

# NOISE IN RADIO/OPTICAL COMMUNICATIONS

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## List of figures: Noise in radio/optical communications

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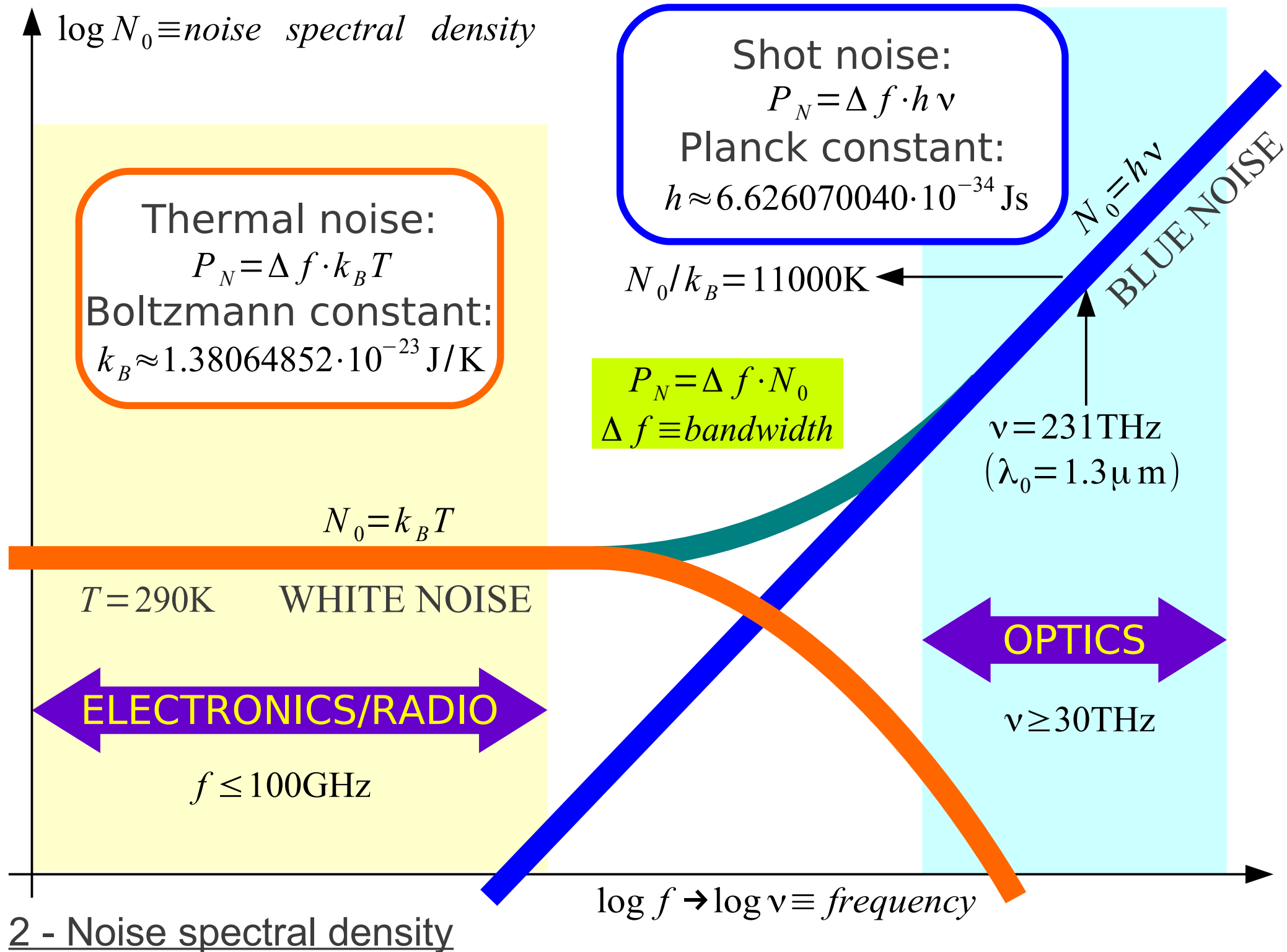
Fifth Solvay International Conference on Electrons and Photons (October 1927). The leading figures Albert Einstein and Niels Bohr disagreed:

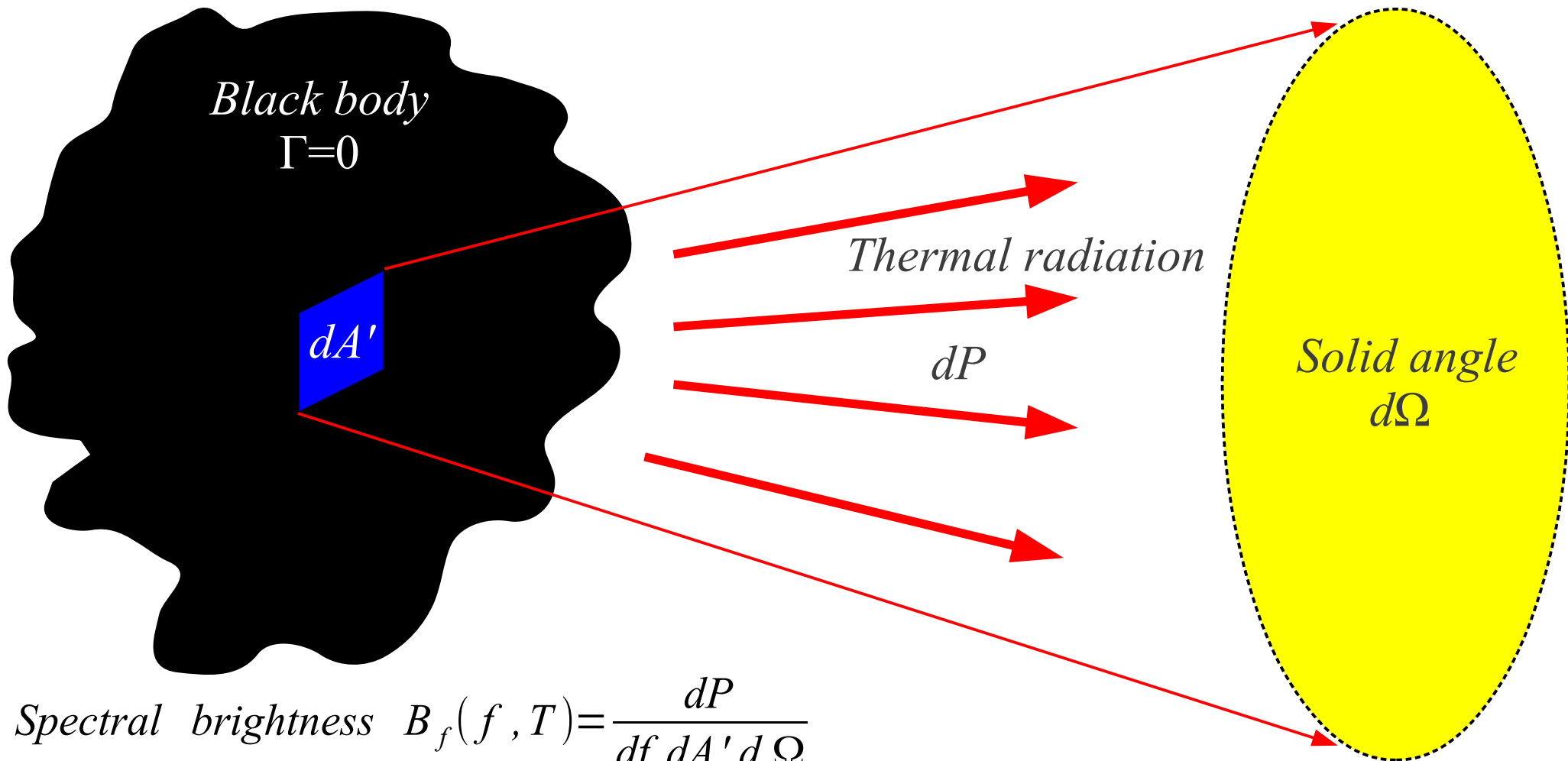
Albert Einstein: „God does not play dice!“

Niels Bohr: „Einstein, stop telling God what to do!“

In telecommunications random signals are called noise. Noise impairs the performance of any communication link.

Noise is a macroscopic description of quantum effects!



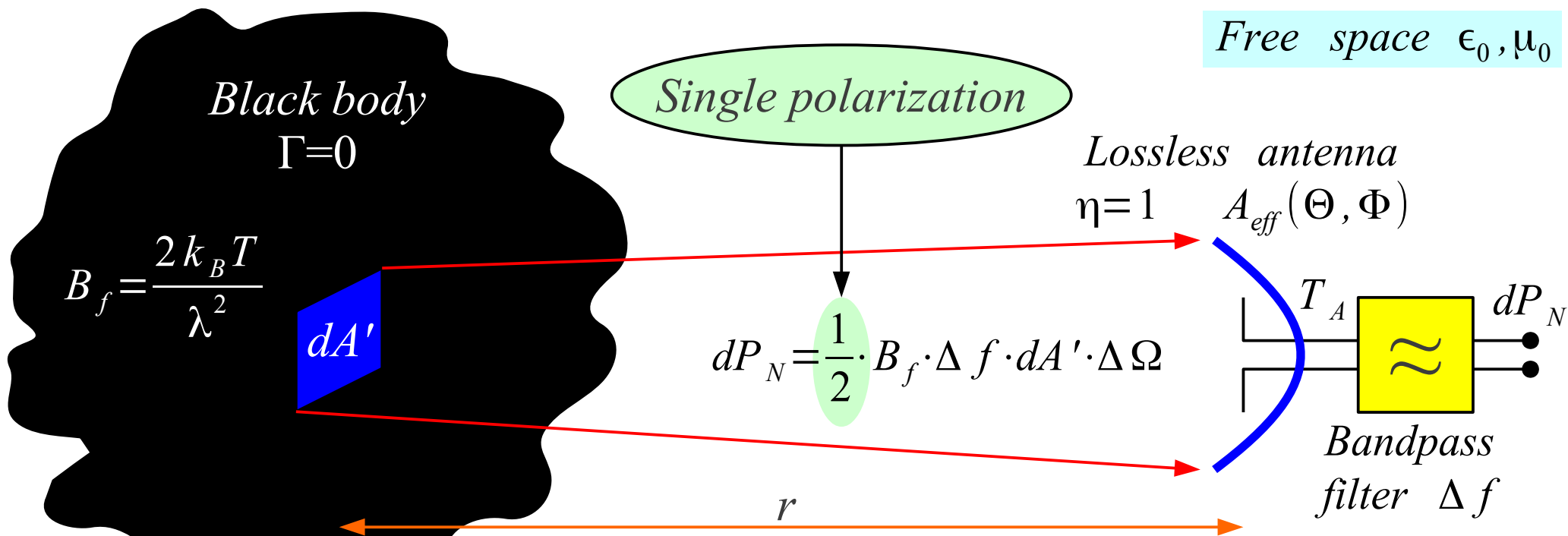


*Planck law*  $B_f(f, T) = \frac{2 h f^3}{c_0^2} \cdot \frac{1}{e^{\frac{h f}{k_B T}} - 1}$

*Free space*  $\epsilon_0, \mu_0$   
 $c_0 = 299792458 \text{ m/s} \approx 3 \cdot 10^8 \text{ m/s}$

*Radio*  $h f \ll k_B T \rightarrow$  *Rayleigh-Jeans approximation*  $B_f(f, T) \approx \frac{2 k_B T f^2}{c_0^2} = \frac{2 k_B T}{\lambda^2}$

### 3 – Black-body thermal radiation



$$\Delta \Omega = \frac{A_{eff}(\Theta, \Phi)}{r^2} = \frac{\lambda^2 D(\Theta, \Phi)}{4\pi r^2} = \frac{\lambda^2 |F(\Theta, \Phi)|^2}{r^2 \iint_{4\pi} |F(\Theta^*, \Phi^*)|^2 d\Omega^*}$$

