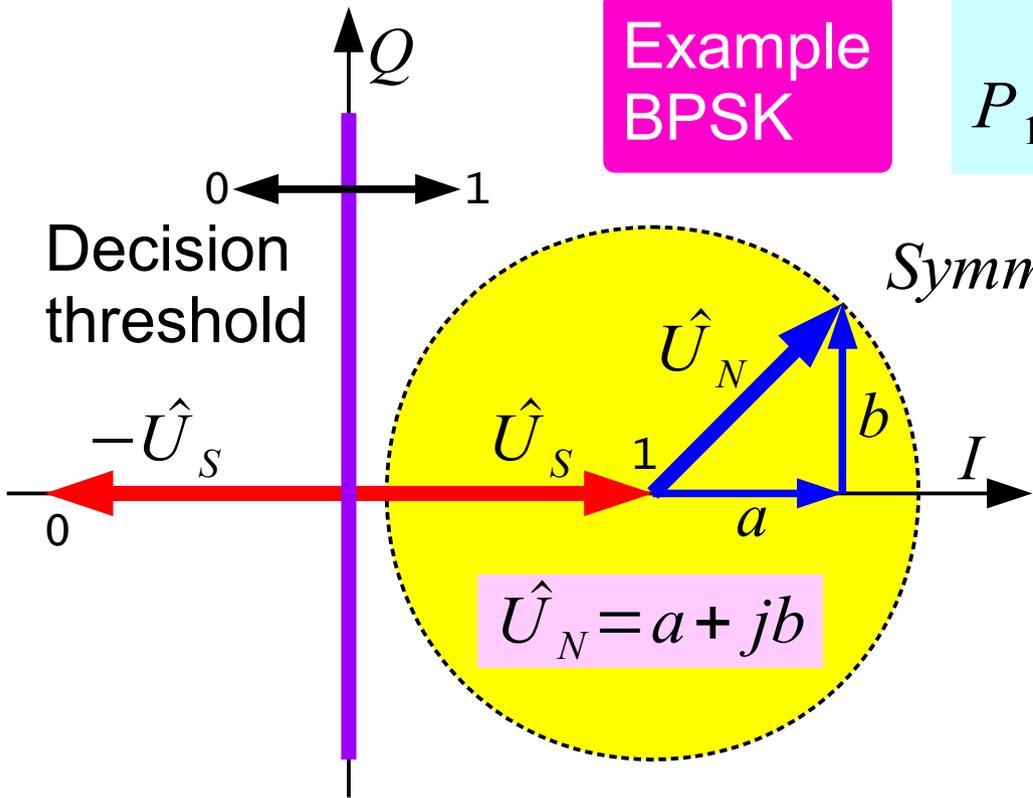
$$T_D = \frac{T_2 - Y \cdot T_1}{Y - 1} - \frac{T_M}{G_D} \leftarrow \text{known } G_D$$

$$F_{dB} = 10 \log_{10} \left[1 + \frac{1}{T_0} \cdot \left(\frac{T_2 - Y \cdot T_1}{Y - 1} - \frac{T_M}{G_D} \right) \right]$$

Four measurements with calibration:

- (1) $P_1 = G_M G_D \Delta f k_B (T_1 + T_D + T_M/G_D)$
- (2) $P_2 = G_M G_D \Delta f k_B (T_2 + T_D + T_M/G_D)$
- (3) $P_3 = G_M \Delta f k_B (T_1 + T_M)$
- (4) $P_4 = G_M \Delta f k_B (T_2 + T_M)$

Solve 4 equations for 4 unknowns:
 T_D, G_D, T_M and $(G_M \Delta f k_B)$



Example BPSK

$$P_{1 \rightarrow 0} = \int_{-\infty}^{-|\hat{U}_s|} p(a) da$$

$$P_{0 \rightarrow 1} = \int_{|\hat{U}_s|}^{\infty} p(a) da$$

Symmetric threshold: $P_{1 \rightarrow 0} = P_{0 \rightarrow 1} = BER$

$$BER = \int_{|\hat{U}_s|}^{\infty} \frac{1}{\sqrt{\pi \langle |\hat{U}_N|^2 \rangle}} e^{-\frac{a^2}{\langle |\hat{U}_N|^2 \rangle}} da$$

$$\text{erfc}(x) = \frac{2}{\sqrt{\pi}} \int_x^{\infty} e^{-u^2} du$$

Gaussian distribution of probability density of the in-phase a and quadrature jb noise components

$$p(a) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{a^2}{2\sigma^2}}$$

$$p(b) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{b^2}{2\sigma^2}}$$

$$BER = \frac{1}{2} \text{erfc} \left(\frac{|\hat{U}_s|}{\sqrt{\langle |\hat{U}_N|^2 \rangle}} \right)$$

$$P_s = \alpha |\hat{U}_s|^2$$

$$P_n = \alpha \langle |\hat{U}_N|^2 \rangle$$

$$\langle |\hat{U}_N|^2 \rangle = \langle a^2 \rangle + \langle b^2 \rangle = 2\sigma^2$$

$$BER = \frac{1}{2} \text{erfc} \left(\sqrt{\frac{P_s}{P_n}} \right)$$

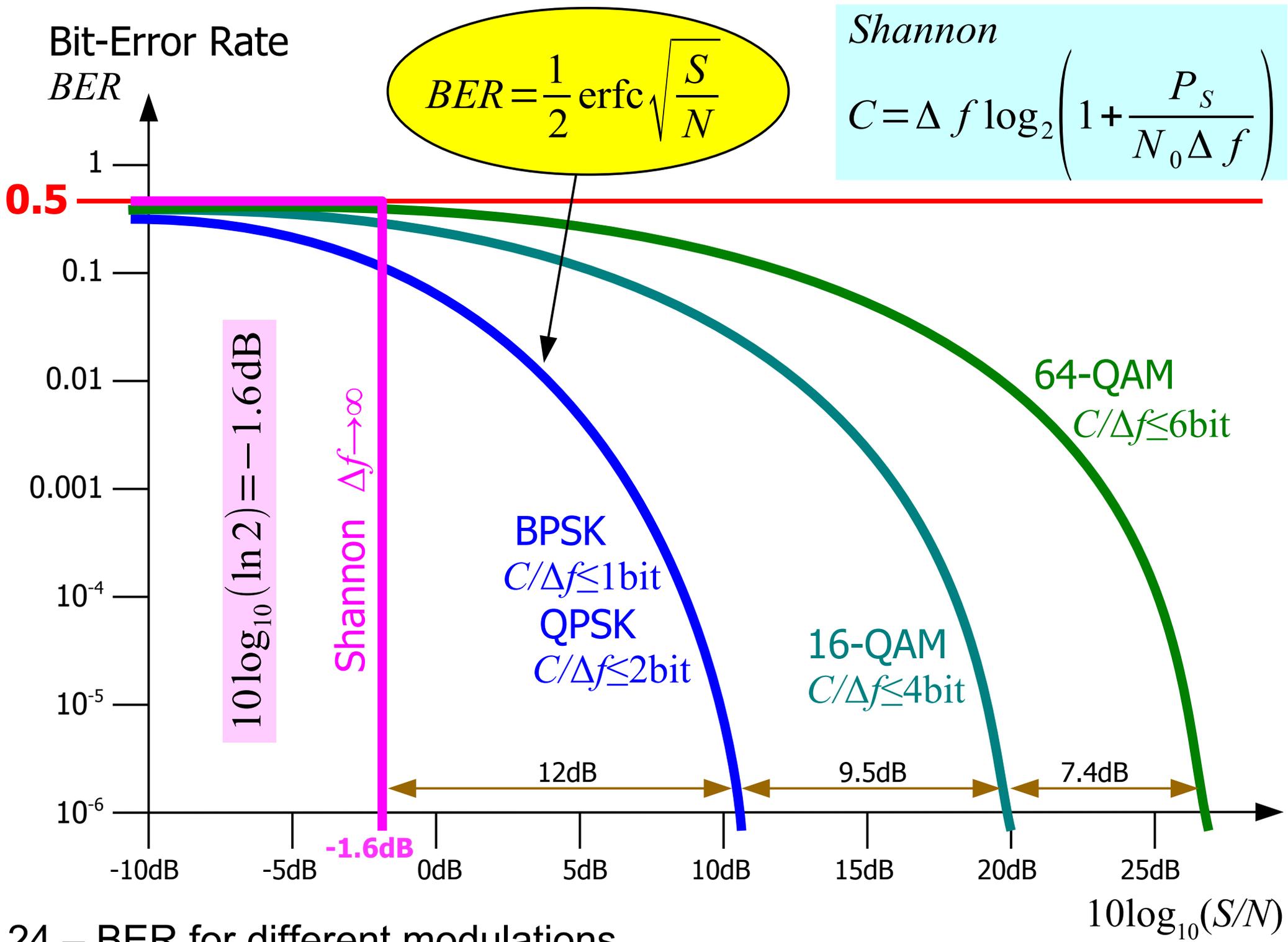
$(S/N)_{dB}$	BER
-5dB	23.6%
-4dB	18.6%
-3dB	15.9%
-2dB	13.1%
-1dB	10.4%
-0dB	7.9%
1dB	5.7%
2dB	3.8%
3dB	2.3%
4dB	1.3%
5dB	0.6%
6dB	0.24%
7dB	$7.7 \cdot 10^{-4}$
$(S/N)_{dB}$	BER