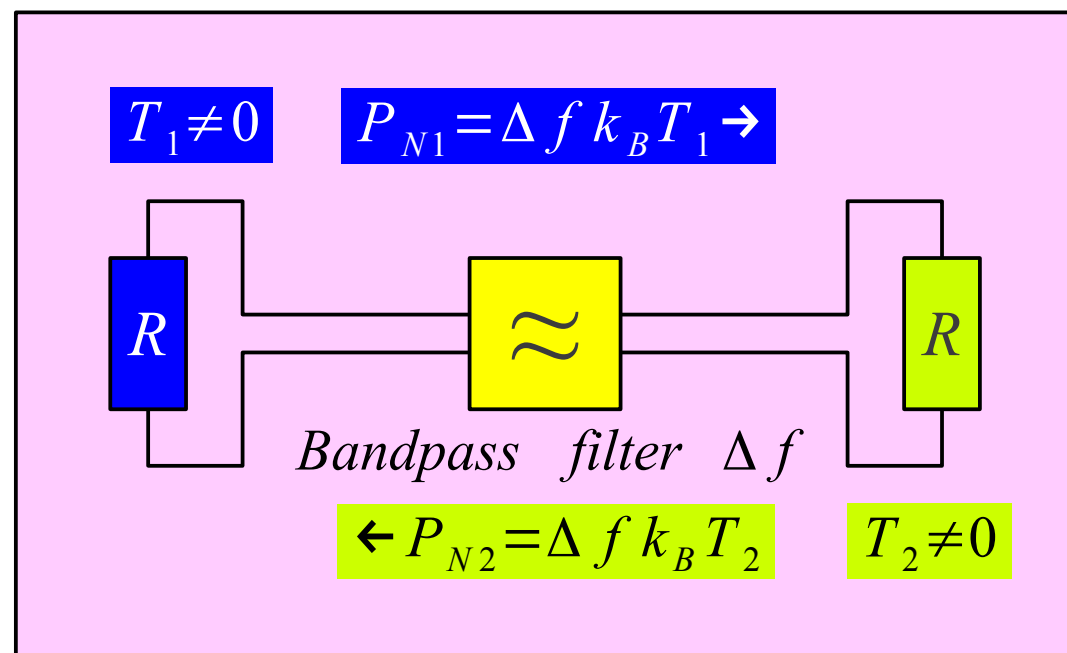
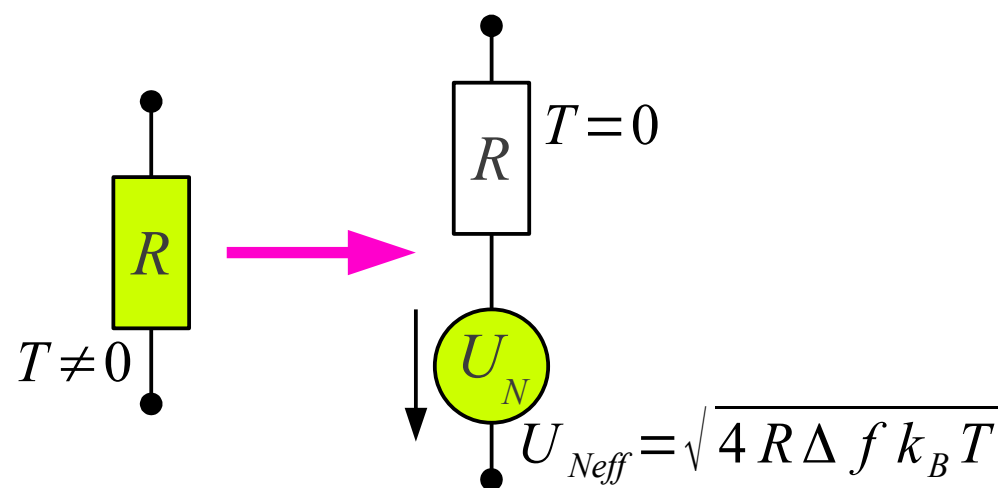
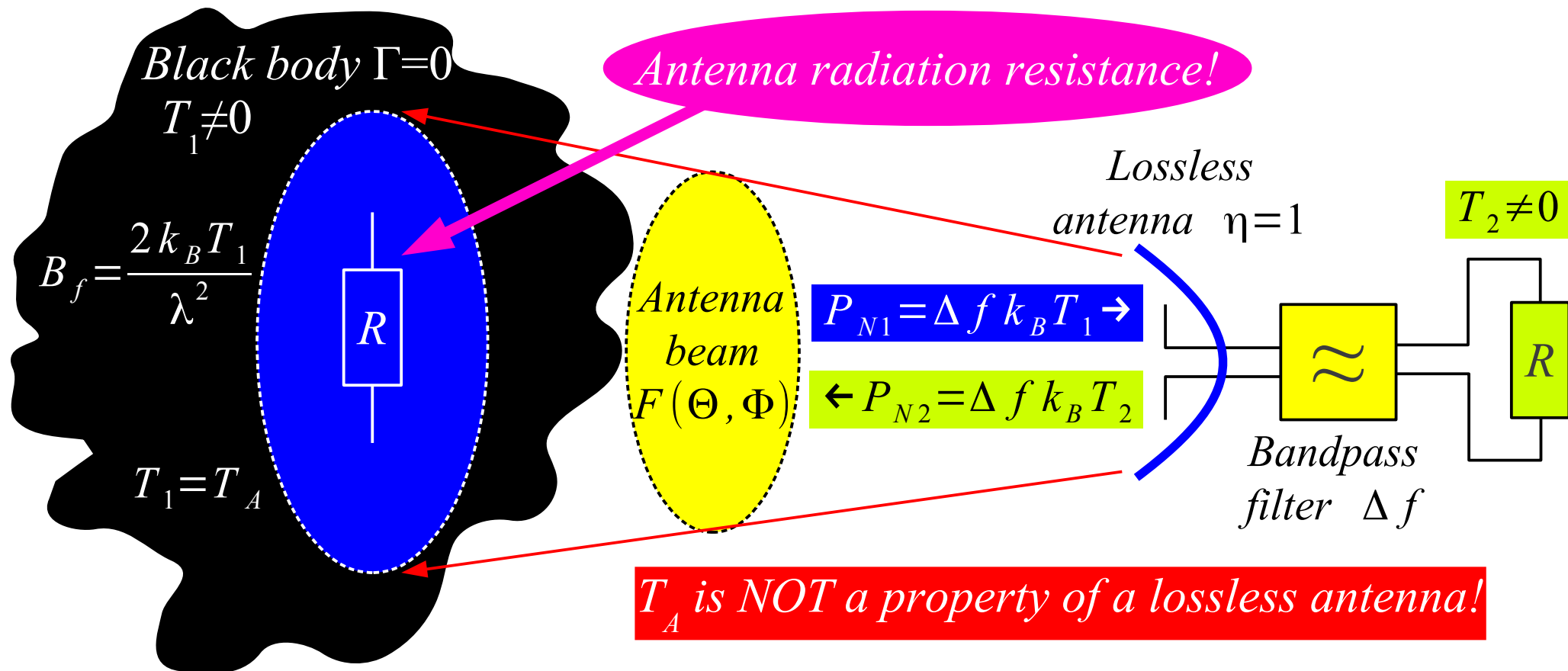
$$dA' = r^2 d\Omega$$

$$P_N = \iint_{A'} \frac{1}{2} \cdot B_f \cdot \Delta f \cdot dA' \cdot \Delta \Omega = \iint_{4\pi} \frac{1}{2} \cdot \frac{2k_B T(\Theta, \Phi)}{\lambda^2} \cdot \Delta f \cdot r^2 d\Omega \cdot \frac{\lambda^2 |F(\Theta, \Phi)|^2}{r^2 \iint_{4\pi} |F(\Theta^*, \Phi^*)|^2 d\Omega^*}$$

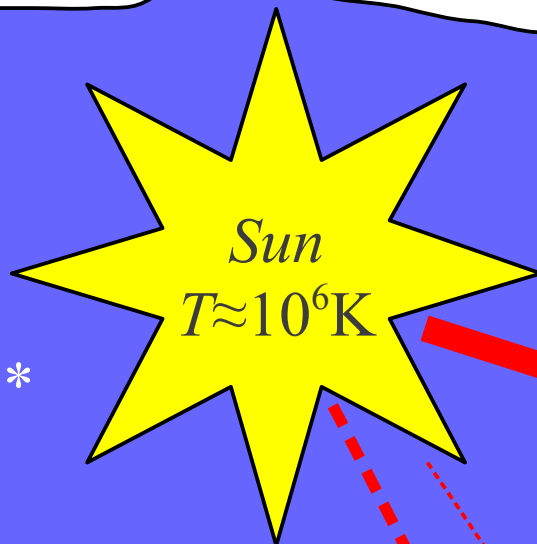
$$P_N = \Delta f k_B \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega} = \Delta f k_B T_A$$

$$T_A = \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

4 – Received thermal-noise power

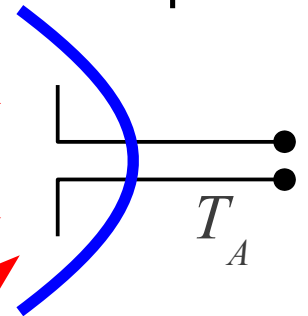


*Cold sky*  
 $T \approx 10\text{K}$   $\Gamma = 0$



$$T_A = \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

*Lossless*  
*antenna*  $\eta = 1$

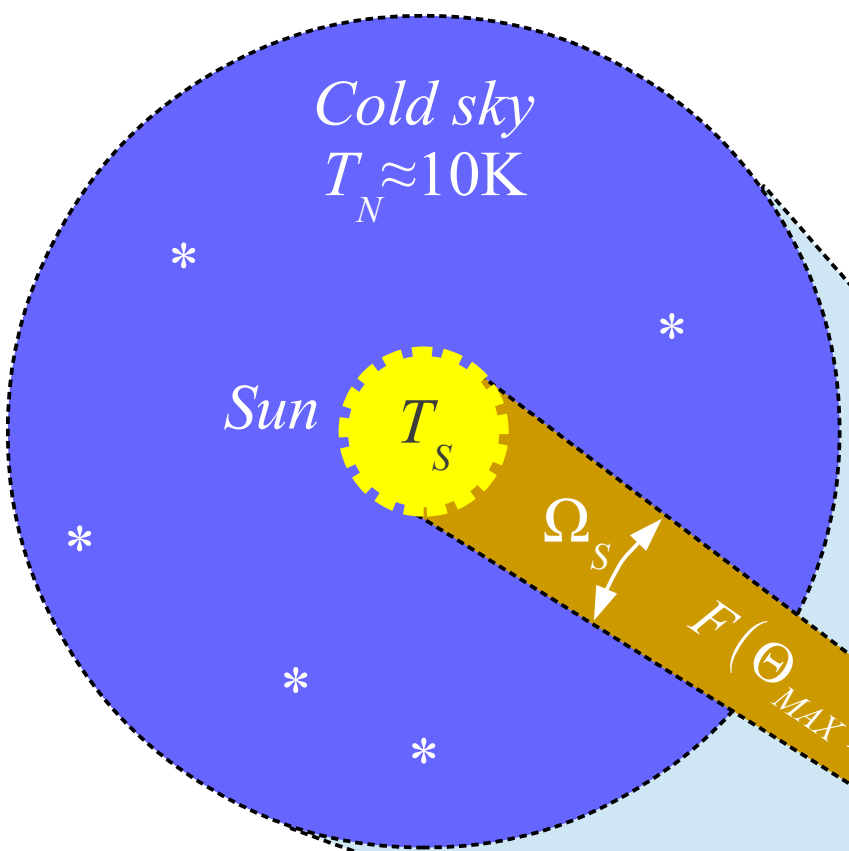


*Vegetation*  
 $T \approx 290\text{K}$   
 $\Gamma \approx 0$

*Ground soil*  
 $T \approx 290\text{K}$   $\Gamma \neq 0$

*Lake*  $|\Gamma| \approx 1 \rightarrow \text{Mirror!}$





$$T_A = \frac{T_S \iint_{\Omega_S} |F(\Theta, \Phi)|^2 d\Omega + T_N \iint_{4\pi - \Omega_S} |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

$$D = \frac{4\pi |F(\Theta_{MAX}, \Phi_{MAX})|^2}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

$$T_A \approx \frac{T_S \Omega_S D}{4\pi} + T_N$$

$$T_A \approx \frac{10^6 \text{ K} \cdot 6 \cdot 10^{-5} \text{ srd} \cdot 100}{4\pi \text{ srd}} + 10 \text{ K}$$

$$T_A \approx 476 \text{ K} + 10 \text{ K} = 486 \text{ K}$$

$$\alpha_S \approx 0.5^\circ \approx 9 \text{ mrd}$$

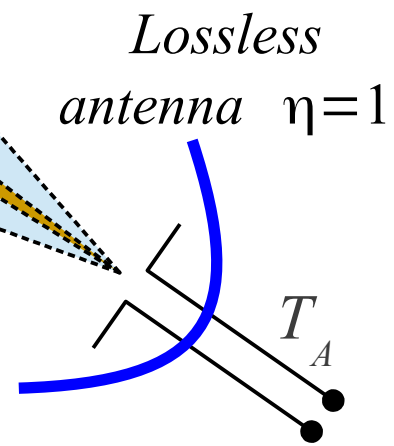
$$T_S \approx 10^6 \text{ K} @ f = 2 \text{ GHz}$$

$$\Omega_S = 2\pi [1 - \cos(\alpha_S/2)] \approx \pi \alpha_S^2 / 4 \approx 6 \cdot 10^{-5} \text{ srd}$$