$(S/N)_{dB}$	BER
8dB	$1.9 \cdot 10^{-4}$
9dB	$3.4 \cdot 10^{-5}$
10dB	$3.9 \cdot 10^{-6}$
11dB	$2.6 \cdot 10^{-7}$
12dB	$9 \cdot 10^{-9}$
13dB	$1.3 \cdot 10^{-10}$
14dB	$6.8 \cdot 10^{-13}$
15dB	$9.2 \cdot 10^{-16}$
16dB	$2.3 \cdot 10^{-19}$
17dB	$6.8 \cdot 10^{-24}$
18dB	$1.4 \cdot 10^{-29}$
19dB	10^{-36}
20dB	10^{-45}
$(S/N)_{dB}$	BER

BER	$(S/N)_{dB}$
30%	-8.6dB
10%	-0.8dB
3%	2.5dB
1%	4.3dB
0.3%	5.8dB
0.1%	6.8dB
$3 \cdot 10^{-4}$	7.7dB
10^{-4}	8.4dB
$3 \cdot 10^{-5}$	9.1dB
10^{-5}	9.6dB
$3 \cdot 10^{-6}$	10.1dB
10^{-6}	10.5dB
$3 \cdot 10^{-7}$	11dB
BER	$(S/N)_{dB}$

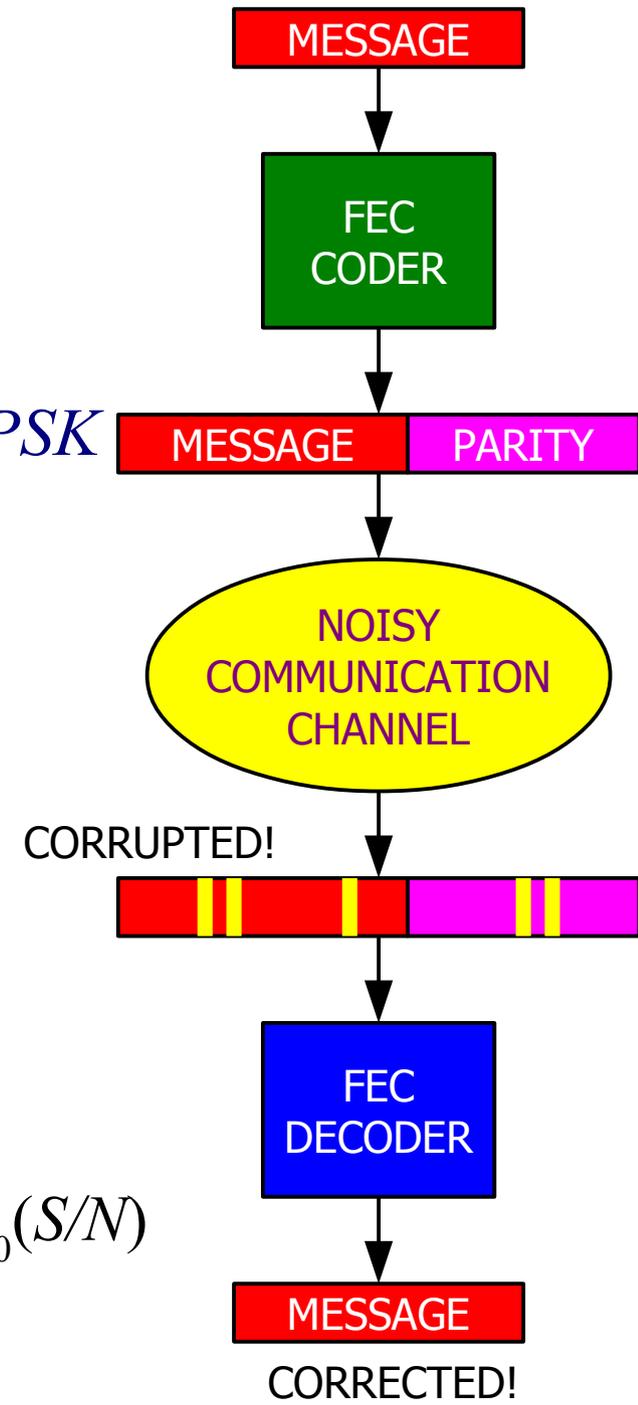
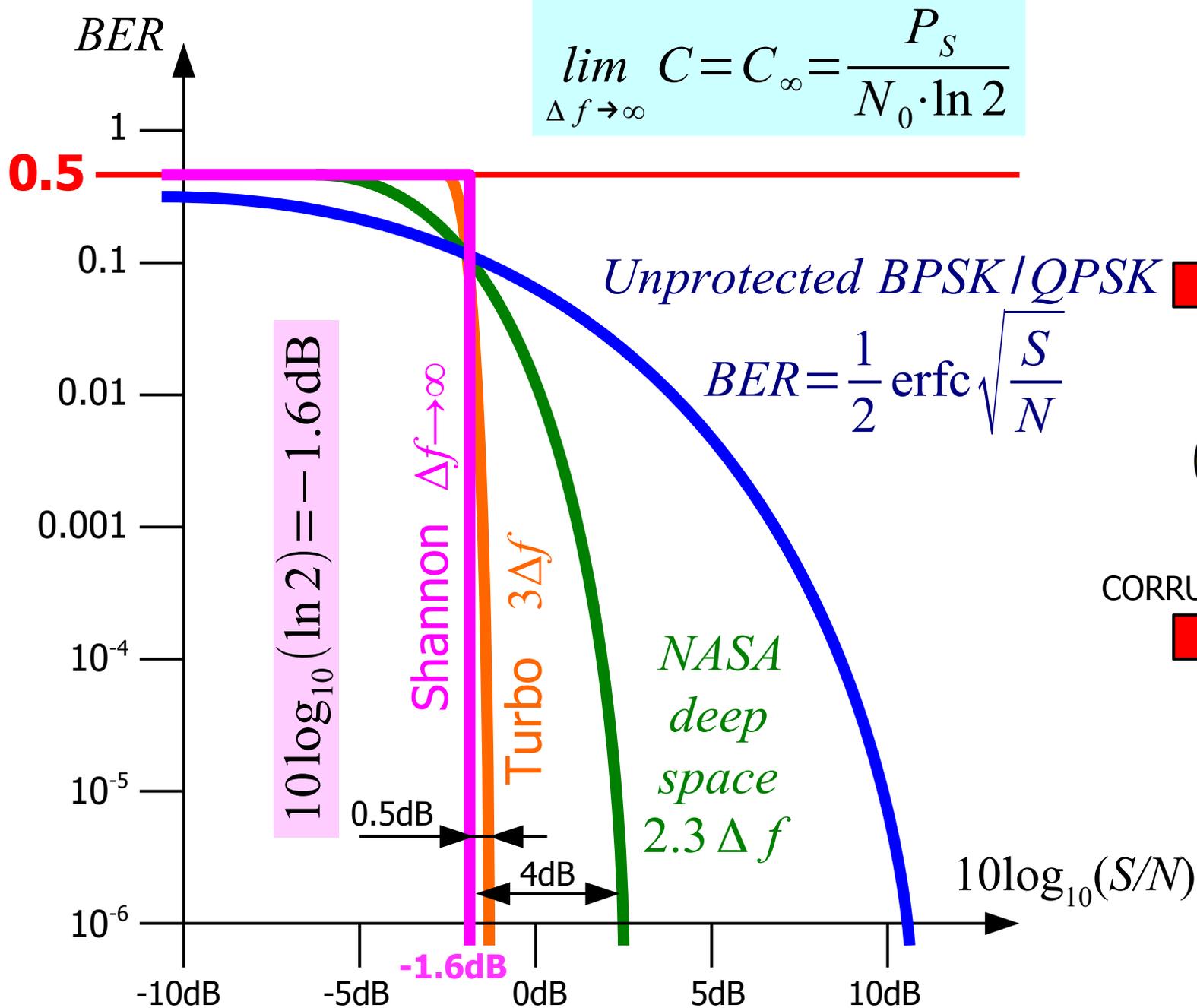
BER	$(S/N)_{dB}$
10^{-7}	11.3dB
$3 \cdot 10^{-8}$	11.7dB
10^{-8}	12dB
$3 \cdot 10^{-9}$	12.3dB
10^{-9}	12.6dB
10^{-10}	13.1dB
10^{-11}	13.5dB
10^{-12}	13.9dB
10^{-13}	14.3dB
10^{-14}	14.7dB
10^{-15}	15dB
10^{-16}	15.3dB
10^{-17}	15.6dB
BER	$(S/N)_{dB}$