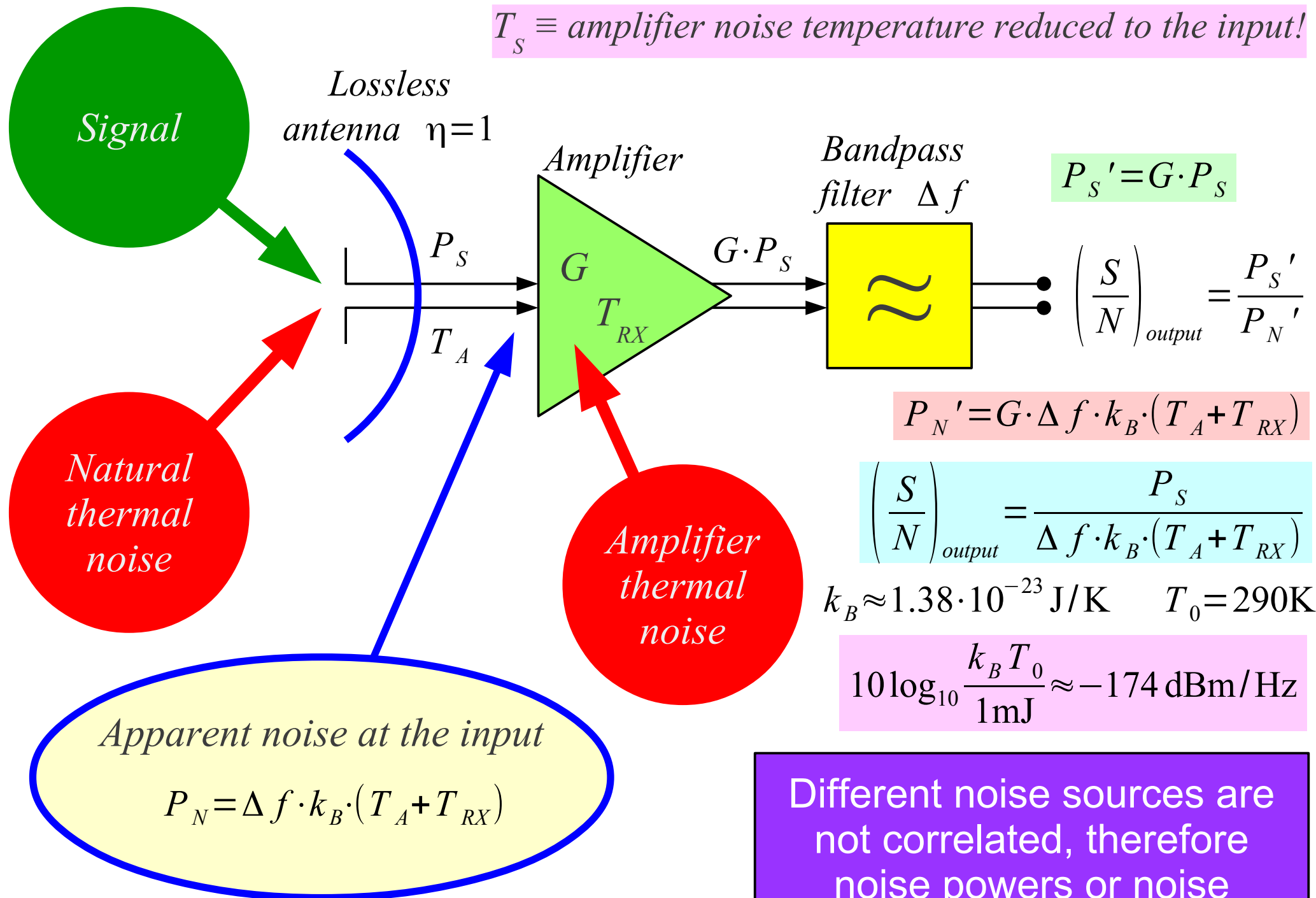
Example:  
D = 20dBi = 100

$$\Omega_A \approx \frac{4\pi}{D} = 0.126 \text{ srd} \gg \Omega_S$$

$F(\Theta, \Phi) = 0$   
outside beam



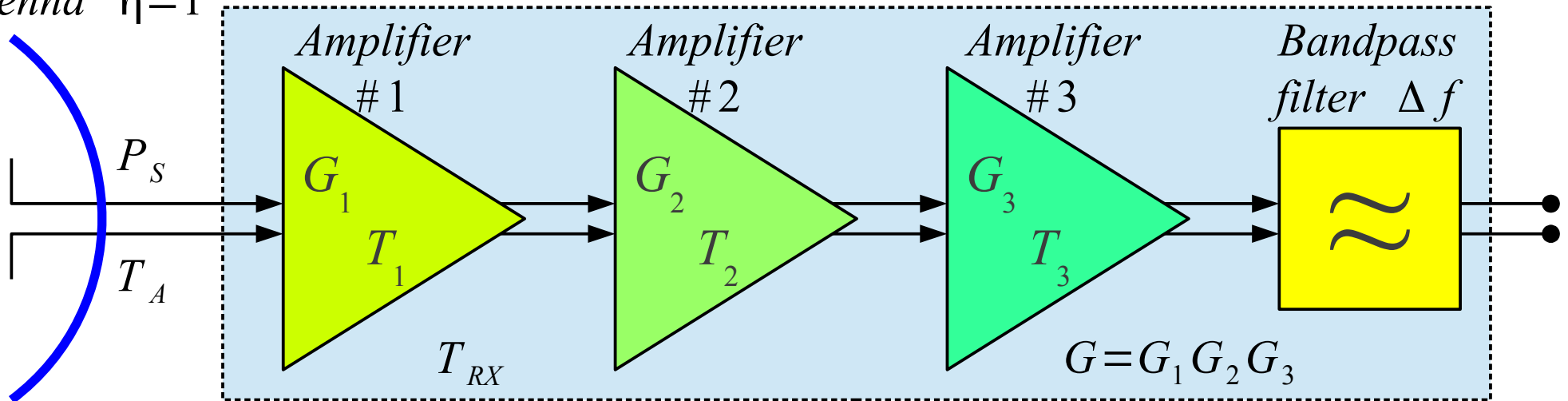
$T_s \equiv$  amplifier noise temperature reduced to the input!



Lossless

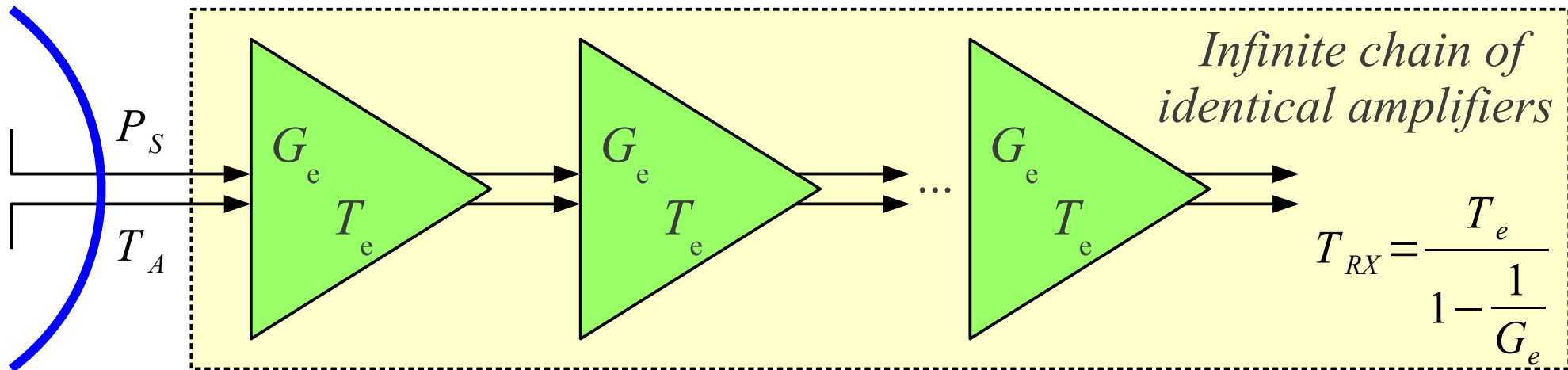
antenna  $\eta=1$

$$P_s' = G_3 G_2 G_1 P_s$$

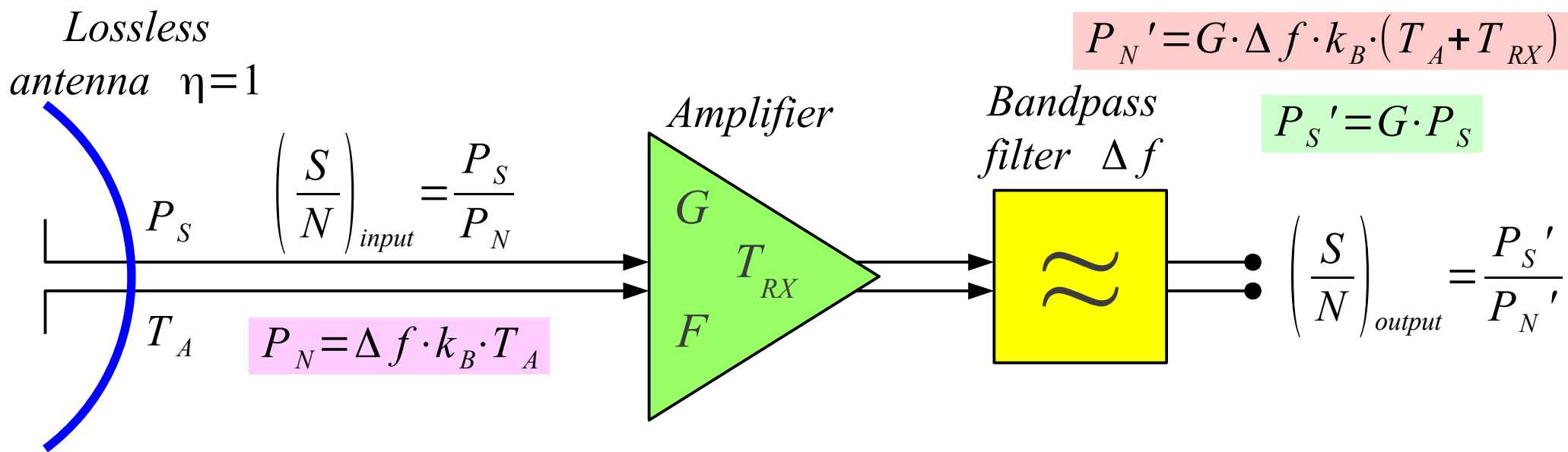


$$P_N' = \Delta f k_B \left[ G_3 G_2 G_1 (T_A + T_1) + G_3 G_2 T_2 + G_3 T_3 \right]$$

$$P_N' = G_3 G_2 G_1 \Delta f k_B (T_A + T_{RX}) \quad \Rightarrow \quad T_{RX} = T_1 + \frac{T_2}{G_1} + \frac{T_3}{G_1 G_2} + \dots$$



$$T_{RX} = \frac{T_e}{1 - \frac{1}{G_e}}$$



*Nonsense definition of the noise figure:*

$$F = \frac{\left(\frac{S}{N}\right)_{input}}{\left(\frac{S}{N}\right)_{output}} = \frac{\frac{P_S}{\Delta f k_B T_A}}{\frac{G P_S}{G \Delta f k_B (T_A + T_{RX})}} = \frac{T_A + T_{RX}}{T_A} = 1 + \frac{T_{RX}}{T_A}$$

*A property of an amplifier can not be a function of  $T_A$ !*

*Sensible definition*

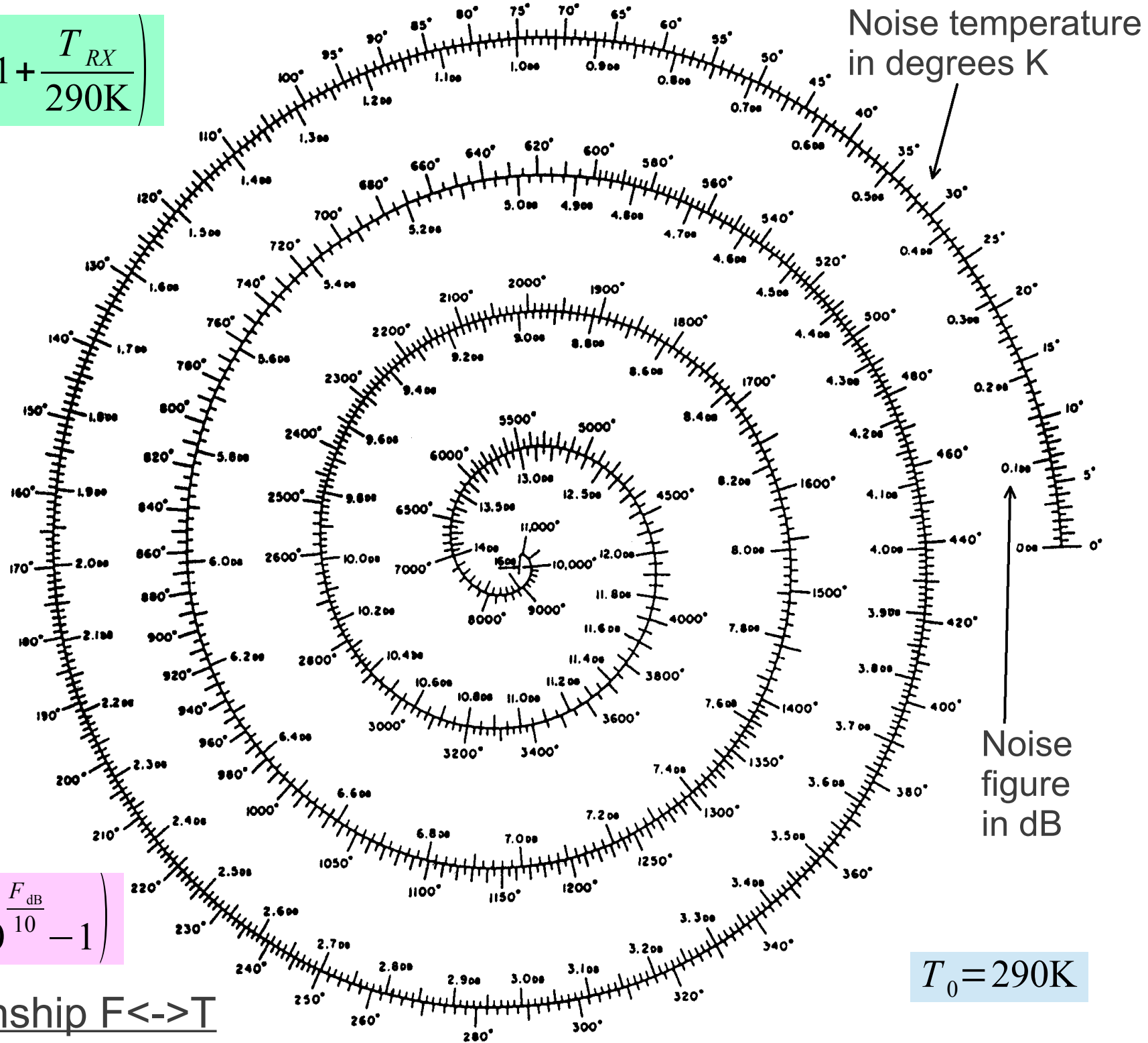
$$F = 1 + \frac{T_{RX}}{T_0} \quad @ \quad T_0 = 290K \quad \leftrightarrow \quad T_{RX} = T_0 (F - 1)$$

*Logarithmic units*

$$F_{dB} = 10 \log_{10} F = 10 \log_{10} \left( 1 + \frac{T_{RX}}{T_0} \right) \quad \leftrightarrow \quad T_{RX} = T_0 \left( 10^{\frac{F_{dB}}{10}} - 1 \right)$$

## 10 – Amplifier noise figure

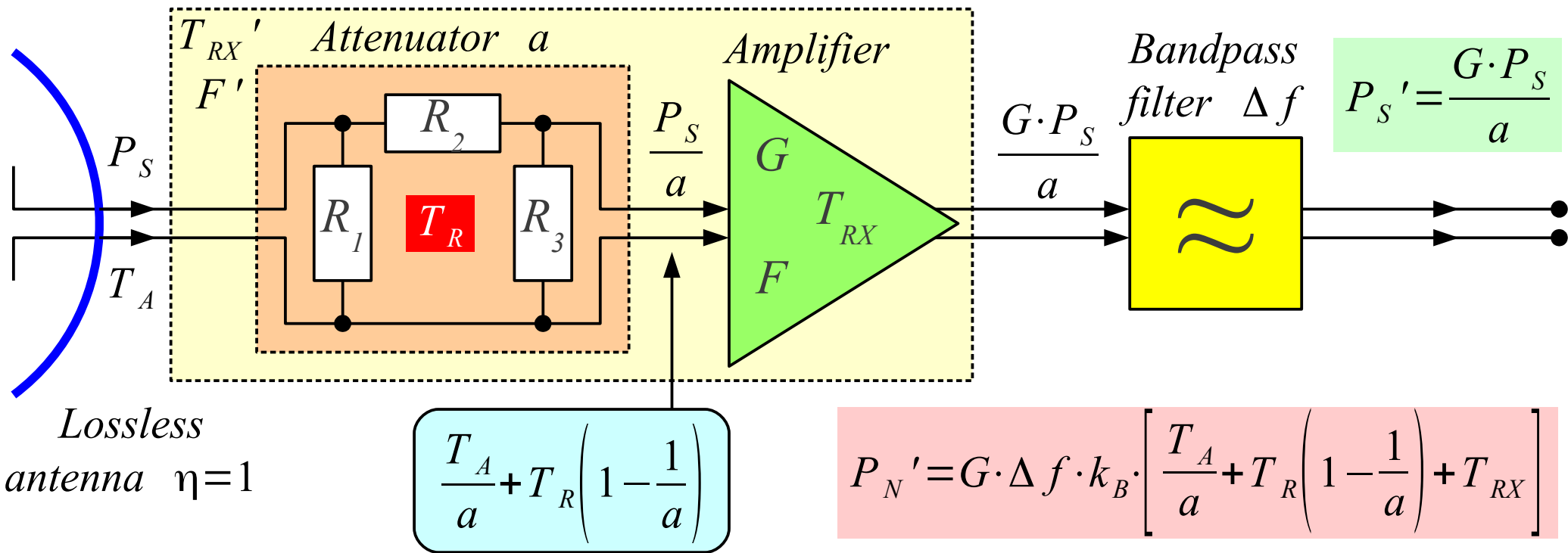
$$F_{\text{dB}} = 10 \log_{10} \left( 1 + \frac{T_{\text{RX}}}{290\text{K}} \right)$$



$$T_{\text{RX}} = 290\text{K} \left( 10^{\frac{F_{\text{dB}}}{10}} - 1 \right)$$

$$T_0 = 290\text{K}$$

11 – Relationship F<->T



$$\left( \frac{S}{N} \right)_{output} = \frac{P_S'}{P_N'} = \frac{P_S}{\Delta f \cdot k_B \cdot [T_A + T_R(a-1) + a T_{RX}]}$$

$$F' = 1 + \frac{T_{RX}'}{T_0} = 1 + \frac{T_R}{T_0}(a-1) + a \frac{T_{RX}}{T_0}$$

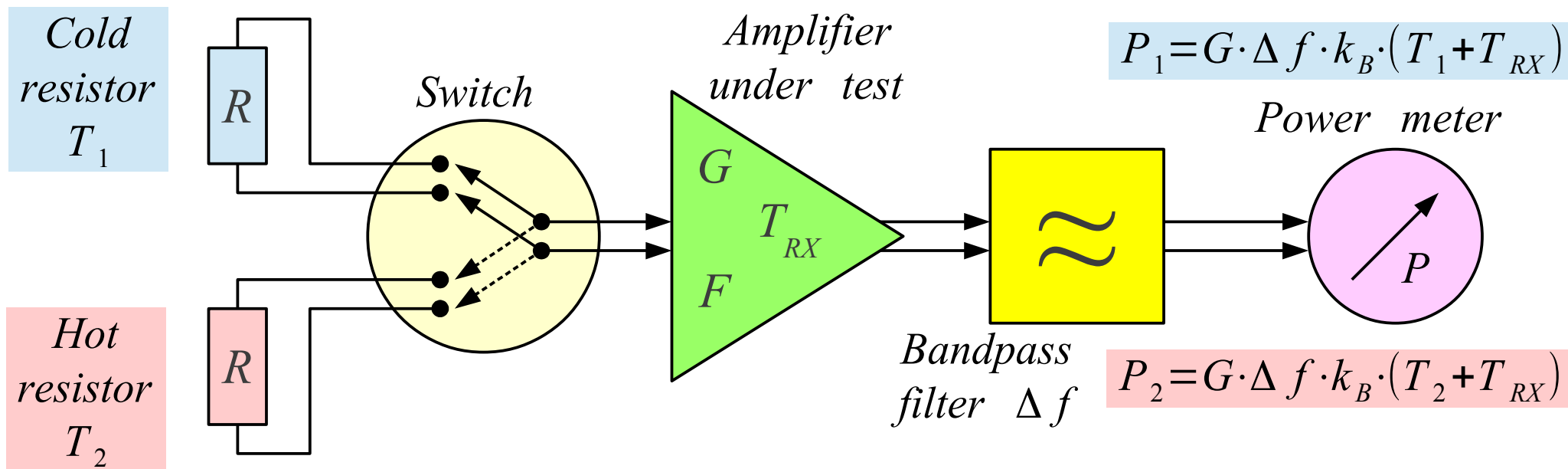
Frequent case  $T_R \approx T_0 = 290K$

$$F' \approx a + a \frac{T_{RX}}{T_0} = a \left( 1 + \frac{T_{RX}}{T_0} \right) = a \cdot F$$

$$F_{dB}' \approx a_{dB} + F_{dB}$$

- Attenuator examples  $T_R \approx T_0 = 290K$
- $F' \approx a \cdot F$        $F_{dB}' \approx a_{dB} + F_{dB}$
- (1) lossy antenna  $a_{dB} = -10 \log_{10} \eta$
  - (2) lossy transmission line  $a_{dB}$
  - (3) lossy bandpass filter  $a_{dB}$
  - (4) passive-mixer loss  $a_{dB}$

Amplifier device	Gain $G$ [dB]	Noise temperature $T_{RX}$ [K]	Noise figure $F_{dB}$ [dB]
Vacuum tube with control grid (triode, pentode)	10↔20	1600↔9000	8↔15
Vacuum tube with speed modulation (klystron, TWT)	20↔50	3000↔30000	10↔20
Parametric amplifier (room temperature)	10↔15	75↔300	1↔3
Si BJT, JFET or MOSFET (room temperature)	10↔20	75↔300	1↔3
GaAs FET or HEMT (room temperature)	10↔15	20↔120	0.3↔1.5
GaAs FET or HEMT (liquid-nitrogen 77K)	10↔15	7↔35	0.1↔0.5
Si or GaAs MMIC amplifier	10↔25	170↔1600	2↔8
Operational amplifier	40↔100	$10^4$ ↔ $10^9$	16↔66



The unknowns  $G \cdot \Delta f \cdot k_B$   
cancel in the  $Y$  ratio!

$$Y = \frac{P_2}{P_1} = \frac{T_2 + T_{RX}}{T_1 + T_{RX}}$$

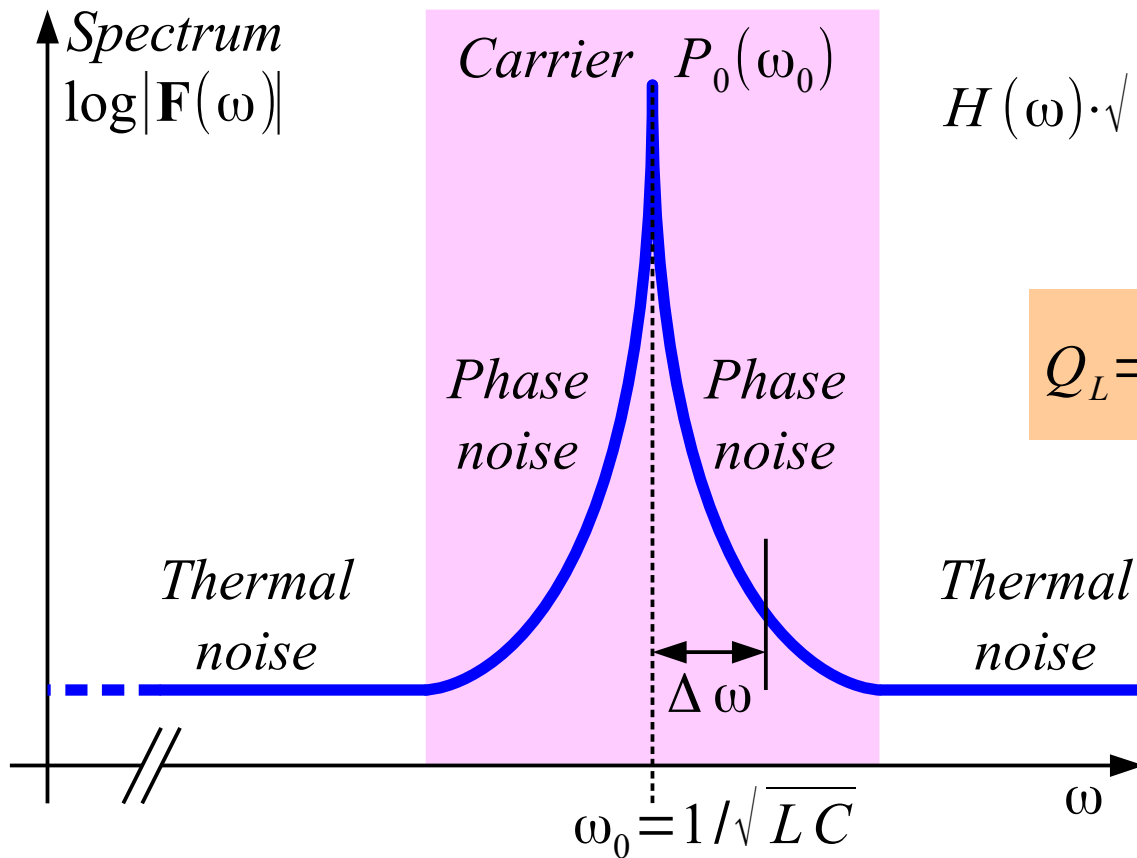
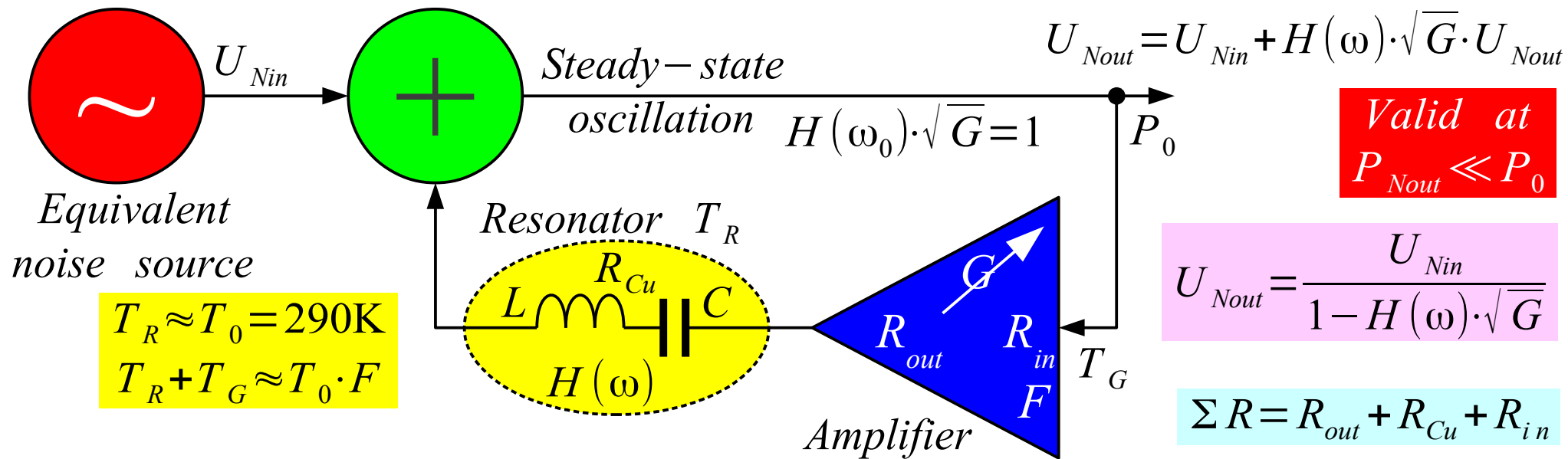
$$T_{RX} = \frac{T_2 - Y \cdot T_1}{Y - 1}$$

$$T_0 = 290\text{K}$$

$$F_{dB} = 10 \log_{10} \left[ 1 + \frac{T_2 - Y \cdot T_1}{(Y - 1) \cdot T_0} \right]$$

Resistor type	Temperature
Antenna into cold sky	$\sim 20\text{K}$
Liquid $\text{N}_2$ cooled R	$\sim 77\text{K}$
Antenna into absorber	$\sim 290\text{K}$
R at room temperature	$\sim 290\text{K}$
Light-bulb filament as R	$\sim 2000\text{K}$
Ionized gas as R	$\sim 10^4\text{K}$
Avalanche breakdown	$\sim 10^6\text{K}$





$$H(\omega) \cdot \sqrt{G} = \frac{\Sigma R}{\Sigma R + j\omega L + \frac{1}{j\omega C}} \approx \frac{1}{1 + j2Q_L \frac{\Delta\omega}{\omega_0}}$$

$$Q_L = \frac{\omega_0 L}{\Sigma R}$$

$$U_{Nout} \approx U_{Nin} \cdot \left( 1 + \frac{\omega_0}{j2Q_L \Delta\omega} \right)$$

$$P_{Nout} \approx P_{Nin} \cdot \left[ 1 + \left( \frac{\omega_0}{2Q_L \Delta\omega} \right)^2 \right]$$

Amplitude and phase noise

$$P_{Nout} \approx P_{Nin} \cdot \left[ 1 + \left( \frac{f_0}{2Q_L \Delta f} \right)^2 \right]$$

Normalized  
phase-noise

Saturation removes amplitude  
noise  $P_\phi = P_{Nout}/2$

$$\frac{dP_{Nin}}{df} = N_0 = k_B(T_R + T_G) \approx k_B T_0 F$$

spectral density

$\log L(\Delta f)$   
[dBc/Hz]

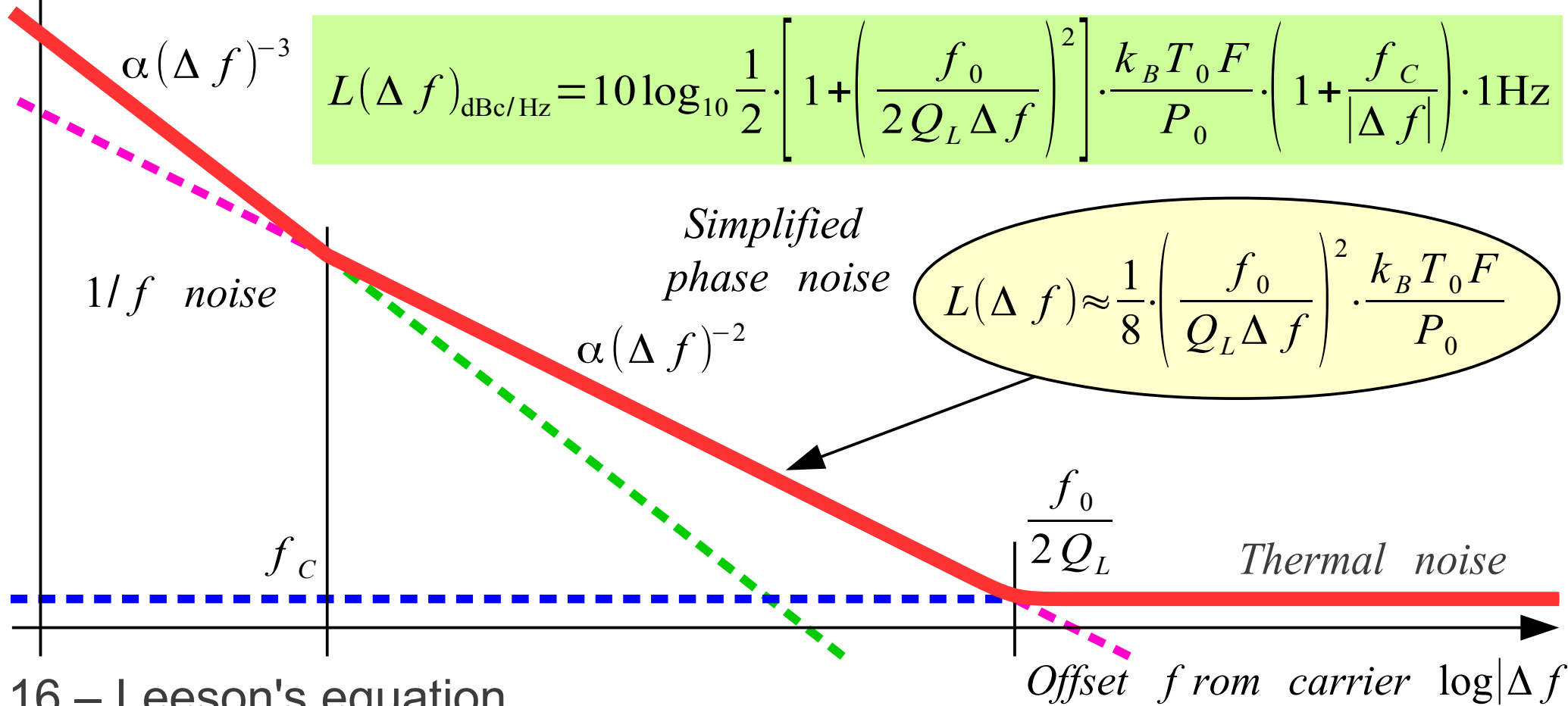
$$L(\Delta f) = \frac{1}{P_0} \cdot \frac{dP_\phi}{df} = \frac{1}{2} \cdot \left[ 1 + \left( \frac{f_0}{2Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left( 1 + \frac{f_c}{|\Delta f|} \right) \quad [\text{Hz}^{-1}]$$