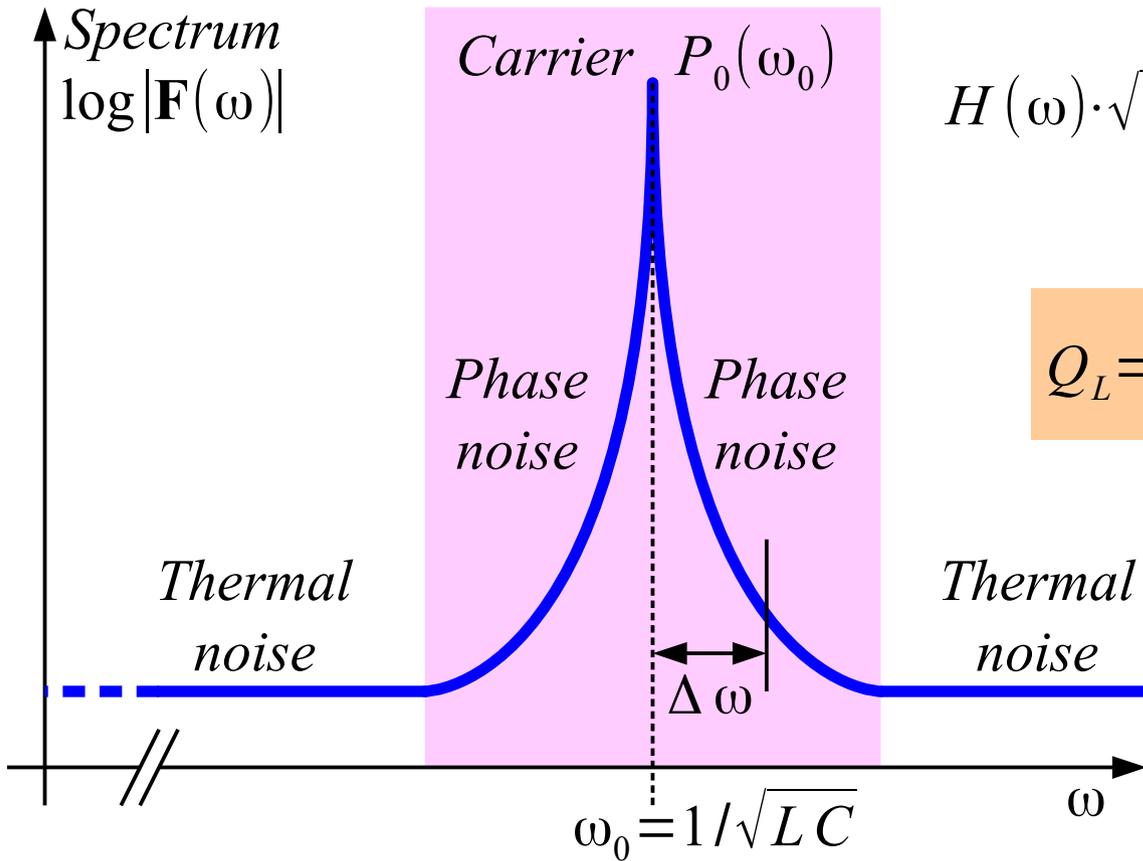
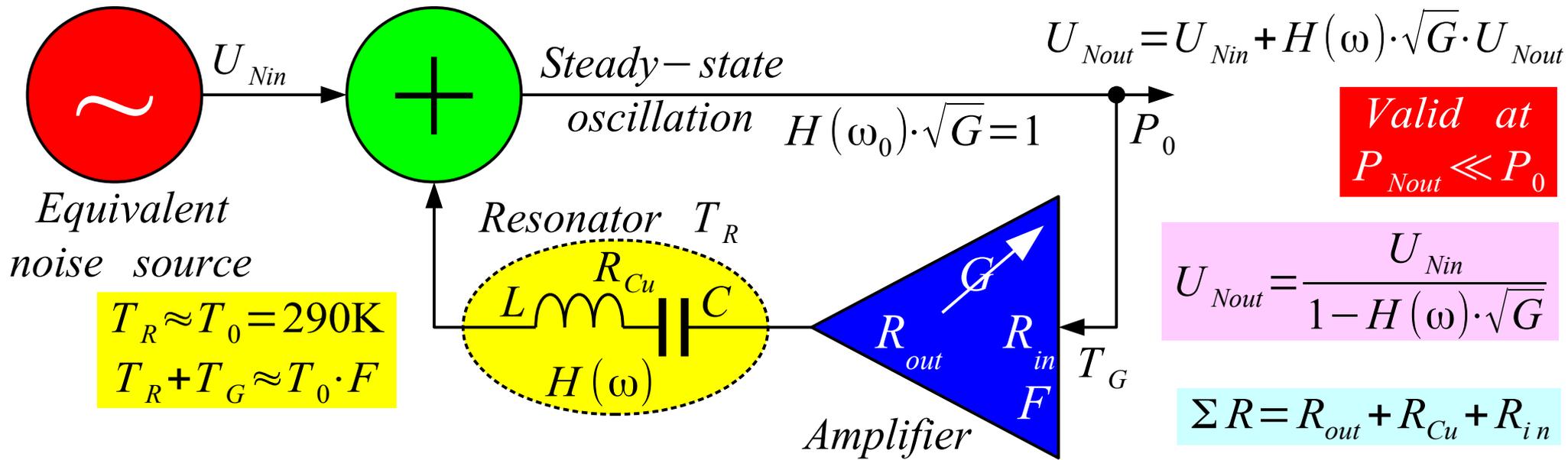
23 – BER <-> S/N table for BPSK



24 – BER for different modulations

Bit-Error Rate





$$H(\omega) \cdot \sqrt{G} = \frac{\Sigma R}{\Sigma R + j\omega L + \frac{1}{j\omega C}} \approx \frac{1}{1 + j2Q_L \frac{\Delta\omega}{\omega_0}}$$

$$Q_L = \frac{\omega_0 L}{\Sigma R}$$

$$U_{Nout} \approx U_{Nin} \cdot \left(1 + \frac{\omega_0}{j2Q_L \Delta\omega} \right)$$

$$P_{Nout} \approx P_{Nin} \cdot \left[1 + \left(\frac{\omega_0}{2Q_L \Delta\omega} \right)^2 \right]$$

Amplitude and phase noise

$$P_{Nout} \approx P_{Nin} \cdot \left[1 + \left(\frac{f_0}{2Q_L \Delta f} \right)^2 \right]$$

Normalized
phase-noise
spectral density

Saturation removes amplitude
noise $P_\varphi = P_{Nout}/2$

$$\frac{dP_{Nin}}{df} = N_0 = k_B(T_R + T_G) \approx k_B T_0 F$$

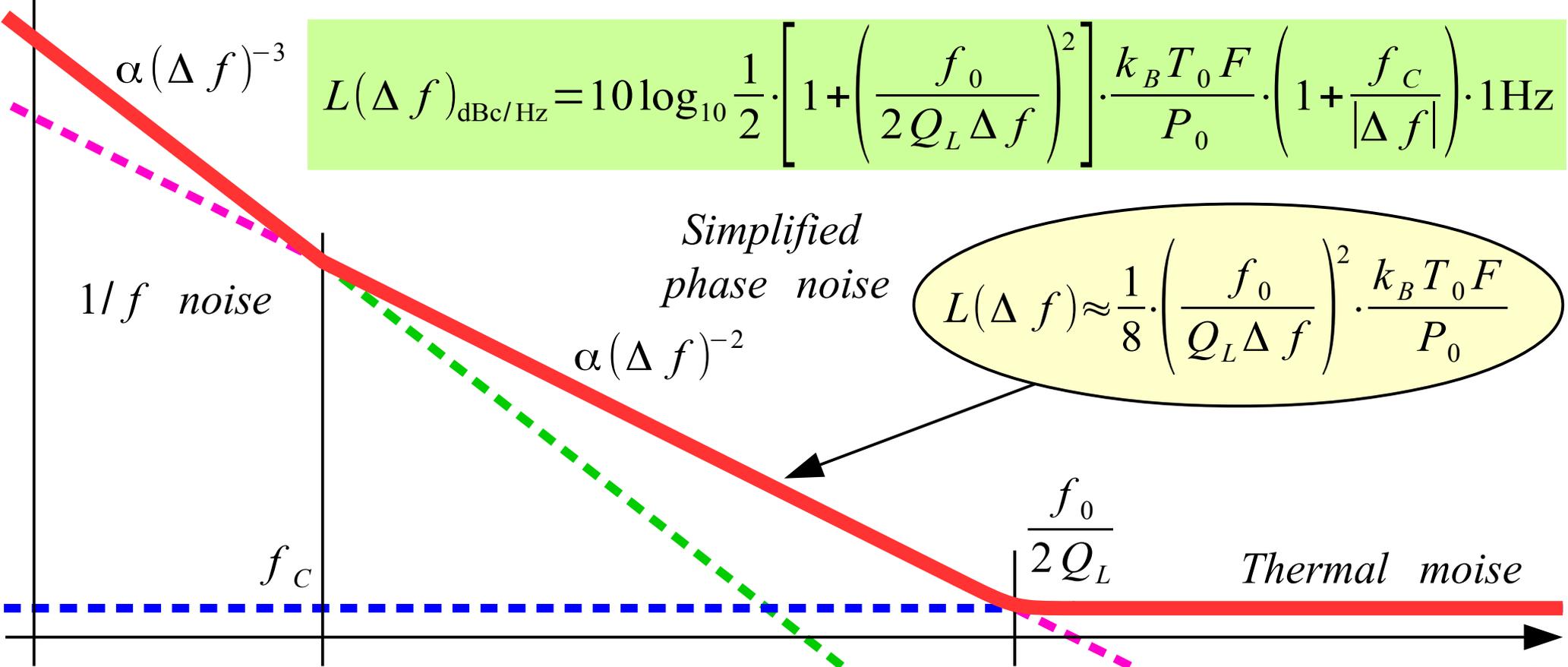
$$L(\Delta f) = \frac{1}{P_0} \cdot \frac{dP_\varphi}{df} = \frac{1}{2} \cdot \left[1 + \left(\frac{f_0}{2Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left(1 + \frac{f_c}{|\Delta f|} \right) \quad [\text{Hz}^{-1}]$$