Valid at  
 $L(\Delta f) \cdot \Delta f \ll 1$

Phase noise only

$P_0 \equiv$  carrier power

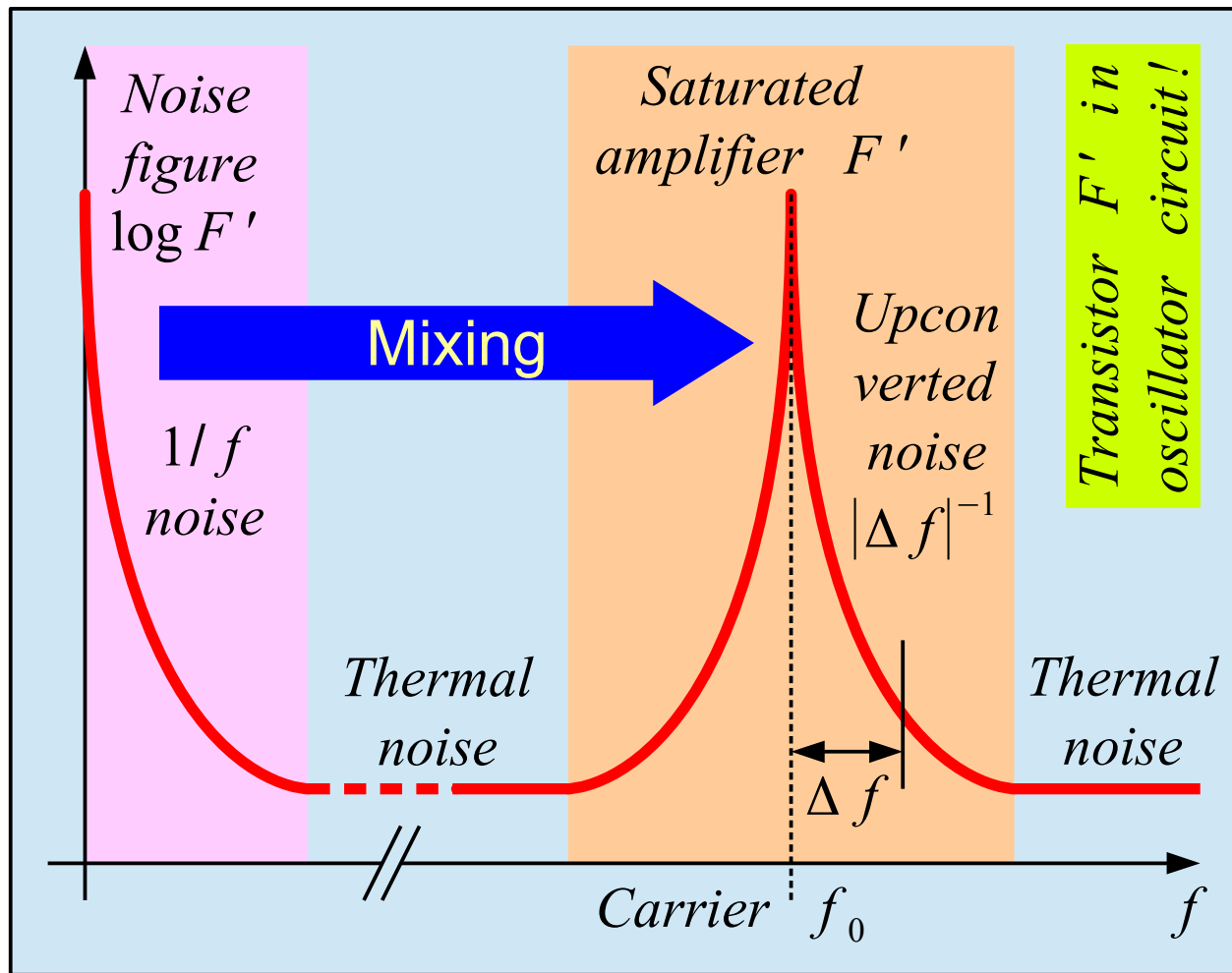
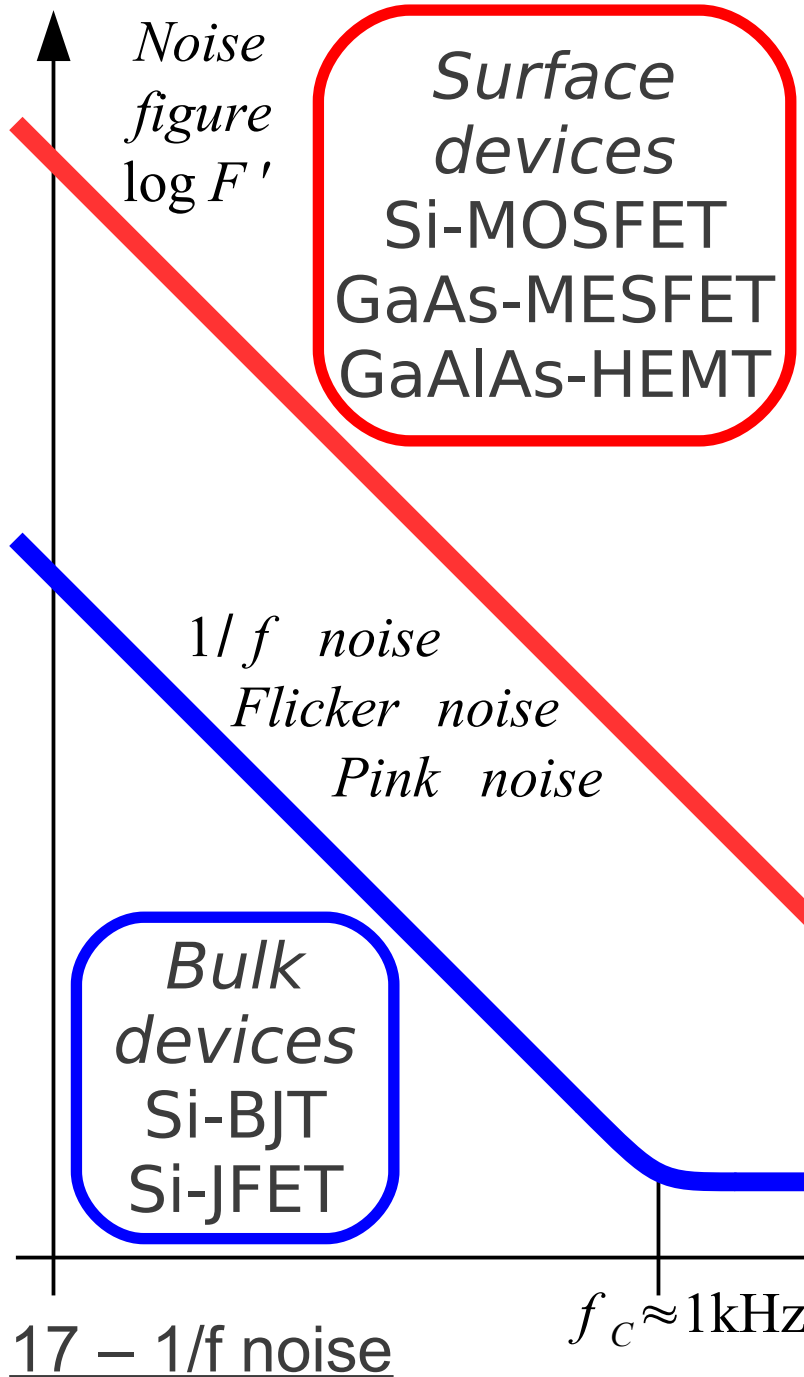
1/f noise



16 – Leeson's equation

*1/f noise usually does not have a clear explanation!*

**Surface devices**  
Si-MOSFET  
GaAs-MESFET  
GaAlAs-HEMT



$$F' = F \left( 1 + \frac{f_c}{f} \right) \equiv \text{increased LF noise!}$$

17 – 1/f noise

$f_c \approx 1\text{kHz}$

$f_c \approx 1\text{MHz}$

Frequency  $\log f$

The loaded resonator quality  $Q_L$  defines the oscillator phase noise!

$$L(\Delta f) = \frac{1}{2} \cdot \left[ 1 + \left( \frac{f_0}{2 Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left( 1 + \frac{f_c}{|\Delta f|} \right)$$

### Variable-frequency oscillators

$Q_L$

RC VCO

$\sim 1$

BWO tube

$\sim 1$

Varactor-tuned LC VCO

$10 \leftrightarrow 30$

YIG ( $\text{Y}_3\text{Fe}_5\text{O}_{12}$ ) oscillator

$300 \leftrightarrow 1000$

### Fixed-frequency oscillators

$Q_L$

RC multivibrator

$\sim 1$

LC resonator

$30 \leftrightarrow 100$

Cavity resonator

$1000 \leftrightarrow 3000$

Ceramic dielectric resonator

$1000 \leftrightarrow 3000$

AT-cut quartz crystal (fundamental mode)

$3000 \leftrightarrow 10000$

AT-cut quartz crystal (third/fifth overtone)

$10000 \leftrightarrow 30000$

Electro-optical delay line (\$)

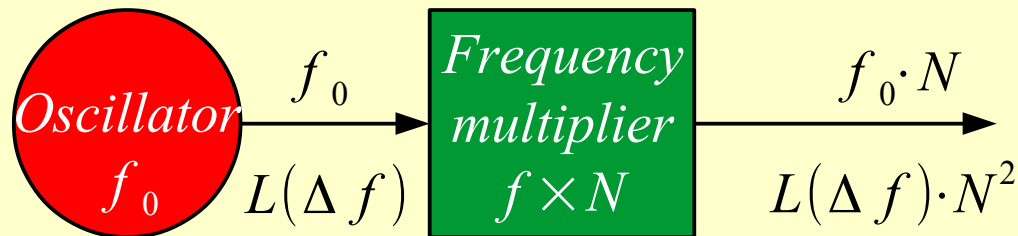
$\sim 10^6$  (noisy!)

Sapphire dielectric resonator (\$\$\$)