Valid at
 $L(\Delta f) \cdot \Delta f \ll 1$

Phase noise only

$P_0 \equiv$ carrier power

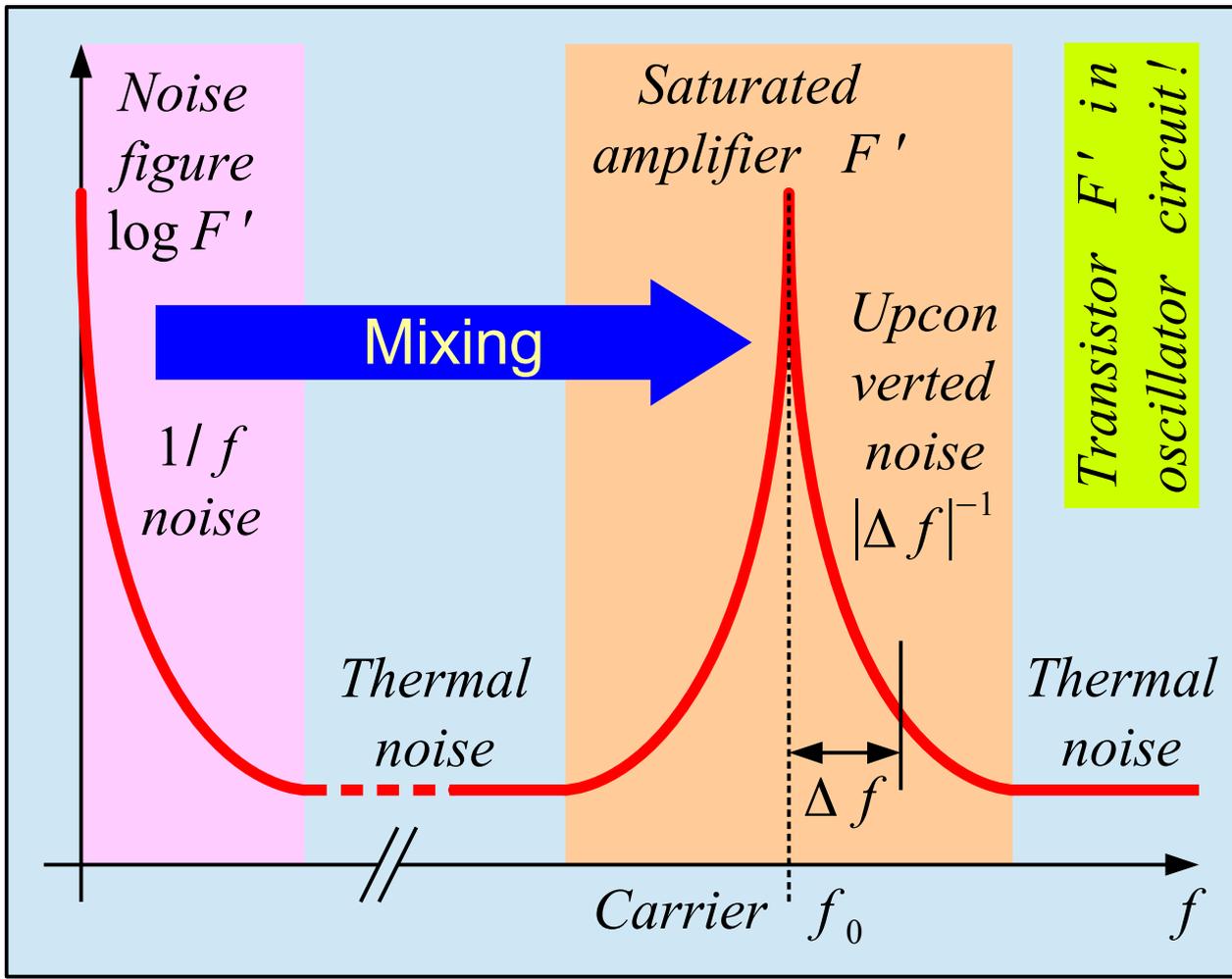
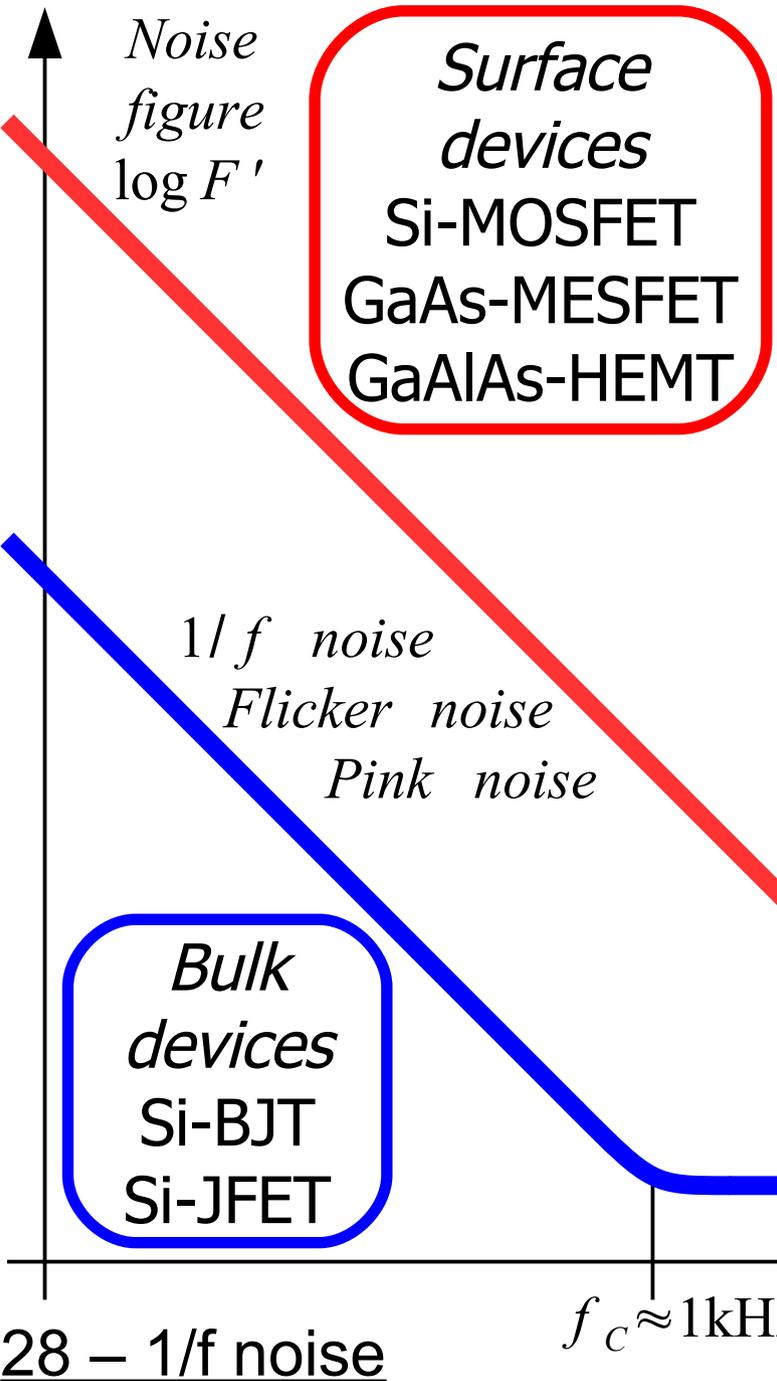
1/f noise



$$L(\Delta f)_{\text{dBc/Hz}} = 10 \log_{10} \frac{1}{2} \cdot \left[1 + \left(\frac{f_0}{2Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left(1 + \frac{f_c}{|\Delta f|} \right) \cdot 1 \text{ Hz}$$

$$L(\Delta f) \approx \frac{1}{8} \cdot \left(\frac{f_0}{Q_L \Delta f} \right)^2 \cdot \frac{k_B T_0 F}{P_0}$$

1/f noise usually does not have a clear explanation!



$$F' = F \left(1 + \frac{f_c}{f} \right) \equiv \text{increased LF noise!}$$

The loaded resonator quality Q_L defines the oscillator phase noise!

$$L(\Delta f) = \frac{1}{2} \cdot \left[1 + \left(\frac{f_0}{2 Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left(1 + \frac{f_c}{|\Delta f|} \right)$$

Variable-frequency oscillators

Q_L

RC VCO

~ 1

BWO tube

~ 1

Varactor-tuned LC VCO

$10 \leftrightarrow 30$

YIG ($Y_3Fe_5O_{12}$) oscillator

$300 \leftrightarrow 1000$

Fixed-frequency oscillators

Q_L

RC multivibrator

~ 1

LC resonator

$30 \leftrightarrow 100$

Cavity resonator

$1000 \leftrightarrow 3000$

Ceramic dielectric resonator

$1000 \leftrightarrow 3000$

AT-cut quartz crystal (fundamental mode)

$3000 \leftrightarrow 10000$

AT-cut quartz crystal (third/fifth overtone)

$10000 \leftrightarrow 30000$

Electro-optical delay line (\$)

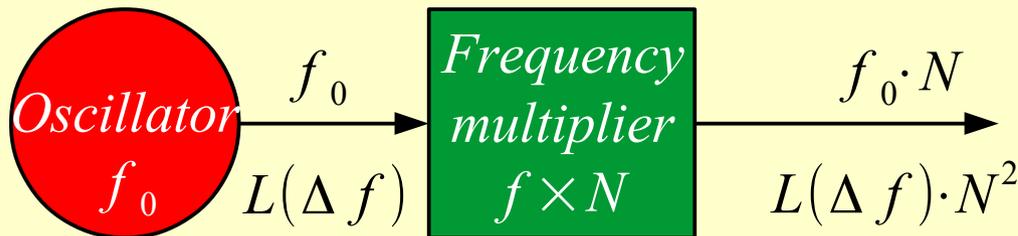
$\sim 10^5$ (noisy!)

Sapphire dielectric resonator (\$\$\$)

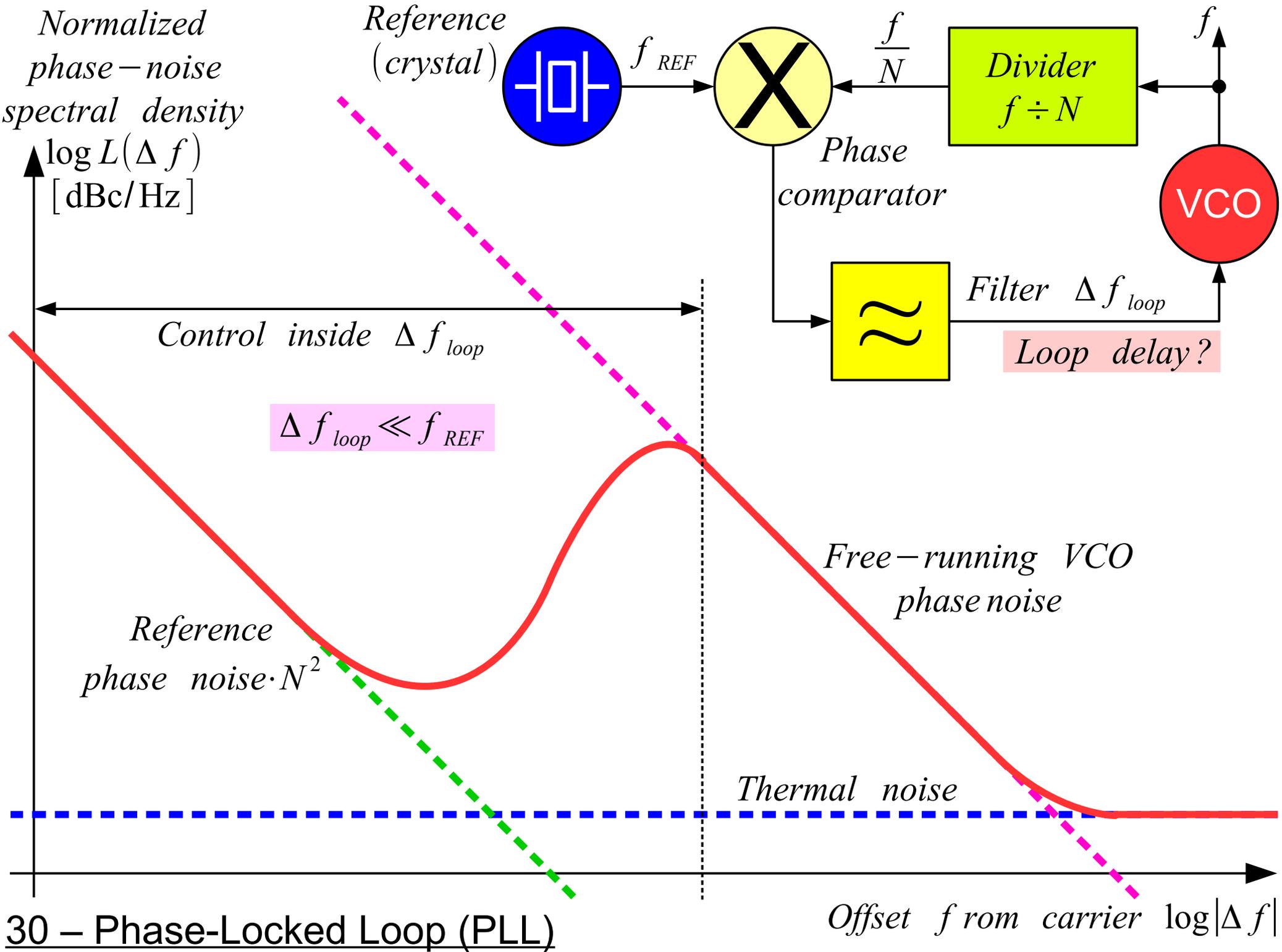
$\sim 3 \cdot 10^5$

Red HeNe LASER

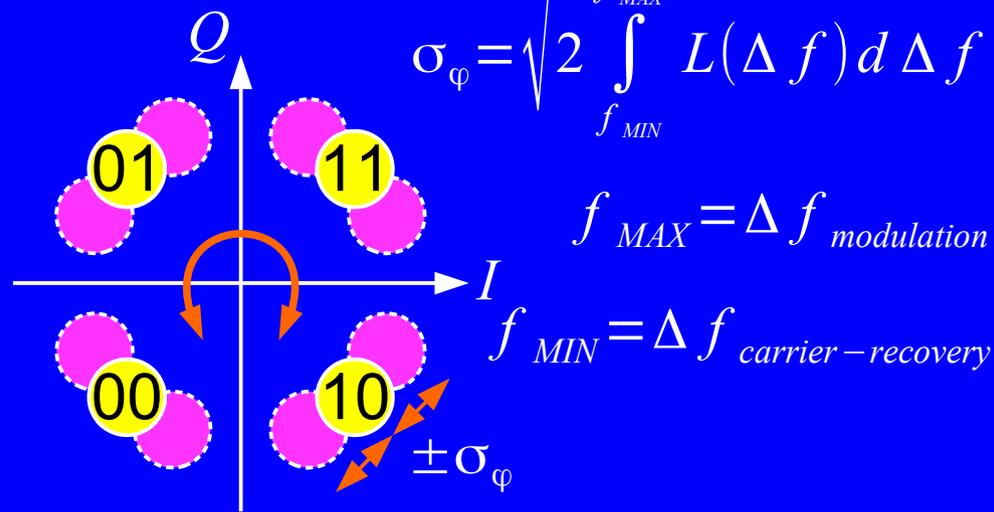
$\sim 10^8$



The phase noise multiplies with the square of the frequency multiplication factor!