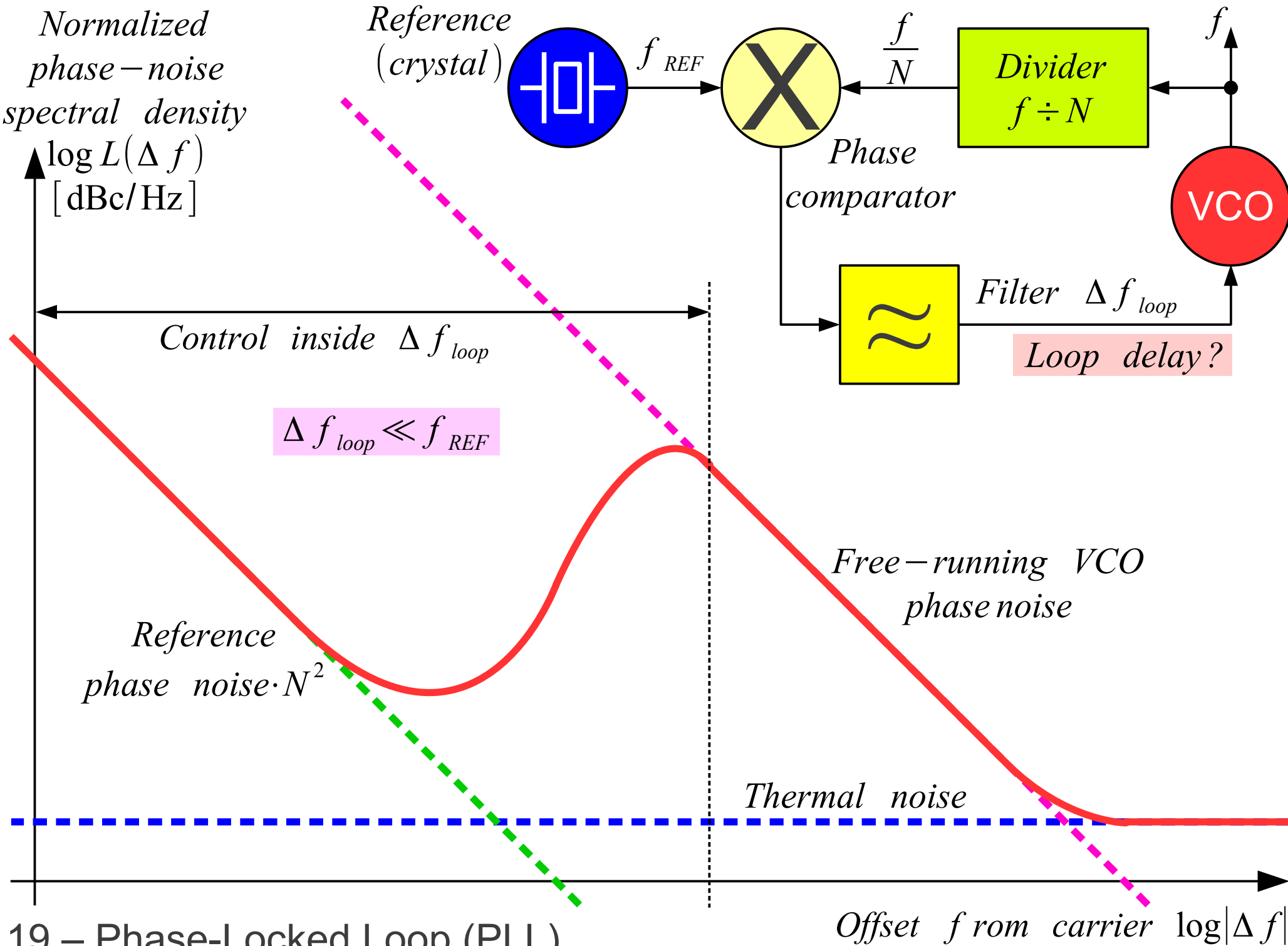
$\sim 3 \cdot 10^5$

Red HeNe LASER

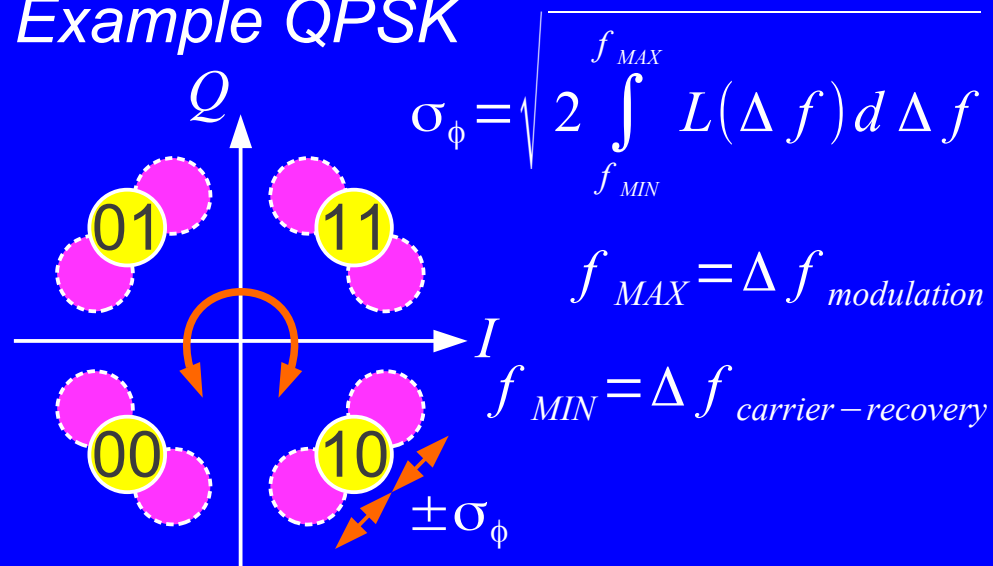
$\sim 10^8$



The phase noise multiplies with the square of the frequency multiplication factor!

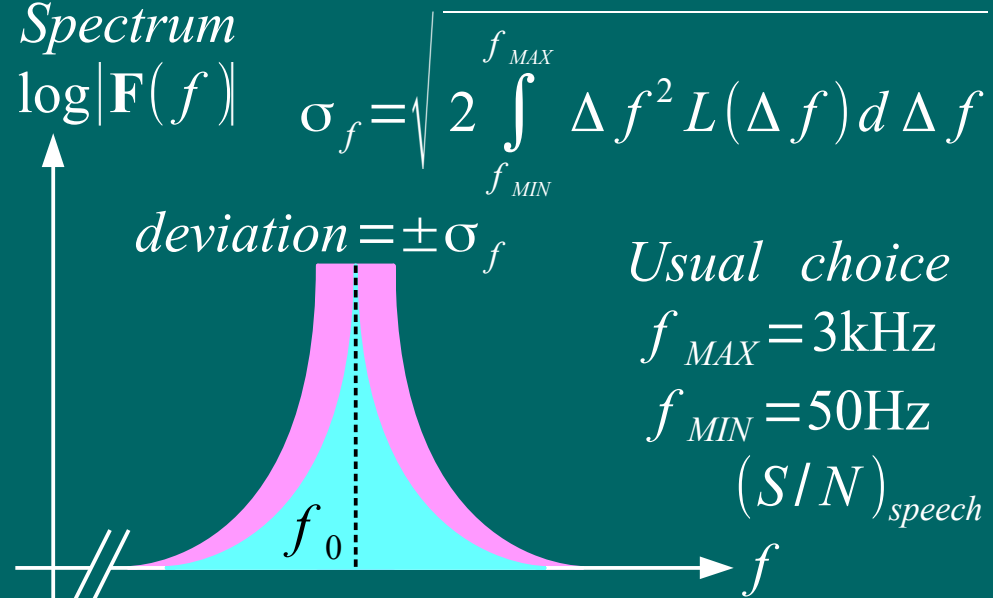


## Example QPSK



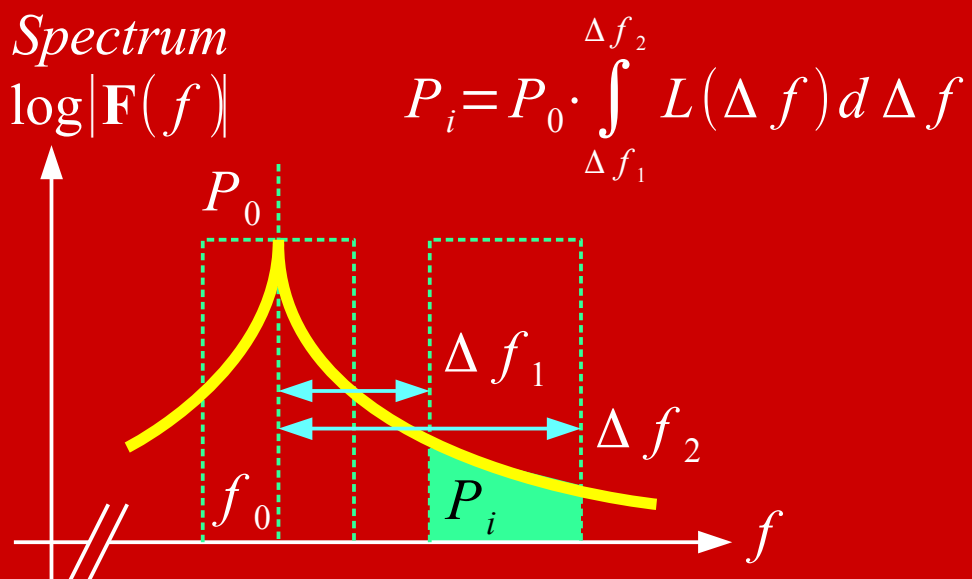
Modulation constellation rotation

## Spectrum



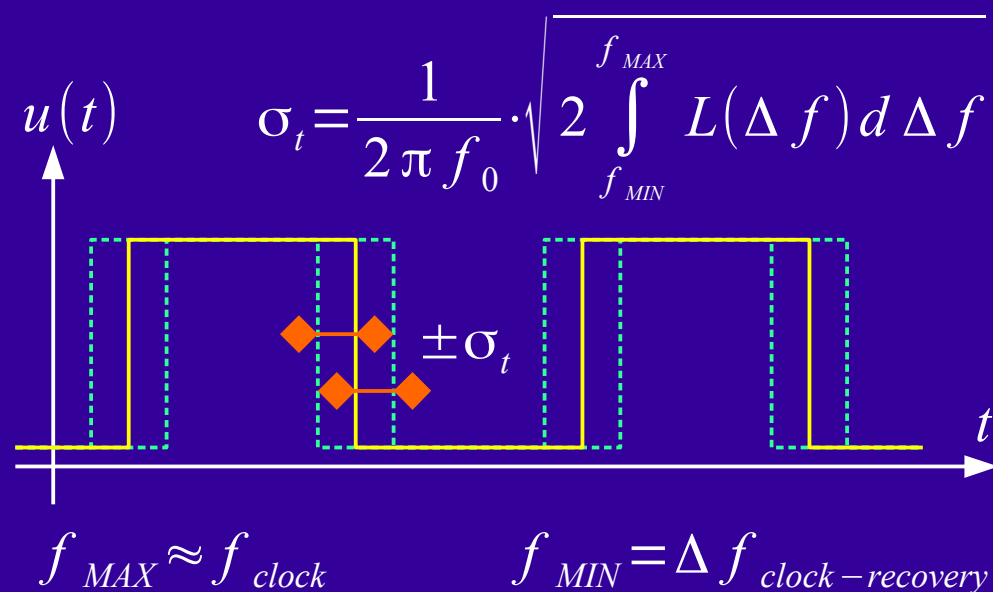
Residual FM

## Spectrum

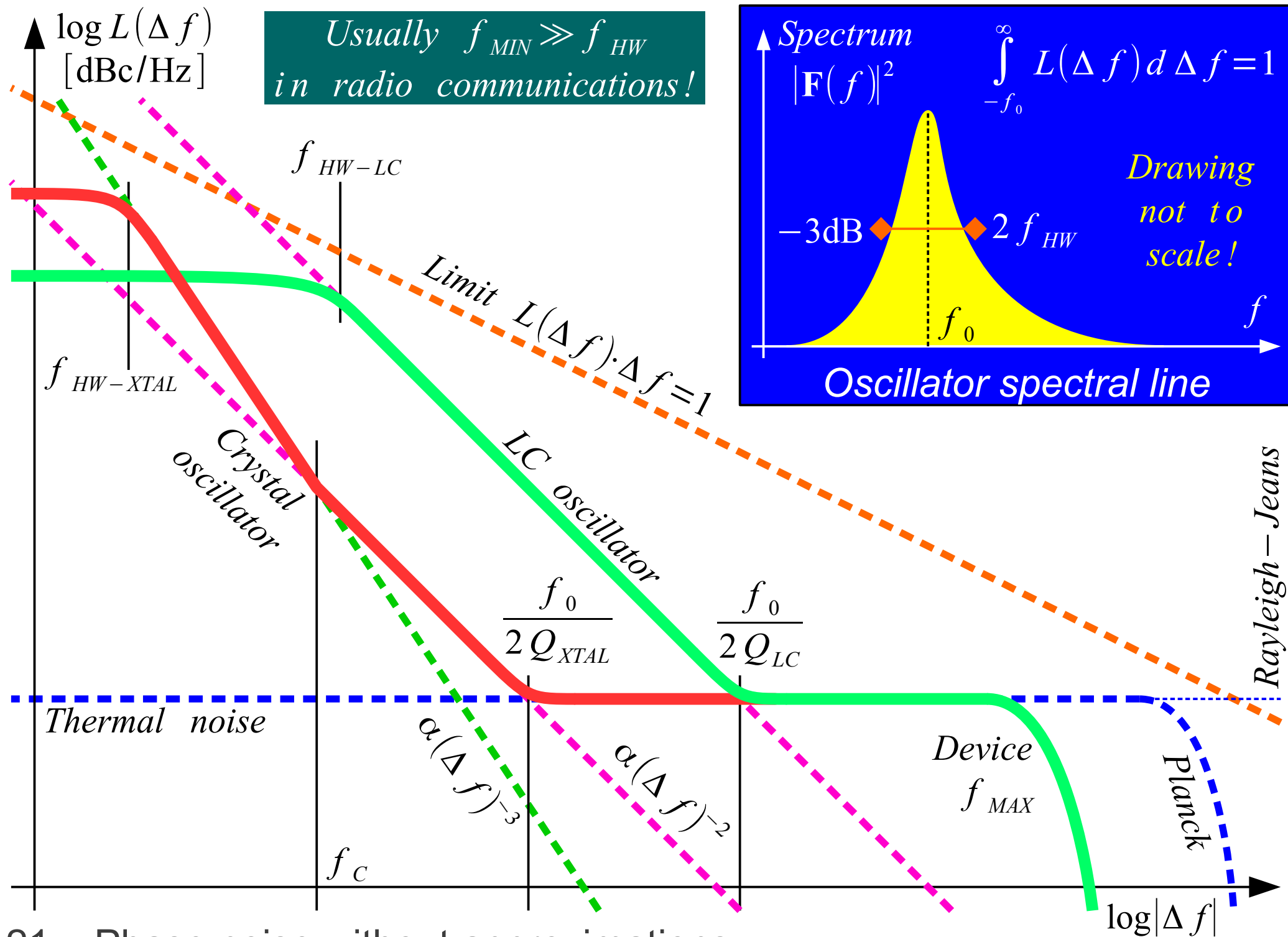


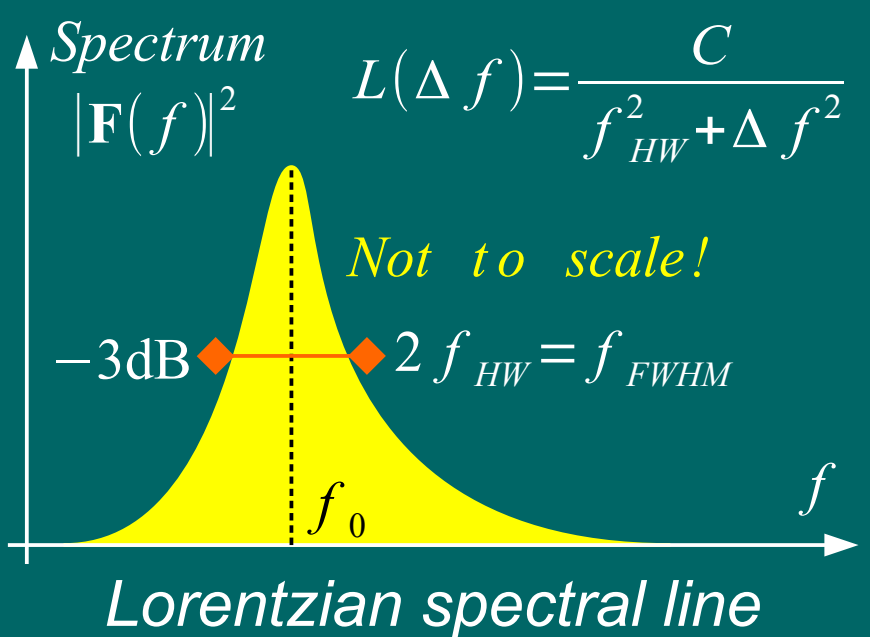
Adjacent-channel interference

## $u(t)$



Clock jitter





*Flat thermal noise can be neglected:  
device  $f_{MAX}$  or Planck law*

*LC-oscillator  $1/f$  noise can be neglected*

$$L(\Delta f) = \frac{1}{8} \cdot \left( \frac{f_0}{Q_L} \right)^2 \cdot \frac{1}{f_{HW}^2 + \Delta f^2} \cdot \frac{k_B T_0 F}{P_0}$$