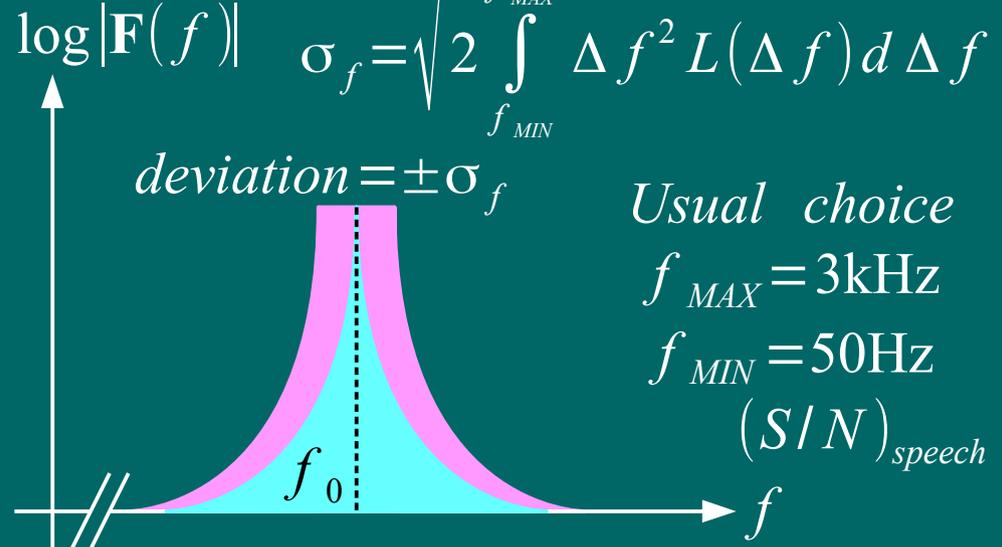


Example QPSK



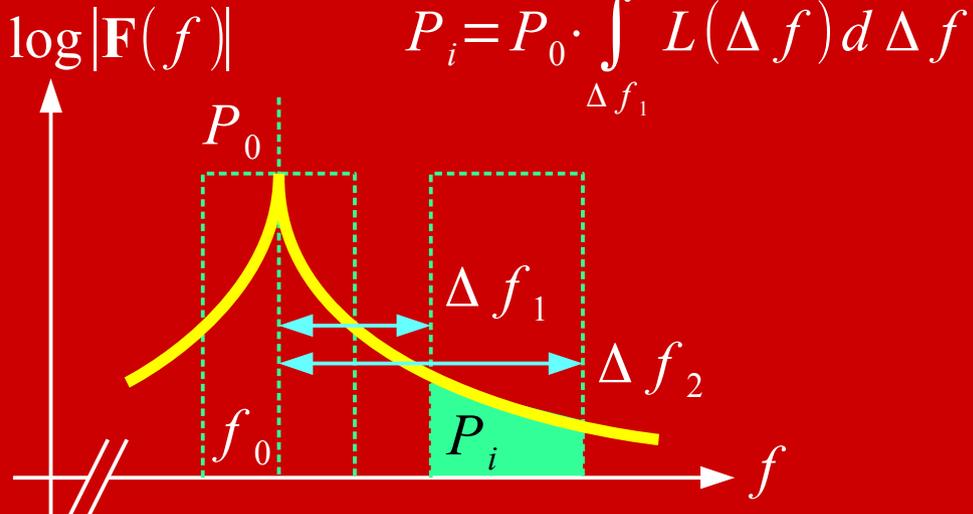
Modulation constellation rotation

Spectrum



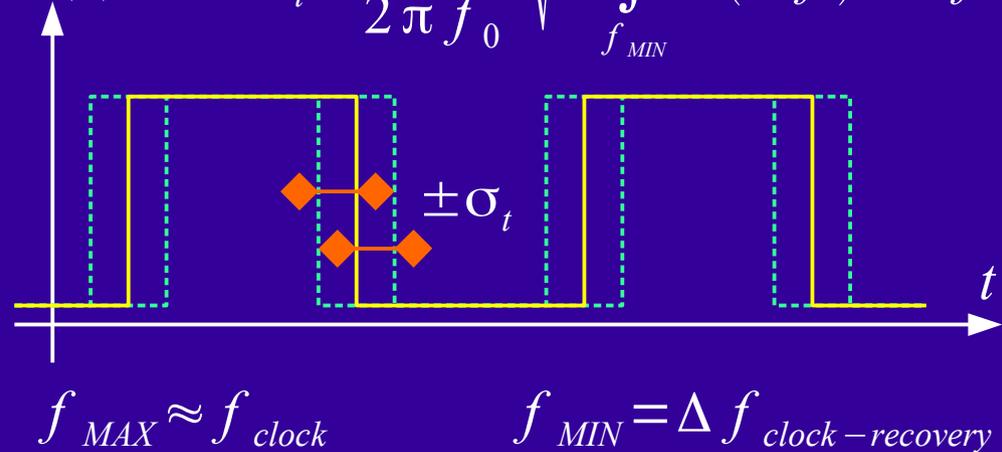
Residual FM

Spectrum

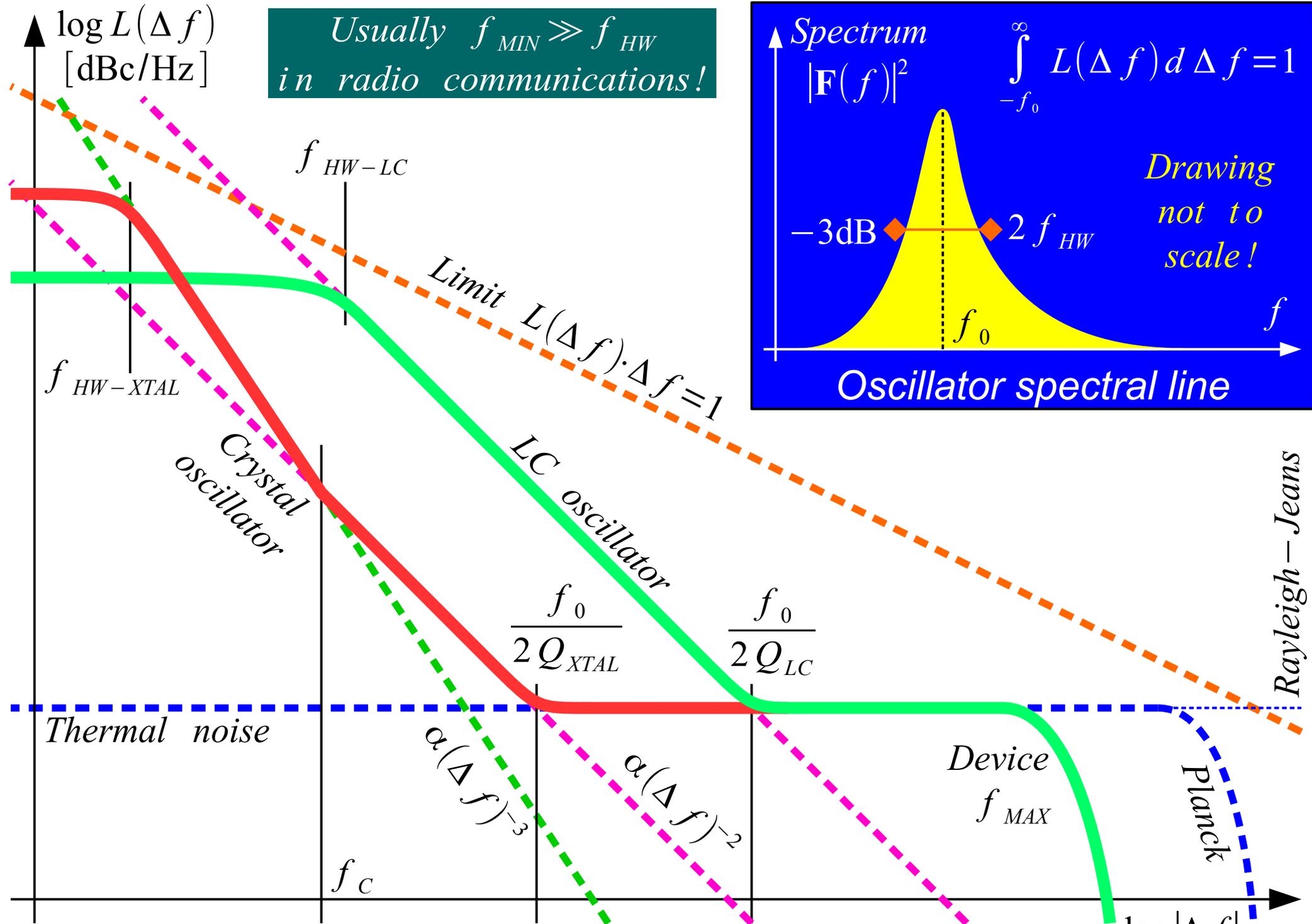


Adjacent-channel interference

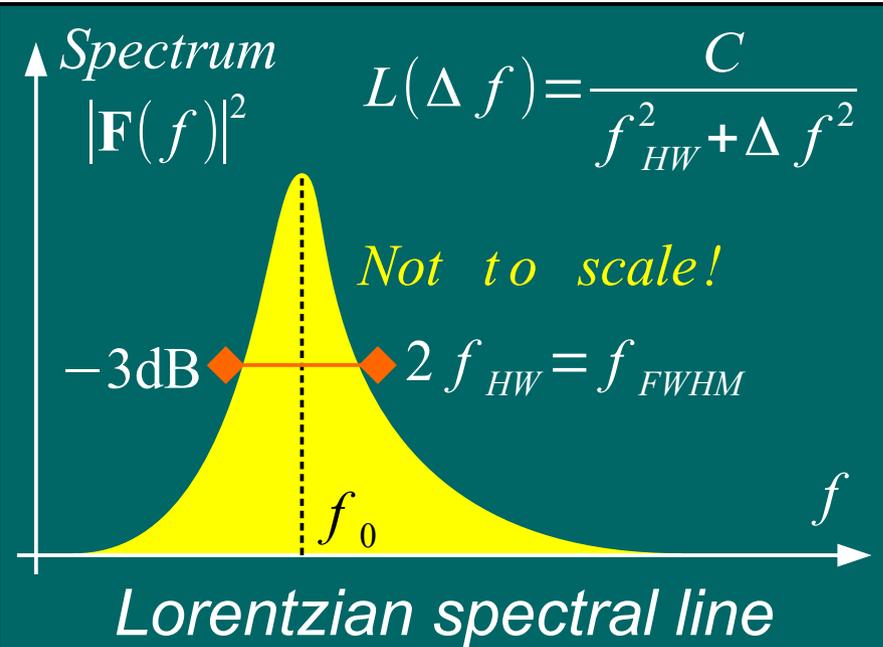
u(t)



Clock jitter



32 – Phase noise without approximations



Flat thermal noise can be neglected: device f_{MAX} or Planck law

LC-oscillator $1/f$ noise can be neglected

$$L(\Delta f) = \frac{1}{8} \cdot \left(\frac{f_0}{Q_L}\right)^2 \cdot \frac{1}{f_{HW}^2 + \Delta f^2} \cdot \frac{k_B T_0 F}{P_0}$$

Lorentzian line in Leeson's equation

$$\int_{-f_0}^{\infty} L(\Delta f) d\Delta f = 1 \approx \int_{-\infty}^{\infty} L(\Delta f) d\Delta f = \frac{1}{8} \cdot \left(\frac{f_0}{Q_L}\right)^2 \cdot \frac{k_B T_0 F}{P_0} \int_{-\infty}^{\infty} \frac{1}{f_{HW}^2 + \Delta f^2} d\Delta f =$$