*Lorentzian line in Leeson's equation*

$$\begin{aligned} \int_{-f_0}^{\infty} L(\Delta f) d\Delta f &= 1 \approx \int_{-\infty}^{\infty} L(\Delta f) d\Delta f = \frac{1}{8} \cdot \left( \frac{f_0}{Q_L} \right)^2 \cdot \frac{k_B T_0 F}{P_0} \int_{-\infty}^{\infty} \frac{1}{f_{HW}^2 + \Delta f^2} d\Delta f = \\ &= \frac{1}{8} \cdot \left( \frac{f_0}{Q_L} \right)^2 \cdot \frac{k_B T_0 F}{P_0} \cdot \left[ \frac{1}{f_{HW}} \cdot \arctan \frac{\Delta f}{f_{HW}} \right]_{\Delta f=-\infty}^{\Delta f=\infty} = \frac{k_B T_0 F}{8 P_0} \cdot \left( \frac{f_0}{Q_L} \right)^2 \cdot \frac{\pi}{f_{HW}} \end{aligned}$$

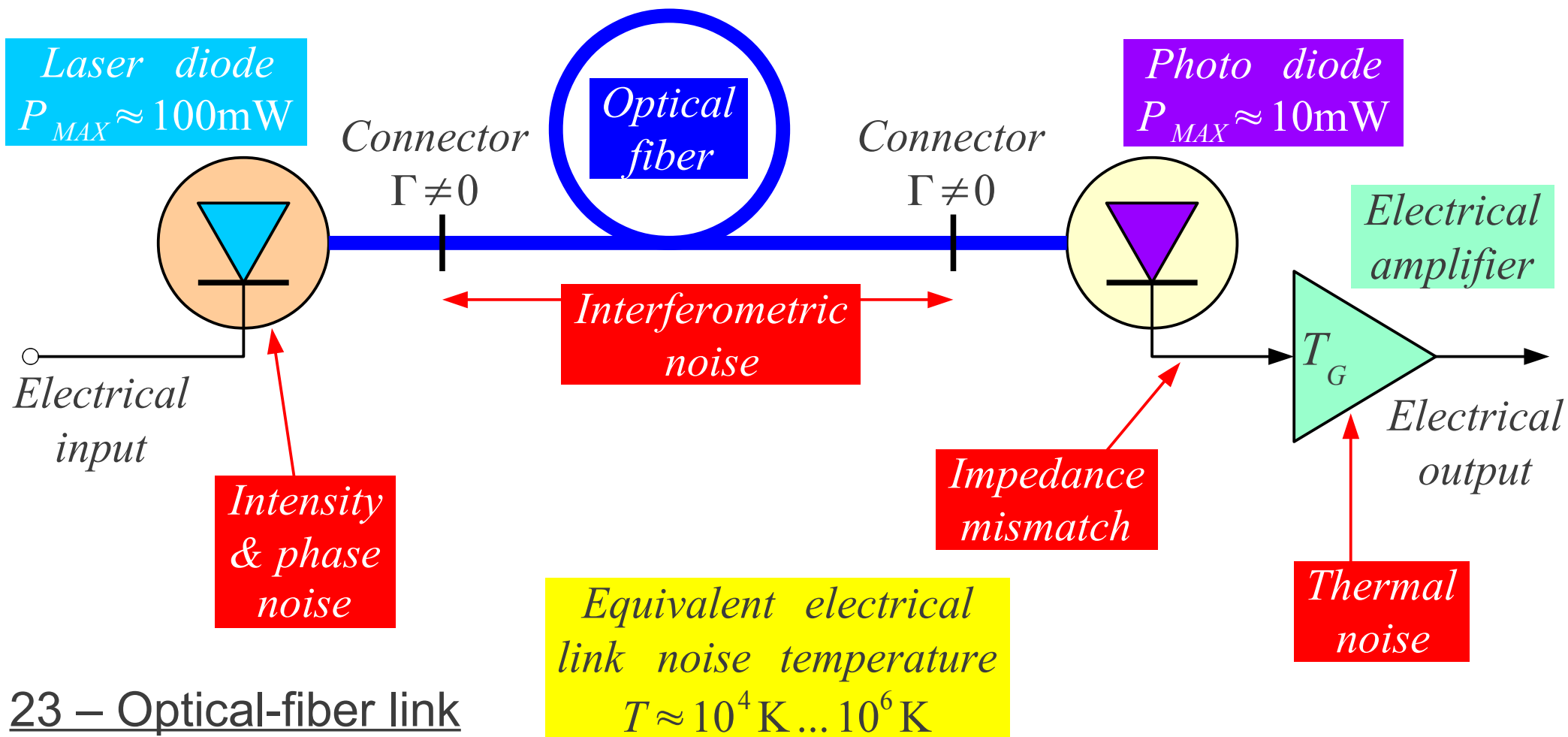
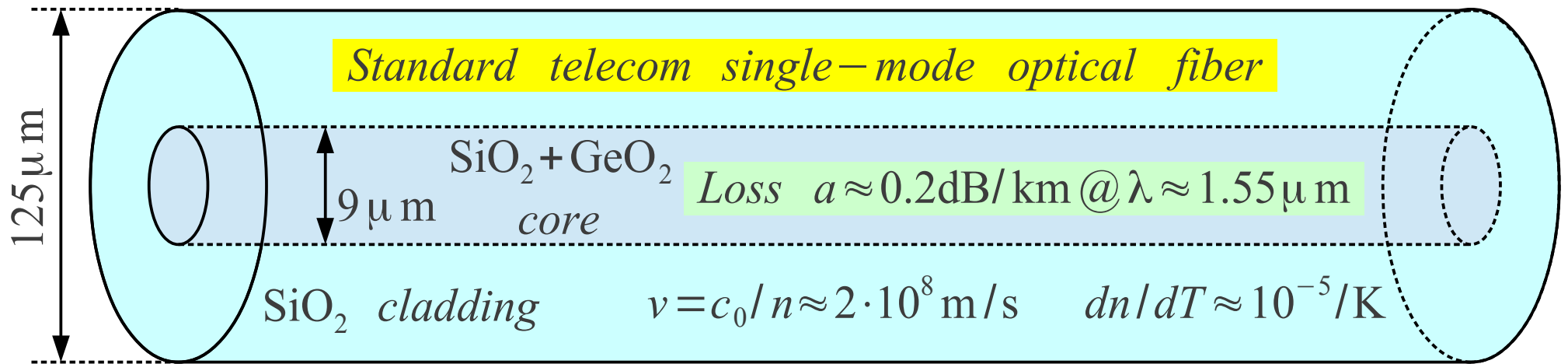
$$f_{HW} = \frac{\pi k_B T_0 F}{8 P_0} \cdot \left( \frac{f_0}{Q_L} \right)^2$$

*Example  $f_0 = 3\text{GHz}$   $Q_L = 10$   
 $P_0 = 0.1\text{mW}$   $F = 10\text{dB}$   
 $f_{HW} = 14\text{Hz}$   $f_{FWHM} = 28\text{Hz}$*

$$C = \frac{k_B T_0 F}{8 P_0} \cdot \left( \frac{f_0}{Q_L} \right)^2 = \frac{f_{HW}}{\pi}$$

$$L(\Delta f) = \frac{f_{HW}/\pi}{f_{HW}^2 + \Delta f^2}$$





*Electro-optical delay line*

$T_R \approx 10^5 \text{ K} \gg T_0$   
 $F$  can not be used

*Laser diode*

*Optical fiber*

*Photo diode*

$l \approx 50 \text{ km} \rightarrow t \approx 250 \mu \text{ s}$   
 $f \approx 10 \text{ GHz} \rightarrow Q_O = \pi f t \approx 7.9 \cdot 10^6$

$P_0 < 1 \text{ mW}$

*Electrical amplifier*

*Mode-select bandpass filter*  
 $Q_M \approx 10\% Q_O$

*Electrical amplifier*

*Advantage:*  
 Very high  
 $Q_L \approx Q_O + Q_M$

*Disdvantages:*  
 Very high  $T_R$   
 Low  $P_0$   
 Difficult  $Q_M$

*Simplified Leeson*  $L(\Delta f) \approx \frac{1}{8} \cdot \left( \frac{f_0}{Q_L \Delta f} \right)^2 \cdot \frac{k_B (T_G + T_R)}{P_0}$

*Electrical output*  
 10GHz

*Electro-optical delay line*

$$T_{RI} \approx 10^5 \text{ K} \gg T_0$$

*Laser diode*

*Optical fiber*

*Photo diode*

$$Q_O = \pi f t$$

*Q multiplier*

*cavity*  
 $Q_E$

$$Q_M = m Q_E$$

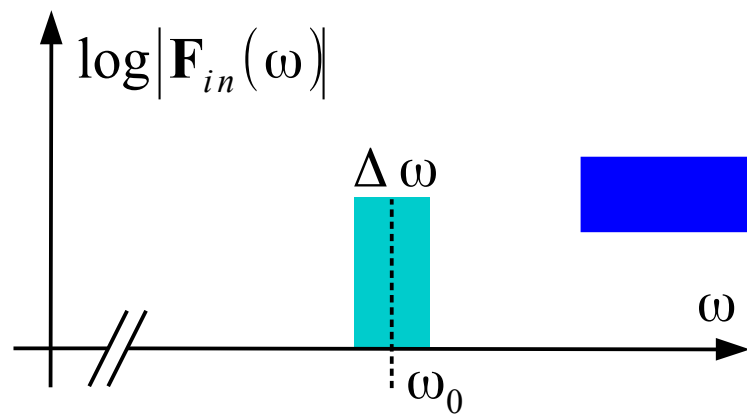
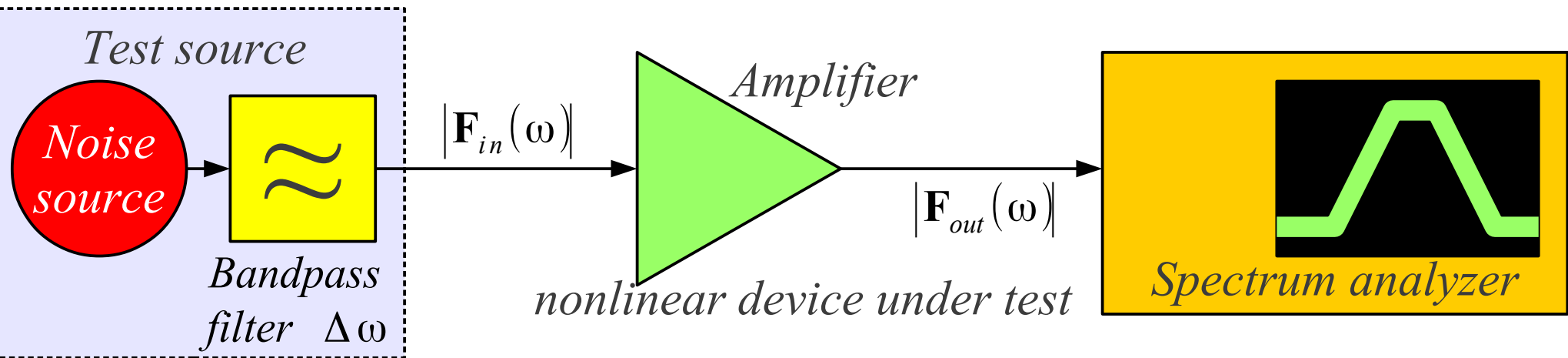
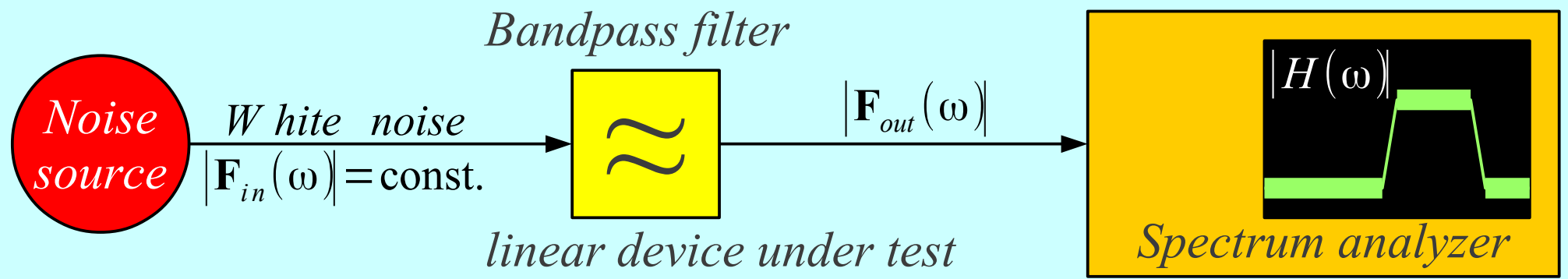
$$T_{R2} \approx m^2 (T_0 + T_G)$$

*Electrical output*  
10GHz

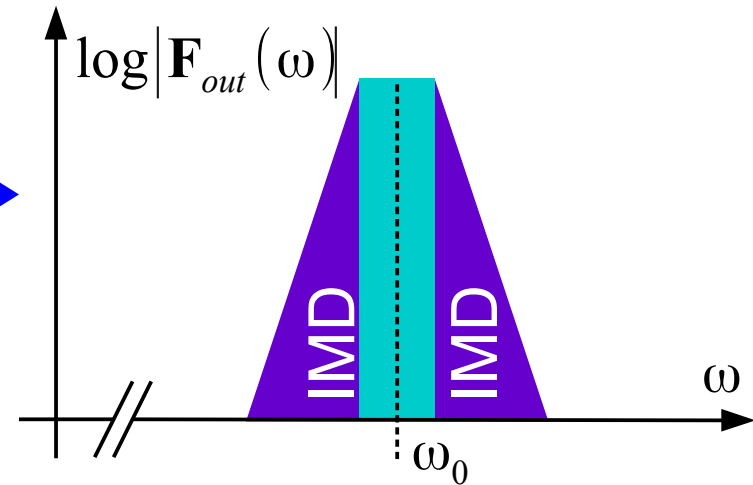
*Simplified Leeson*  $L(\Delta f) \approx \frac{1}{8} \cdot \left( \frac{f_0}{Q_L \Delta f} \right)^2 \cdot \frac{k_B (T_{RI} + T_{R2})}{P_0}$

$$Q_L \approx Q_O + Q_M$$

BOGATAJ, Luka,  
VIDMAR, Matjaž,  
BATAGELJ, Boštjan:  
*Opto-electronic oscillator  
with quality multiplier*,  
*IEEE transactions on  
microwave theory and  
Techniques*,  
ISSN 0018-9480.  
Feb. 2016, vol. 64,  
no. 2, pp. 663-668.