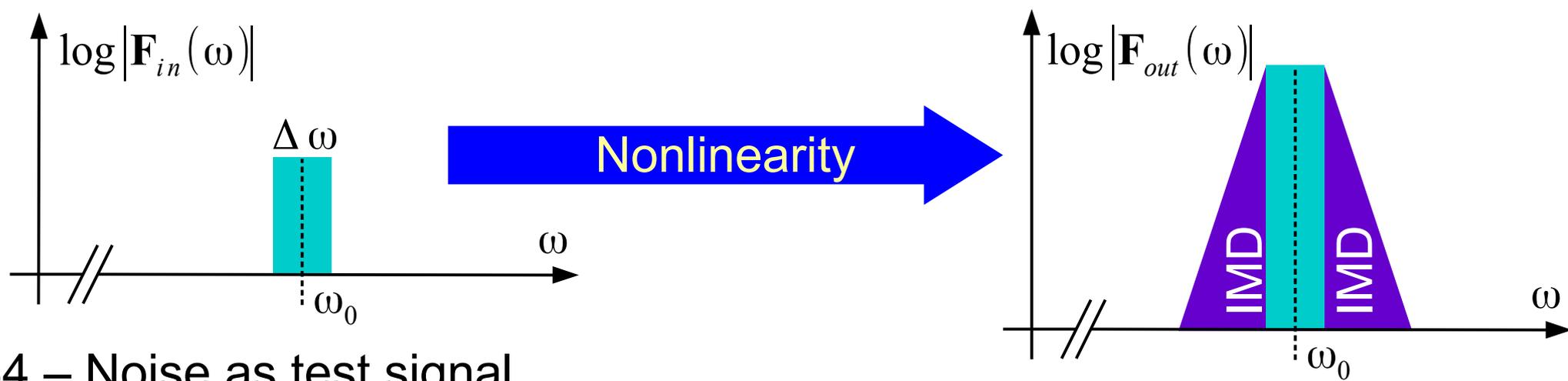
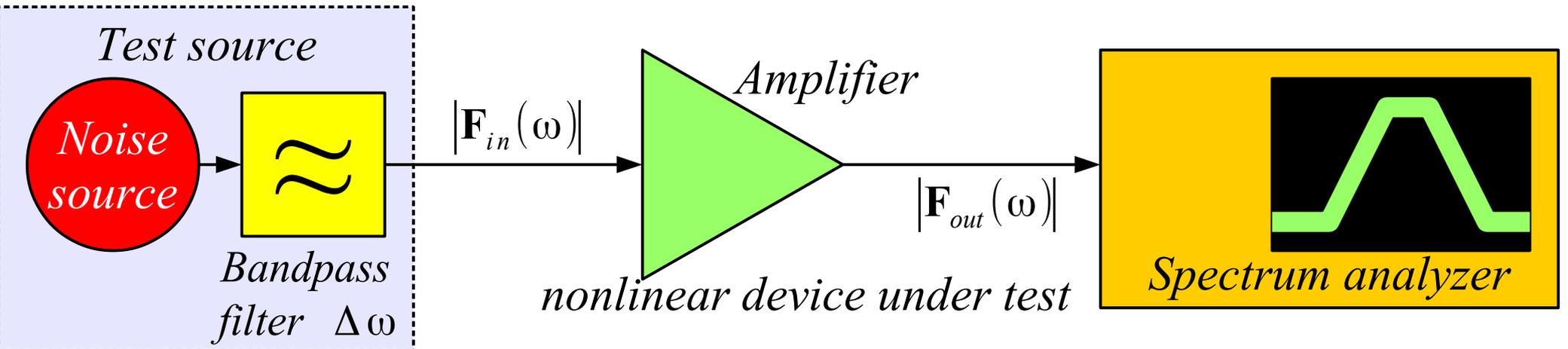
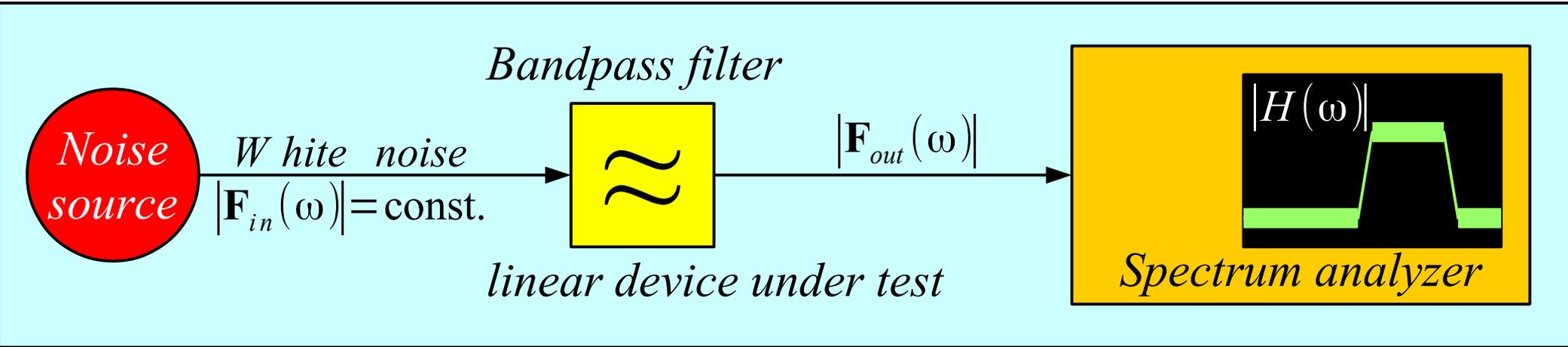
$$= \frac{1}{8} \cdot \left(\frac{f_0}{Q_L}\right)^2 \cdot \frac{k_B T_0 F}{P_0} \cdot \left[\frac{1}{f_{HW}} \cdot \arctan \frac{\Delta f}{f_{HW}} \right]_{\Delta f = -\infty}^{\Delta f = \infty} = \frac{k_B T_0 F}{8 P_0} \cdot \left(\frac{f_0}{Q_L}\right)^2 \cdot \frac{\pi}{f_{HW}}$$

$$f_{HW} = \frac{\pi k_B T_0 F}{8 P_0} \cdot \left(\frac{f_0}{Q_L}\right)^2$$

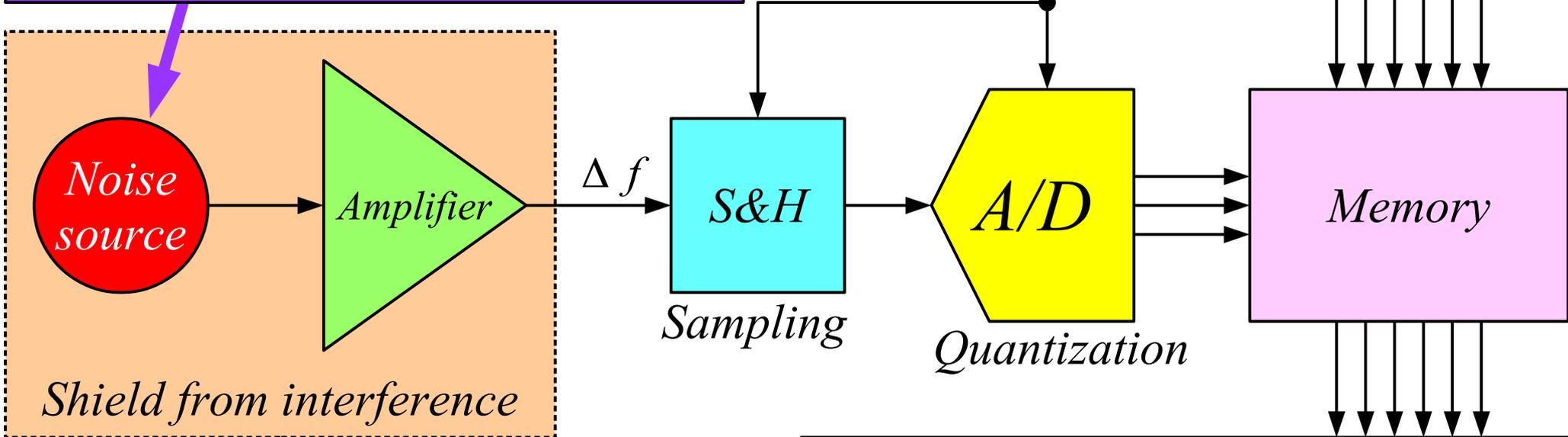
Example $f_0 = 3\text{GHz}$ $Q_L = 10$
 $P_0 = 0.1\text{mW}$ $F = 10\text{dB}$
 $f_{HW} = 14\text{Hz}$ $f_{FWHM} = 28\text{Hz}$

$$C = \frac{k_B T_0 F}{8 P_0} \cdot \left(\frac{f_0}{Q_L}\right)^2 = \frac{f_{HW}}{\pi}$$

$$L(\Delta f) = \frac{f_{HW} / \pi}{f_{HW}^2 + \Delta f^2}$$

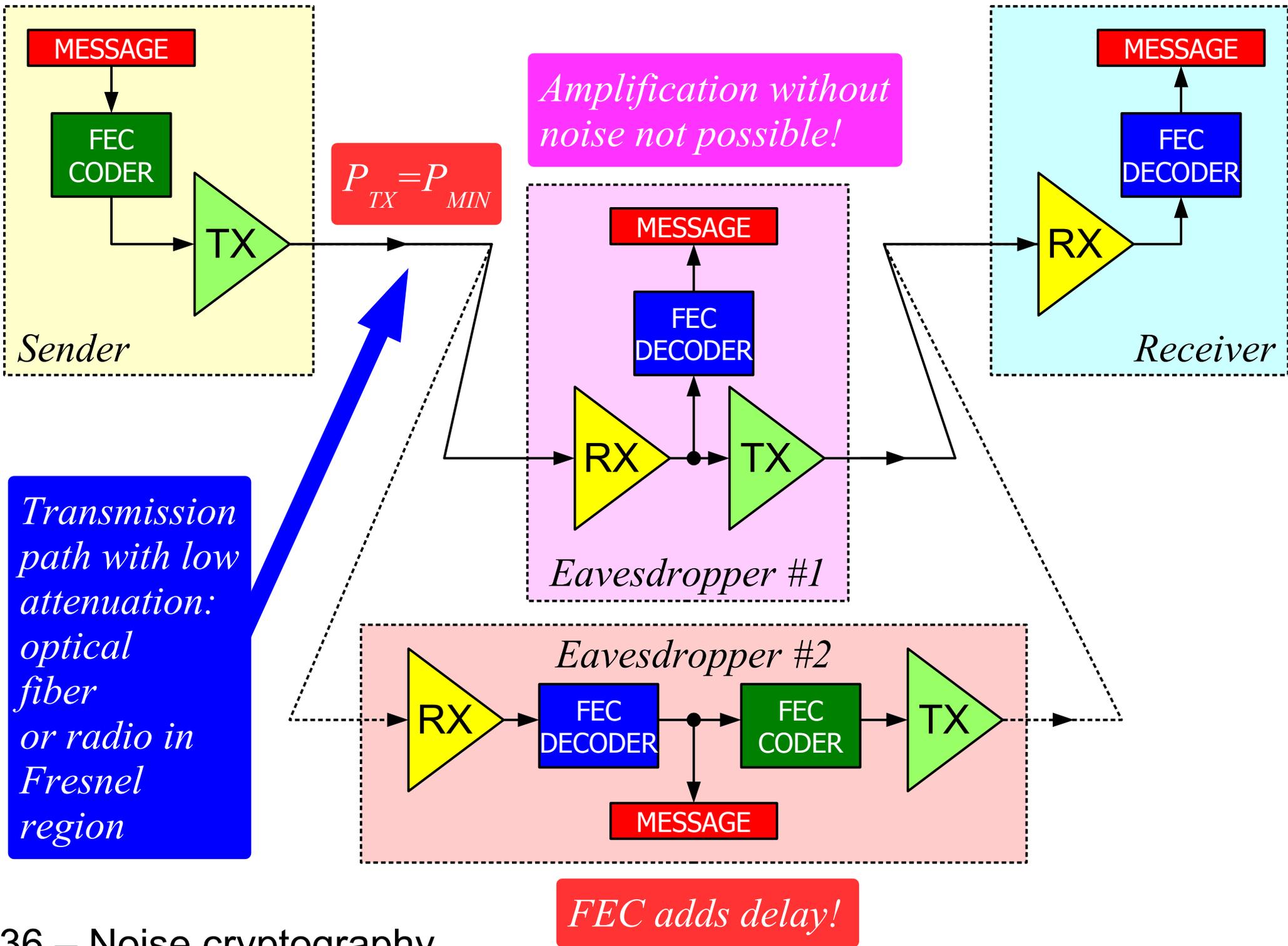


Natural sources of random signals:
Thermal noise
Shot noise
Avalanche breakdown
Radioactive decay
...

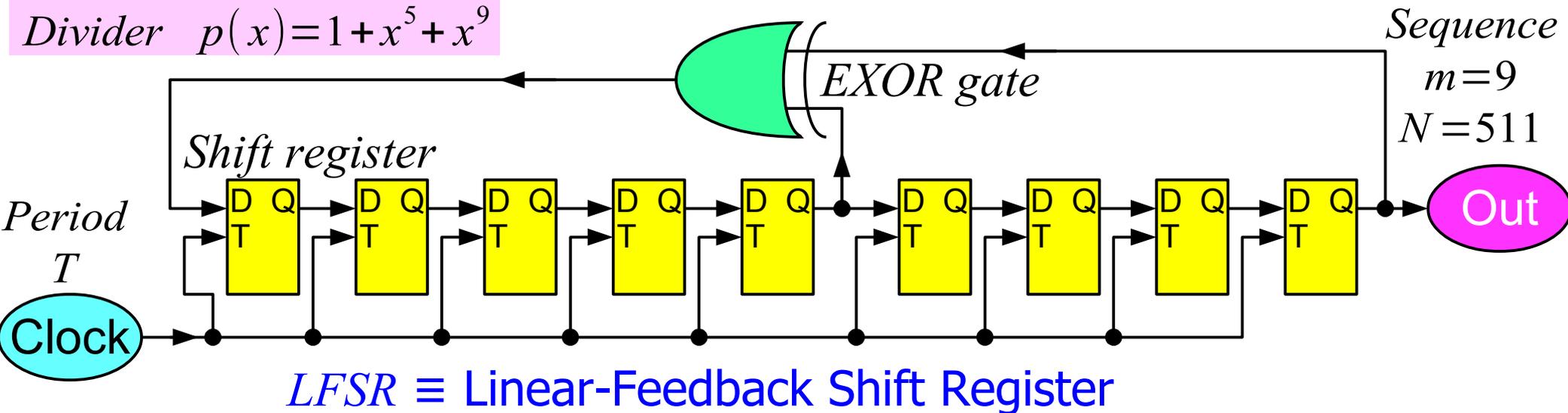


Interference is not random!
Interference might be intentional!

Arbitrary-length cryptographic key:
Password (rather short key...)
Keys for DES, AES etc
One-time pad
(very long but unbreakable key!)



Divider $p(x) = 1 + x^5 + x^9$



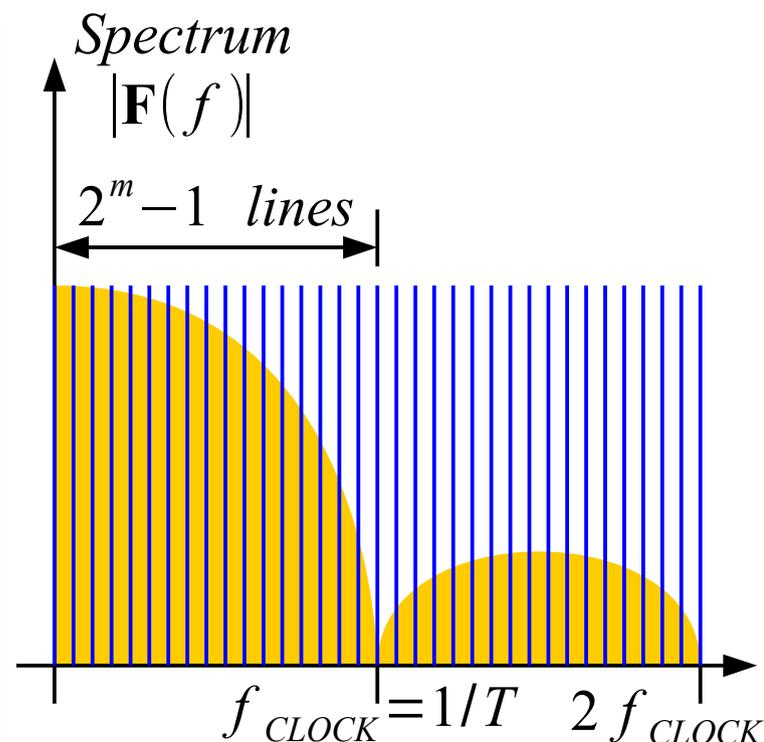
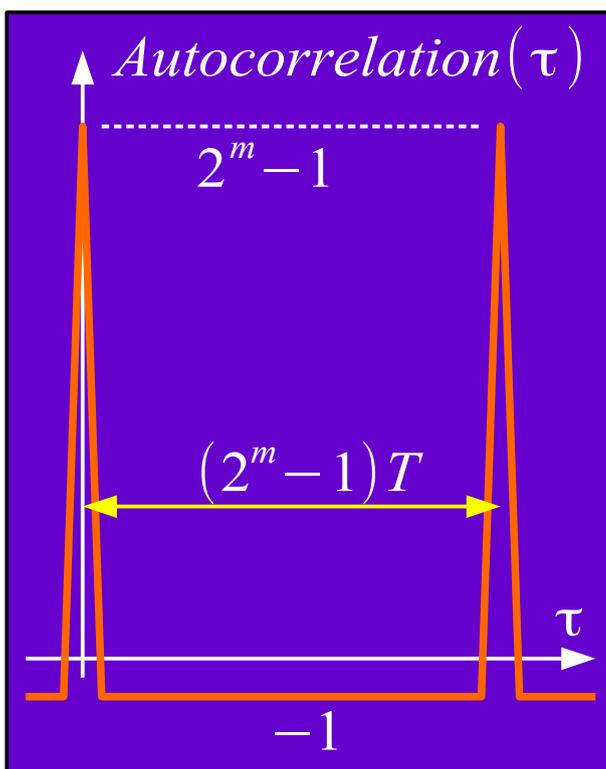
Primitive polynomial $p(x) = 1 + x^l + x^m \rightarrow$ max sequence length $N = 2^m - 1$

2^{m-1} ones and $2^{m-1} - 1$ zeros
arranged in groups of

- 1X m ones, m-1 zeros
- 1X m-2 ones and zeros
- 2X m-3 ones and zeros
- 4X m-4 ones and zeros

.....

- 2^{m-5} groups 111 and 000
- 2^{m-4} groups 11 and 00
- 2^{m-3} individual 1 and 0



Two-valued autocorrelation with a single very pronounced peak:

- synchronization headers for data frames
- spreading sequences in CDMA
- accurate time transfer in radio navigation (GPS, GLONASS)

Perfect frequency spectrum of uniformly-spaced lines and simple generation/checking:

- test sequences for all kinds of telecommunication links
- data scrambling (randomization) as part of line coding

Peak-to-average power ratio:

$$LFSR: \frac{P_{MAX}}{\langle P \rangle} \approx 1 \quad \text{Noise: } \frac{P_{MAX}}{\langle P \rangle} \rightarrow \infty$$

LFSR pseudo-random sequences are of NO cryptographic value: algorithm Berlekamp-Massey 1969

LFSR sequences are the result of pure mathematics that does not appear anywhere else in nature!

*How to present ourselves to the inhabitants of a neighbor galaxy?
How to find out that they look for us?*