Nonlinearity



Natural sources of random signals:

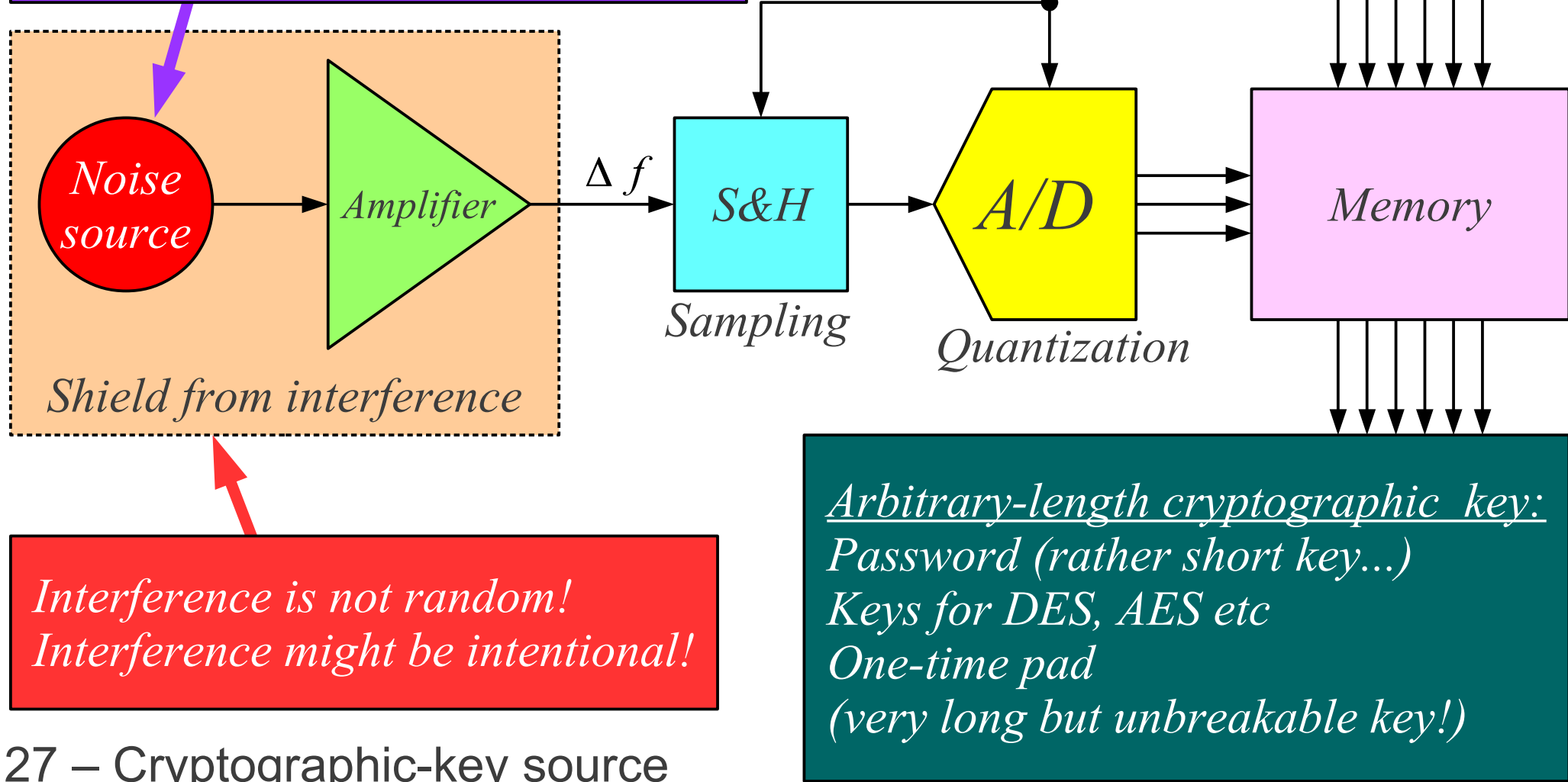
*Thermal noise*

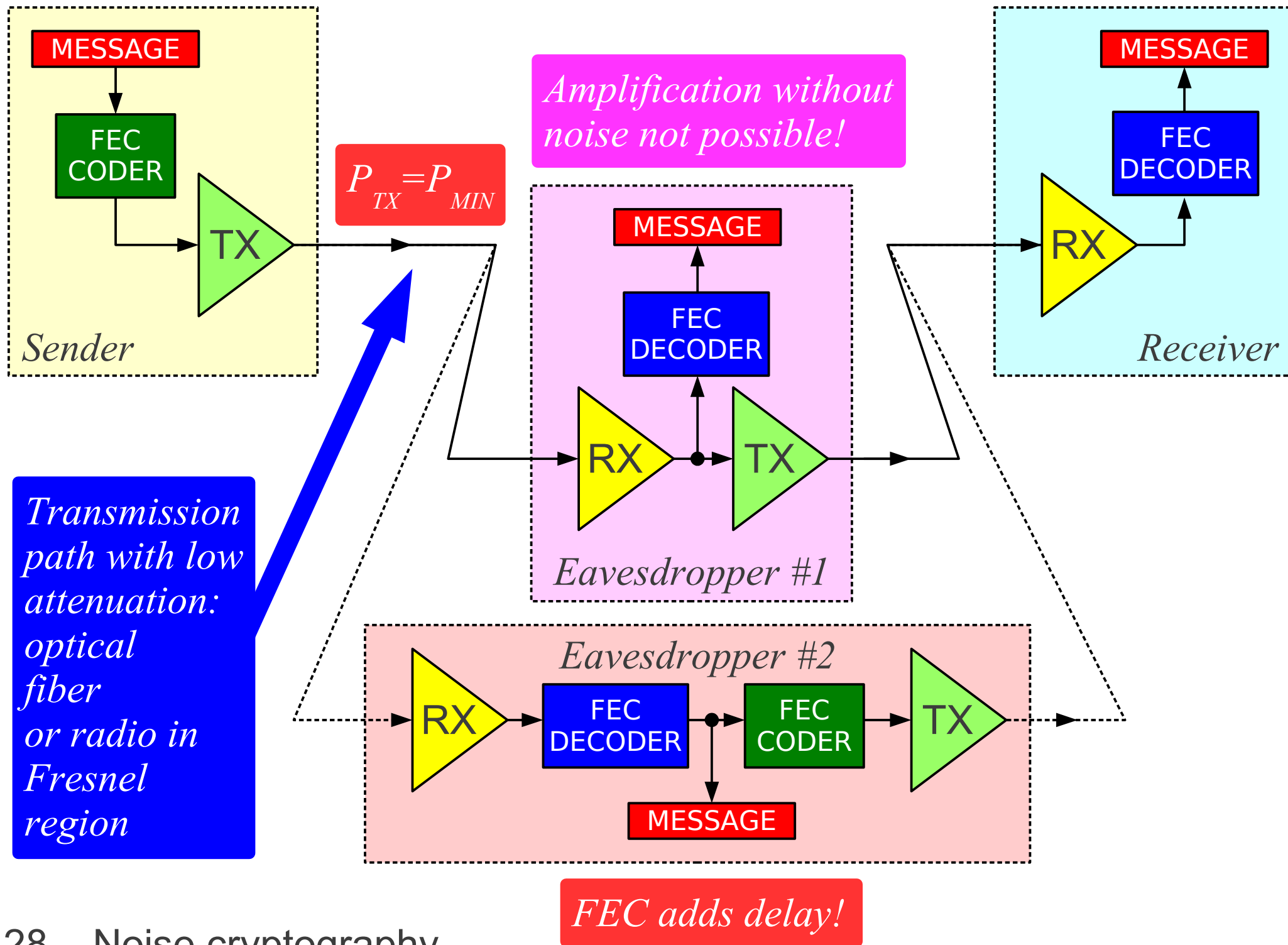
*Shot noise*

***Avalanche breakdown***

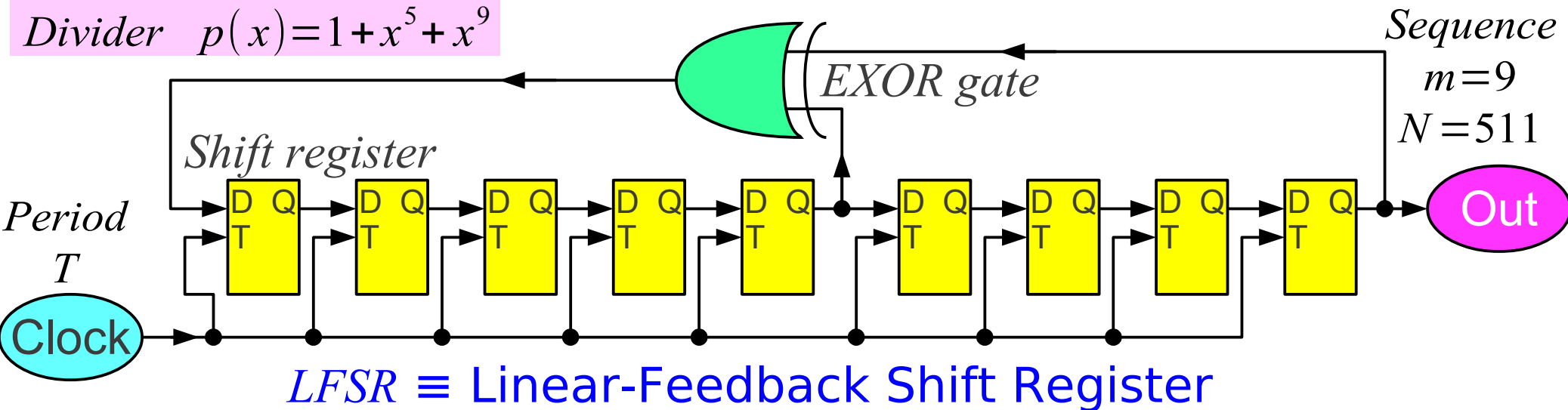
*Radioactive decay*

...





Divider  $p(x) = 1 + x^5 + x^9$



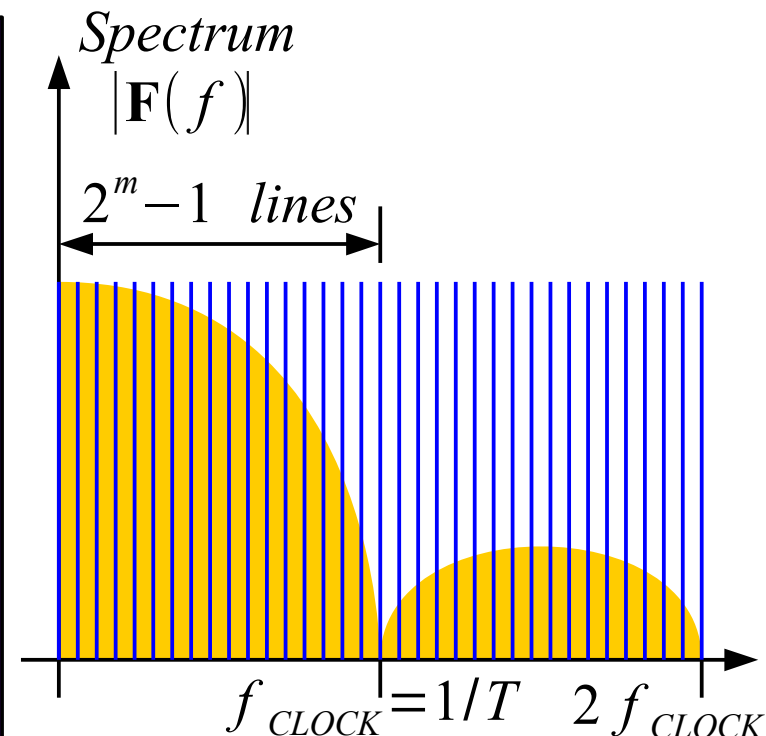
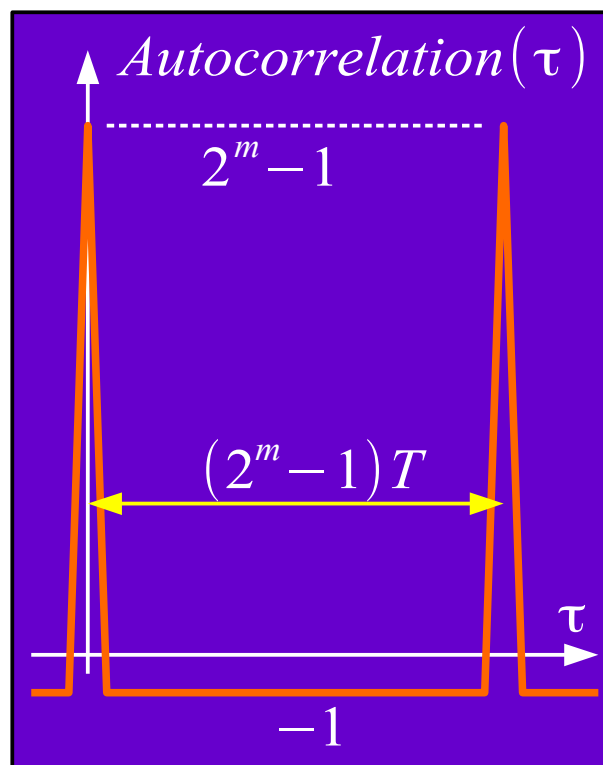
Primitive polynomial  $p(x) = 1 + x^l + x^m \rightarrow$  max sequence length  $N = 2^m - 1$

$2^{m-1}$  ones and  $2^{m-1}-1$  zeros  
arranged in groups of

- 1X m ones, m-1 zeros
- 1X m-2 ones and zeros
- 2X m-3 ones and zeros
- 4X m-4 ones and zeros

.....

- $2^{m-5}$  groups 111 and 000
- $2^{m-4}$  groups 11 and 00
- $2^{m-3}$  individual 1 and 0





Two-valued autocorrelation with a single very pronounced peak:

- synchronization headers for data frames
- spreading sequences in CDMA
- accurate time transfer in radio navigation (GPS, GLONASS)

Perfect frequency spectrum of uniformly-spaced lines and simple generation/checking:

- test sequences for all kinds of telecommunication links
- data scrambling (randomization) as part of line coding

Peak-to-average power ratio:

$$LFSR: \frac{P_{MAX}}{\langle P \rangle} \approx 1 \quad \text{Noise: } \frac{P_{MAX}}{\langle P \rangle} \rightarrow \infty$$

LFSR pseudo-random sequences are of NO cryptographic value:  
Berlekamp-Massey algorithm 1969

*LFSR sequences are the result of pure mathematics that does not appear anywhere else in nature!*

*How to present ourselves to the inhabitants of a neighbor galaxy?  
How to find out that they look for us?*